

1 **Manuscript Number: AES6163**

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3 **Title: Crustal thickness estimates beneath four seismic stations in Ethiopia inferred from p-**
4 **wave receiver function studies**

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6 Keywords: Receiver Functions; Crustal structure; Moho; Partial melt; Anomalous Moho;
7 Velocity model

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14 **Abstract**

15
16 Moho depths beneath four stations of the Ethiopian Seismic Station Network (ESSN) are estimated
17 from p-wave receiver functions (RF). We used high quality seismic data recorded at ANKE
18 (Ankober), DILA (Dilla), HARA (Harar) and SEME (Semera) stations for earthquakes located at
19 epicentral distances ranging from 30 to 100 degrees with magnitude $m_b \geq 5.5$. We applied a
20 frequency domain deconvolution technique to remove source and propagation path effects from
21 the earthquakes waveforms to make the RFs dependent only on the structure beneath the seismic
22 stations. A linearized-iterative inversion is applied on the generated radial component of the
23 receiver functions. The minimum number of teleseismic earthquakes used is 14 for HARA while
24 the maximum is 39 for SEME station. A linearized-iterative inversion is applied on the generated
25 radial component of the receiver functions to obtain p-wave velocity models beneath the stations.
26 We achieved a reasonably good fit between the observed and synthetic RFs, which demonstrated
27 the high quality of the inversion process. From the obtained models we estimated Moho depths of
28 26 ± 2 km for SEME, 36 ± 2 km for DILA, 38 ± 2 km for HARA and 42 ± 1.7 km for ANKE. We
29 have achieved a reasonably good fit between the observed and synthetic RFs which demonstrates
30 the quality of the inversion. The lowest Moho depth is observed at Semera station which implies
31 a thinned crust while the highest crustal thickness is observed at Ankober, which lies along the
32 North western plateau margin. Our results agree with previous observations which strengthen the
33 hypothesis that Moho depths estimated for stations that lie within the rift and rift margins are lower
34 than those located in the plateaus. Our RFs inversions show a low velocity gradient at about 16km
35 depth at Semera station, interpreted as evidence for lower crustal storage of partial melt.

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43 1. Introduction

44 Most continental rifts are thought to extend by some component of mechanical extension in which
45 faulting and stretching of the tectonic plate defines the primary architecture of the rift (Weissel
46 and Karner, 1989). Ultimately, however, the locus of strain must shift towards a narrow zone that
47 becomes the newly formed seafloor spreading center. It is here that magma formed from
48 decompression melting of the mantle intrudes the plate and creates a new ocean floor (Whitehead
49 et al., 1985). Despite the importance of continental breakup in plate tectonic theory, it remains
50 unclear how and when the transition from mechanical to magmatic extension of the plate occur
51 (Hayward and Ebinger, 1996; Hopper et al., 2004). For example, it is still unclear what proportion
52 of rift extension is taken up by the intrusion and where the magma is stored in the crust beneath
53 Afar.

54 In Afar, mechanical extension is thought to have initiated around ~30 Ma on large offset border
55 faults that still define the edge of the rift (Wolfenden et al., 2004). These observations have led to
56 the interpretation that mechanical extension by ductile stretching occurred beneath the fault bound
57 rift (Bastow and Keir, 2011). The geological record suggests the locus of extension migrated
58 progressively in-rift, with exposed dikes and lava flows suggesting that the progressive extension
59 included a magmatic component. The current locus of dike intrusion (Keranen et al., 2004), recent
60 volcanism, young faults (Corti et al., 2009), earthquakes, coupled with the geodetically constrained
61 ground motions (Wright et al., 2006; Dumont et al., 2017) suggest extension is now mainly focused
62 in ~20km-wide and ~50—100 km long volcanic segments that define the axis of the rift (Ebinger
63 and Casey, 2001).

64 This study conducted, P to S wave conversion receiver function analysis at four newly occupied,
65 permanent, seismic stations. Therefore, this work is aimed at estimating the Moho depth beneath
66 Semera (Central Afar), Ankober (Northwestern plateau), Harar (Southeastern Plateau) and Dila
67 (Southern Main Ethiopian Rift) (Figure 1) using teleseismic data recorded from September 2014
68 to June 2015 at Ethiopian permanent broadband seismic stations. The aims of this study is to
69 constrain crustal thickness and internal crustal structure in the rift and adjoining plateau using
70 passive source teleseismic receiver functions. In particular, we aim to better understand how the
71 lower crust has been modified by magma since this has implications for the thickness, strength of
72 extending crust, as well as for quantifying the mode of extension. Though similar receiver function

73 studies have been done in the rift and adjoining plateaus, this work focuses on Moho depth
74 estimation beneath three new sites with few prior constraints. We also conduct a re-appraisal of
75 one station.

76 **1.1 Previous constraints on crustal structure in Ethiopia**

77 The crustal structure beneath parts of Ethiopia has been constrained using both passive and
78 controlled source seismic techniques, such as magnetotelluric imaging, and inversion of gravity
79 data (Didana et al., 2014; Lewi et al., 2015). Seismic imaging is in broad agreement with other
80 geophysical methods such as magnetotelluric and gravity methods and shows that crustal thickness
81 beneath the Northwestern plateau is 35—45km, with the thicker crustal regions along the Southern
82 Red Sea margins and central Northwestern plateau, including up to 10km of high velocity lower
83 crust (P-wave velocity of 7.4—7.7 km/s) (Dugda et al., 2005; Maguire et al., 2006; Stuart et al.,
84 2006; Dugda and Nyblade, 2006; Ebinger et al., 2011). This high velocity lower crustal layer has
85 been interpreted as “underplating” of mafic rock or high density lower crustal sill intrusion during
86 Oligocene to Recent magmatism (Stuart et al., 2006; Cornwell et al., 2010; Maguire et al., 2006;
87 Hammond et al., 2011; Hammond, 2014). This contrast with the South Eastern plateau where the
88 receiver functions show the crustal thickness is consistently $\sim 35 \pm 1$ km thick (Hammond et al.
89 2011). From seismic refraction results V_p through the upper mantle beneath the Northwestern
90 plateau is about 8.0 km/s (Makris and Ginzburg, 1987). We summarized previous findings of
91 crustal thickness in Ethiopia estimated from receiver function studies in Table 1.

92 The crust is generally thinner and has higher V_p/V_s ratio in the Main Ethiopian Rift (MER) and
93 Afar than the Ethiopian and Southeastern Plateau. Thickness of the crust beneath the rift
94 reduces northward from ~ 36 km in the MER, to ~ 25 km in most of Afar, to ~ 15 km in northern Afar
95 (Hammond et al., 2011). Recent multidisciplinary studies have helped to better image the crust in
96 the region of the Tendaho graben (station SEME). Magnetotelluric surveys showed the presence of
97 a highly electrically conductive body ranging 15 to 28 km depth and ~ 13 km width beneath the rift
98 axis, which is in turn interpreted as evidence of considerable melt/partial melt fraction in the crust
99 (Kind et al., 1996; Desissa et al., 2013; Didana et al., 2014; Hammond, 2014; Lewi et al., 2015).
100 Modeling of dense micro-gravity measurements supports the presence of low density material
101 (basaltic melt in a magma reservoir) whilst MT profile results show that high electrical
102 conductivity material occupies space of ~ 13 km extent at a depth of about 15—28 km at the rift

103 axis (Johnson, 2012. On a regional scale, the crustal thickness constrained by modeling gravity
104 data is broadly consistent with that constrained using seismology (Tiberi et al., 2005; Mammo,
105 2013) (Table 2).

106 **2. Methods**

107 The receiver function method (RF) is a well established time series technique, which is widely
108 used to study the structure of the Earth's crust and upper mantle (Langston, 1979; Ammon et al.,
109 1990; Ammon, 1991). It is the transfer function between the direct P and converted S waves with
110 all P and Ps multiples and reverberations as shown in figure (A) and (B) in Figure 3 (Ammon,
111 1991; Jesse and Douglas, 2004). In order to address thickness of the crust and its layered structures
112 before and during progressive stages of rift evolution, we applied this method using high quality
113 seismic data obtained from 4 permanent Ethiopian stations. Using Langston's [1979] source
114 equalization procedure to remove the effects of near-source and near receiver structure are a major
115 procedure for receiver function analysis. The procedure includes phase information, complex
116 frequency-domain ratio and inverse transforming back into the time domain. A deconvolution
117 approach employed in this work used a water-level stabilization method. Then, a low-pass
118 Gaussian filter removes high-frequency noises that are not filtered by the water-level. Following
119 a low pass Gaussian filter (reducing spectral artifacts to acquire better P-onset arrival time),
120 rotation of the coordinates from ZNE (Z = vertical, N = north, E = east) to ZRT (Z = vertical, R =
121 radial, T = transversal) was done to isolate the converted S phase of the direct P wave using Seismic
122 Analysis Code (SAC) software (Goldstein, 1999). Radial and transverse receiver functions were
123 calculated using the source equalization approach to derive the structural response beneath the
124 recording station (Langston, 1979). This approach was done using the frequency domain water-
125 level deconvolution method (Clayton and Wiggins, 1976; Midzi and Ottemöller, 2001).

126 The powerful water-level technique was done by deconvolving the vertical from radial and
127 transverse components in the frequency domain to remove the signature of source, travel path and
128 instrumental response effects (Langston, 1979; Ammon et al., 1990; Ammon, 1991; Dugda et al.,
129 2005; Tuluka, 2010). Then, by multiplying the result obtained from water-level deconvolution
130 with low-pass Gaussian filter high-frequency signals were excluded. Finally, receiver functions
131 were stacked before being inverted to obtain the plane layered crustal velocity structure. Various
132 starting forward velocity models ranging from 10 to 30 layers were synthetically generated for

133 each receiver function. A smoothness factor in the range of 0.0 and 2.0 and Gaussian width factor
134 of about 1.0 (Ammon et al., 1990; Ammon, 1991; Randall, 1989; Zandt and Ammon, 1995;
135 Tuluka, 2010) were used to obtain a better root mean square fit between synthetic and observed
136 receiver functions. Receiver functions were calculated using the singular-value decomposition
137 method with a water level parameter c (in the range of 1.0 to 0.0001) to avoid division instabilities
138 during the deconvolution process (Clayton and Wiggins, 1976; Randall, 1989; Ammon et al., 1990;
139 Ammon, 1991; Cassidy, 1992; Mangino et al., 1993; Midzi and Ottem, 2001; Tuluka, 2010) for a
140 better resolution of arrival times. Linearized inversion methods were used to model RF using
141 several different starting models chosen by referring to a priori information from the previous work
142 conducted near the study areas to make an initial model as close as the true Earth velocity structure
143 as possible (Ammon et al., 1990). The receiver functions (RFs) were produced using the programs
144 written by Ammon, and others (e.g., Ammon, 1991).

145 By calculating the time difference between the direct P (tp) and the converted Ps (tps) phase of the
146 receiver function plot, the crustal thickness can then be estimated by multiplying average velocity
147 above the Moho with the time difference. This estimation provides a good “point” measurement
148 at the station because of the steep incidence angle of the teleseismic P wave (Zhu and Kanamori,
149 2000).

150 To obtain the fittest synthetic receiver function with its corresponding inverted velocity model, we
151 applied an iterative inversion technique that employed an initial model which presumed downward
152 increments of velocity. We then used a gradient-based inversion algorithm to evaluate at which
153 iteration the synthetic and observed waveforms are matched. Once we obtained the minimum
154 number of iterations by which calculated synthetic value converges to the observed model, then
155 that is the end of selecting a station’s smoothest model that matches the observations (Ammon,
156 1991). The iteration halts when the calculated receiver function begins to repeat itself without
157 having significant modification during fitting process with the observed receiver function (Ligorria
158 and Ammon, 1999). Fittest selection criteria used in this study is by visual coherency examinations
159 of corresponding Ps conversions of synthetic RF with stacked observed RF that is obtained by
160 careful elimination of outliers and noisy records before the start of inversion processes (Tkalčić et
161 al., 2011).

162 This work chose frequency domain over time domain method because the latter method is less
163 efficient than simpler methods such as water level deconvolution for moderate earthquake source
164 though it is efficient to estimate from large earthquake's source (Legorria and Ammon, 1999). The
165 second reason we chose the frequency domain over time domain is due to the easiness of applying
166 Water-level and Gaussian transform to limit frequency band by excluding high frequency signals
167 (artifacts that are not obviously present in the original recording) (Langston, 1979; Clayton and
168 Wiggins 1976) in addition to its simplicity and time efficiency (Bona, 1998). But both methods
169 are equally effective for estimating receiver functions from high magnitude signals (Legorria and
170 Ammon, 1999).

171 Frequency domain deconvolution, a spectral division technique, has an advantage to resist leakage
172 of the low amplitude portions from P-wave receiver functions (Jeffrey and Vadim, 2000). But, it
173 has certain shortfalls; two of these disadvantages are exposed to instability caused by the very low
174 denominator (exposed to spectral hole) and seismic noise (Bone, 1998). To circumvent the side
175 effect of frequency domain deconvolution, a modified spectral division is used by applying water
176 level parameters that avoids numerical instable zeroes of the denominator (Clayton and Wiggins,
177 1976; Ammon, 1991; Jeffrey and Vadim, 2000). When the data are wide band with good signal to
178 noise levels most deconvolution methods such as frequency domain and time domain approaches
179 work well, and the advantage of one method over the other is insignificant (Ligorria and Ammon,
180 1999) and no deconvolution approach outshines all others in all occasions and complications. In
181 this work we used both receiver function and inverted velocity versus depth plot for a conclusion
182 whether the point is Moho or not as Ps phase may be hidden in the noise and/or display complex
183 masks from near-surface reverberations and P-wave multiple reflections and may not fully judged
184 by the velocity versus plot alone (Yuan et al., 2006).

185 When unconsolidated sediments are unearthed below a station, the strong reverberations generated
186 at its base mask the Ps conversions at deep structure such as the Moho. To minimize such hidden
187 phase shift caused by masked structure we included velocity of sedimentary structure in the initial
188 model. However, application of this technique alone didn't fully remove the occurrences of misfit.

189 In this work we used both receiver functions and inverted velocity versus depth plots for a
190 conclusion whether the point is Moho or not as Ps phase may be hidden in the noise and/or display

191 complex masks from near-surface reverberations and P-wave multiple reflections and may not
192 fully judged by the velocity versus plot alone (Yuan et al., 2006).

193 **3 Results and Discussion**

194 Using the steps and processes described in Section 2 observed waveform receiver functions were
195 estimated for the four stations (Figure 4A). Corresponding four calculated synthetic receiver
196 functions were chosen from the smoothest model that matches (Figure 4B) the observations by the
197 use of visual match. From calculated smoothest models various crustal structures were obtained
198 (Table 3).

199 The number of iterations in any inversion depended on the complexity of the structure and the
200 closeness of the initial model to a structure that matches the observations (Ammon et al., 1990 and
201 Ammon, 1991). The P wave velocity models obtained from iterative inversion technique (Figure
202 5) displayed the velocity contrast of ANKE, DILA, HARA and SEME stations at various depths,
203 which describe different crustal structural units. By using good quality teleseismic signals a
204 reliable crustal thickness and P wave velocity (V_p) contrast were estimated. However, the
205 numerical value of V_p in SEME is not in the range of known V_p passing through Moho, the
206 anomalous Moho depth beneath SEME station is about 26km with a V_p value of 5 km/s. This
207 anomalous low velocity zone was obtained in the range of 15 km to 26 km deep. The station DILA
208 located at the southern main Ethiopian rift was obtained to be the second shallowest Moho in our
209 study with a depth of 36 km and with V_p 7.53 km/s, however, the value of V_p at different depths
210 were found to be less heterogeneous (i.e. ranging from 6.0—7.0 km/s). There is a relatively high
211 velocity contrast at the depth of 36km from the top beneath this station. This contrast might be due
212 to the existence of Moho beneath DILA, a rift margin of the southeastern plateau (Mahatsente et
213 al., 1999).

214 For the station HARA, which is in the southeastern plateau, there is a relatively high change in V_p
215 from 7.8 km/s to 8.0 km/s at a depth of $38\text{km} \pm 2\text{km}$ from the surface. The velocity contrast at this
216 depth could be the location of the Moho interface. The Moho depth obtained in HARA is similar
217 to the results obtained by Mahatsente et al. [1999] and Hammond et al. [2011] beneath the
218 Southeastern plateau, which range from 36 to 38km. Significant P-wave velocity contrast of about
219 7.8km/s is observed at the depth of $42\text{km} \pm 1.7\text{km}$ beneath ANKE which is thickest crust with

220 normal V_p beneath the station and this result has a good agreement with the previous studies
221 (Mahatsente et al., 1999; Dugda et al., 2005; Stuart et al., 2006; Hammond et al., 2011; Mammo,
222 2013). From inversion results of velocity versus depth a high velocity zone of about 7.1km/s is
223 obtained in the range of ~ 7 —10 km. The average value of V_p increases towards the Northwestern
224 plateau and the reason for this higher than average upper crustal seismic velocity could be a mafic
225 dike intrusion. In Figure 4B the observed and calculated receiver function obtained at HARA and
226 SEME station look like they don't fit. This might be due to shallow volcanic and sedimentary
227 layers which manifest high velocity contrast at its base since P wave arrivals overwhelmed by the
228 converted phase and reverberations (Ligorria and Ammon, 1999). The locality beneath the station
229 SEME could be layered soft and thick unconsolidated sediment that mixes the direct P and
230 converted P_s phase (Chen and Niu, 2016).

231 Overall, the thinnest crust is observed in SEME with the existence of anomalous Moho, while the
232 thickest crust is in the Northwestern plateau beneath ANKE with the observed normal Moho
233 (Table 3). Similar patterns of Moho depths were found near these two localities in previous studies
234 (Dugda and Nyblade, 2006; Stuart et al., 2006; Cornwell et al., 2010). From the velocity models
235 obtained from the inversion process, the top most crustal layer of about 4 km at SEME has a very
236 low velocity, V_p . Beneath this low velocity there is a high velocity (~ 6.6 km/s—7.6 km/s) thick
237 crustal layer of about 8km thick. At 16km the velocity decreases steeply to a low velocity layer of
238 about 10km thick. The anomalous Moho depth beneath SEME station is about $26\text{km} \pm 2\text{km}$, while
239 the Moho depth beneath ANKE is at the depth of $42\text{km} \pm 1.7\text{km}$.

240 The crustal structure observed beneath SEME is distinctly different from that observed under the
241 other three stations. The upper crustal layers in the top 4km have very low P wave velocities
242 ranging from 2.4 to 3.2 km/s. This might be interpreted as a cover sediment (maybe evaporates)
243 or volcanic layer. Top most layers beneath SEME are likely to be volcanic rocks intercalated with
244 thin sediments. This is similar to the results obtained by Lewi et al. [2015] for a site with surface
245 sediments of a thickness not more than 1.2 km is decreasing to 0.75 km northwards.

246 Beneath this station a high velocity layer (~ 7 —7.6 km/s) between 6 km to 14km is consistent with
247 the presence of cooled magmatic intrusion. Below this high velocity zone, a very low V_p value
248 was obtained (Figure 6) in the depth range of 16 km—26 km, which is unusual compared to earlier
249 receiver function studies near the area (Dugda et al., 2005; Dugda and Nyblade, 2006; Stuart et

250 al., 2006; Hammond et al., 2011; Hammond, 2014). Forward modeling shows that the observed
251 very low velocity is not an artifact of the inversion process, but can potentially be explained by
252 deep seated fluid accumulation. This conclusion is justified by the fact SEME station is located at
253 the southeastern terminus of the active Manda Hararo Rift system which is near the Tendaho basin
254 where deep seated magma and hydrothermal reservoirs are potentially possible at the specified
255 depth (Hernández-Antonio et al., 2015). A similar inference is given by Field et al. [2012] as there
256 is peralkaline magmas beneath Dabbahu Volcano, Afar. Using secondary ion mass spectrometry
257 (SIMS) analysis of volatile contents in melt inclusions (Kind et al., 1996) and olivine from
258 pantellerite obsidians which represent the youngest eruptive phase (<8 ka) that includes H₂O
259 contents ≤ 5.8 wt. % and CO₂ contents below 500 ppm (Field et al., 2012). Petrological methods
260 for constraining magma storage depths include the identification of dissolved volatiles (principally
261 H₂O and CO₂) in phenocryst-hosted melt inclusions from Dabbahu, which are H₂O-rich in the
262 range of 3 to 5.8 wt.%. However, high magmatic H₂O is consistent with other findings in
263 peralkaline rocks (Maksimov, 2008; Blundy and Cashman, 2008; Field et al., 2012) in Dabbahu
264 near SEME station. Touret and Van Den Kerkhof [1986] gave a possible evidence for the presence
265 and the nature of free fluid phase to be found at depths down to 35 km below the surface of the
266 continent. The existence of melt and/or fluid beneath SEME makes the anomalous Moho difficult
267 (since observed V_p value beneath SEME is very small compared to global Moho velocity) to
268 identify only using V_p value from the velocity versus depth plot alone unless a priori V_p value is
269 used. However, the source of this low velocity beneath the station could be the release of magmatic
270 fluid such as CO₂ and H₂O from the mantle to the crust (Bucher-Nurminen, 1990). This inference
271 may get accompanied by either the accumulation of H₂O-rich fluid composition, heat and/or
272 volatile in the lower crust (Zandt and Ammon, 1995) which percolates upwards through the crust
273 by a porous-media or fracture depending on the melting of the crust. If the temperature of the lower
274 crust is high enough, this fluid could induce partial melting of material of suitable composition
275 (Amundsen, 1987; Bucher-Nurminen, 1990). This has a similar inference with the work of
276 Thompson and Connolly [1990] as fluid accumulation is a common phenomenon in the lower
277 crust. The maximum fluid content transferred from the mantle to the lower crust depends on both
278 Moho temperature and depth (Bucher-Nurminen, 1990). The existence of the partial melt in the
279 lower crust beneath Afar is supported by previous geophysical studies such as Hammond et al.

280 [2011], Guidarelli et al. [2011], Desissa et al. [2013] and Hammond [2014], which are interpreted
281 as stored melt in the sill-like features in the lower crust.

282 The Moho depth beneath SEME is about $26\text{km} \pm 2\text{km}$ and the result is very close to an inference
283 given by (Berckhemer, 1975). Gravity modeling in the southern central part and margin of the Red
284 Sea Rift shows that there is a possibility of partial melt existence at a depth of 8.5 to 25 km
285 (Johnson, 2012; Lewi et al., 2015). In their work, Knox et al. (1998) used the inversion of Rayleigh
286 wave dispersion and found low S-wave velocities (0.2—0.8 km/s) and that were inferred as partial
287 melt.

288 **4 Global comparisons and the major points**

289 Crustal P wave seismic velocities and crustal thicknesses observed beneath HARA and DILA are
290 similar to the global average crustal structure (Christensen and Mooney, 1995), suggesting that the
291 crust on the Southeastern side of the MER and Afar have not been significantly modified by plate
292 thinning and intrusion. The relatively normal seismic velocities in the crust make a relatively sharp
293 contrast with what is slower than the global average upper mantle seismic velocities.

294 ANKE is characterized by a 42 km thick crust that has higher seismic velocities than HARA and
295 DILA, and thus higher than the global average. Specifically, the observed velocities at 5—10 km
296 depth is ~ 0.5 to 1 km/s faster than the globally most commonly observed range of 6—6.5 km for
297 this depth range [e.g., Christensen and Mooney, 1995]. In addition, the seismic velocities of
298 $\sim 7\text{km/s}$ in the 30-40 km deep lower crust is towards the higher end of the globally most common
299 velocities of 6.5—7.1 km/s. Since frozen mafic rock has a seismic velocity of over 7 km/s, the
300 modification of the crust at ANKE by mafic intrusion provides a simple explanation for high
301 seismic velocity through lower crust underplating and recent magmatic activity beneath the
302 Northwestern plateau (Mackenzie et al., 2005). Intruded upper crust is consistent with observed
303 mafic dykes exposed near the surface along the western margin of Afar (Wolfenden et al., 2004).
304 Elevated lower crustal velocities are consistent with the presence of frozen lower crustal, stacked
305 sill complexes imaged beneath the western margin of the MER and the Northwestern plateau by
306 controlled source and passive seismic techniques in project EAGLE (Stuart et al., 2006; Keir et
307 al., 2009). In light of previous results, the observed crustal structure beneath ANKE is consistent
308 with the interpretation that the crust beneath the northwestern side of Afar and the MER has been

309 significantly modified by magmatism. Our results suggest that intrusion into the upper crust along
310 the western Afar margin may be more significant than previously thought.

311 The seismic velocity structure of SEME is quite different to the other seismic stations, and to the
312 global average continental crust. Seismic velocities in the uppermost crust are exceptionally low,
313 consistent with the presence of a sedimentary basin including evaporites. Seismic velocities at
314 10—13 km depth are over 7 km/s, and consistent with the presence of upper crustal mafic intrusion.
315 Seismic velocities are then anomalously low in the 15 - 25 depth range. In light of independent
316 constraints of high conductivities observed in MT studies at the same depth range and interpreted
317 as evidence for the presence of partial melt (Dessiese et al., 2013), we interpret the low seismic
318 velocities in the lower crust beneath SEME to be caused by the presence of partial melt. The
319 interpretation is strengthened by InSAR analysis which shows a broad zone of subsidence near
320 SEME that has been modeled as magma withdrawal from a lower crustal sill complex (Grandin et
321 al., 2009). Our observations, in light of previous constraints from Afar as well as from seismic
322 refraction images across rifted continental margins such as the eastern north Atlantic (White et al.,
323 2008), provides evidence that the lower crust is an important melt storage region during the
324 breakup of continents.

325 **Conclusions**

326 By applying frequency domain deconvolution techniques followed by a linearized-iterative
327 inversion on the radial component of receiver functions we obtained Moho depths of 26 ± 2 km
328 for SEME, 36 ± 2 km for DILA, 38 ± 2 km for HARA and 42 ± 1.7 km for ANKE. We achieved
329 a reasonably good fit between the observed and synthetic RFs by employing high quality seismic
330 dataengaged in plain usage of water level and Gaussian filter methods. Though additive noises
331 preceding direct P pulse caused uncertainty, the seemingly misfit result obtained at SEME and
332 HARA stations could be due to local sediments overlaid on thick high velocity material in addition
333 to the filters' pre-signal remnants and ambient seismic noises. The lowest Moho depth is estimated
334 for SEME station, which implies a thinned crust while the highest crustal thickness is achieved for
335 ANKE that lies along the Northwestern plateau margin. Our results agree with previous
336 observations which intensify the hypothesis that Moho depths estimated for stations that lie within
337 the rift and rift margins are lower than those located in the plateaus. Our RF inversions show a low
338 velocity gradient at a depth of about 16km at SEME station, interpreted as evidence of lower

339 crustal storage of partial melt, which might have a defined contribution for the transition from
340 continental to incipient oceanic rifting. A relatively high velocity zone probably due to cooled
341 magma from previous dike or sill intrusions is found at a depth of ~7 to 20 km at ANKE station.
342 In contrast, Vp values beneath HARA are consistently high throughout the crust, and this might
343 be interpreted as low crustal heterogeneity.

344 Acknowledgments

345 The authors would like to thank the technical staff of the Institute of Geophysics Space Science
346 and Astronomy (IGSSA) Addis Ababa University. The Figs. 1 and 3 were done using Generic
347 Mapping Tools (GMT) [Wessel and Smith, 1991]. We thank Charles Ammon for providing the
348 inversion codes freely. We are grateful to the International Science Program of Uppsala University
349 for financing the Ethiopian Seismic Station Network (ESSN). D.K. is supported by NERC grant
350 NE/L013932/1.

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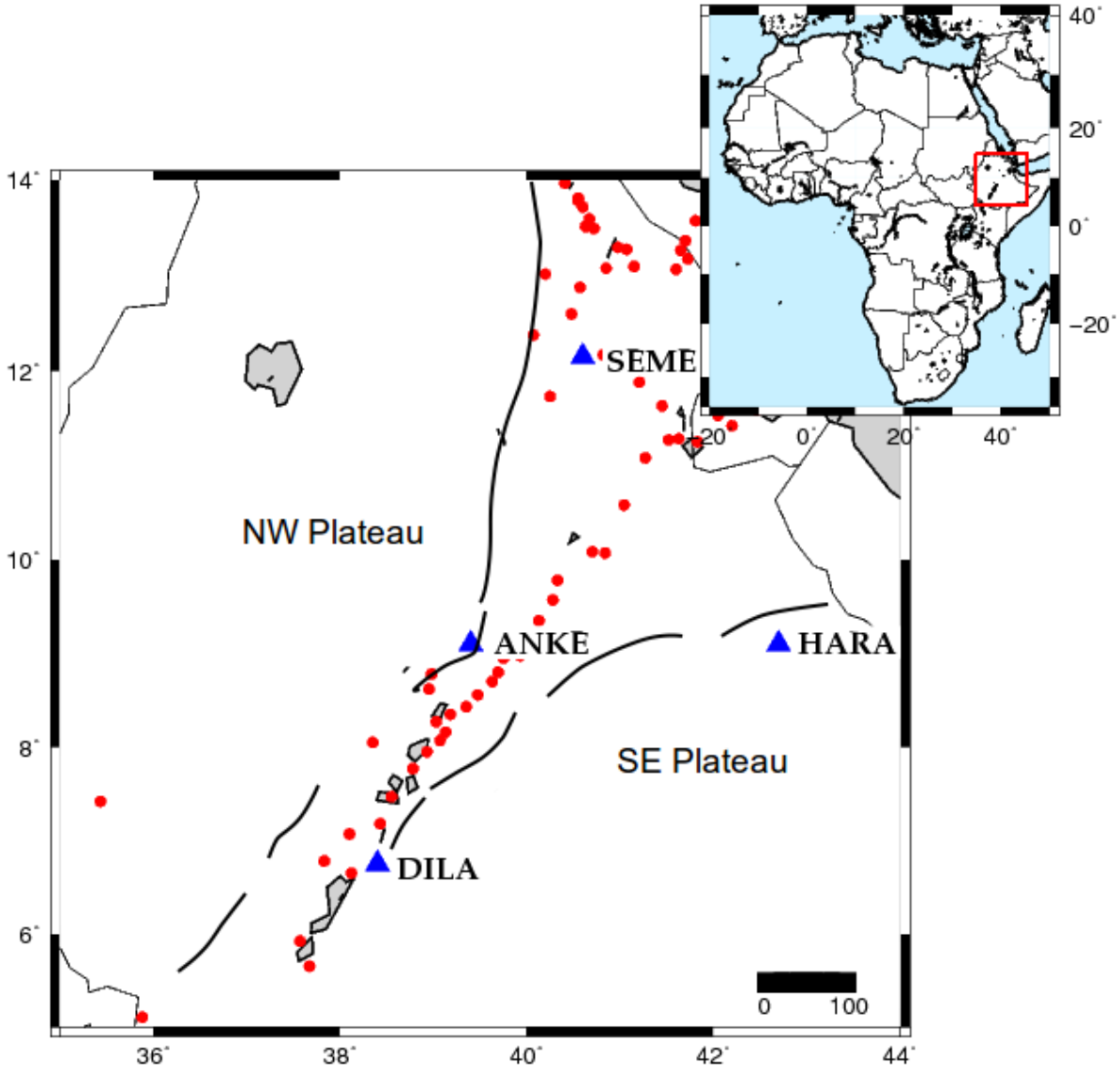
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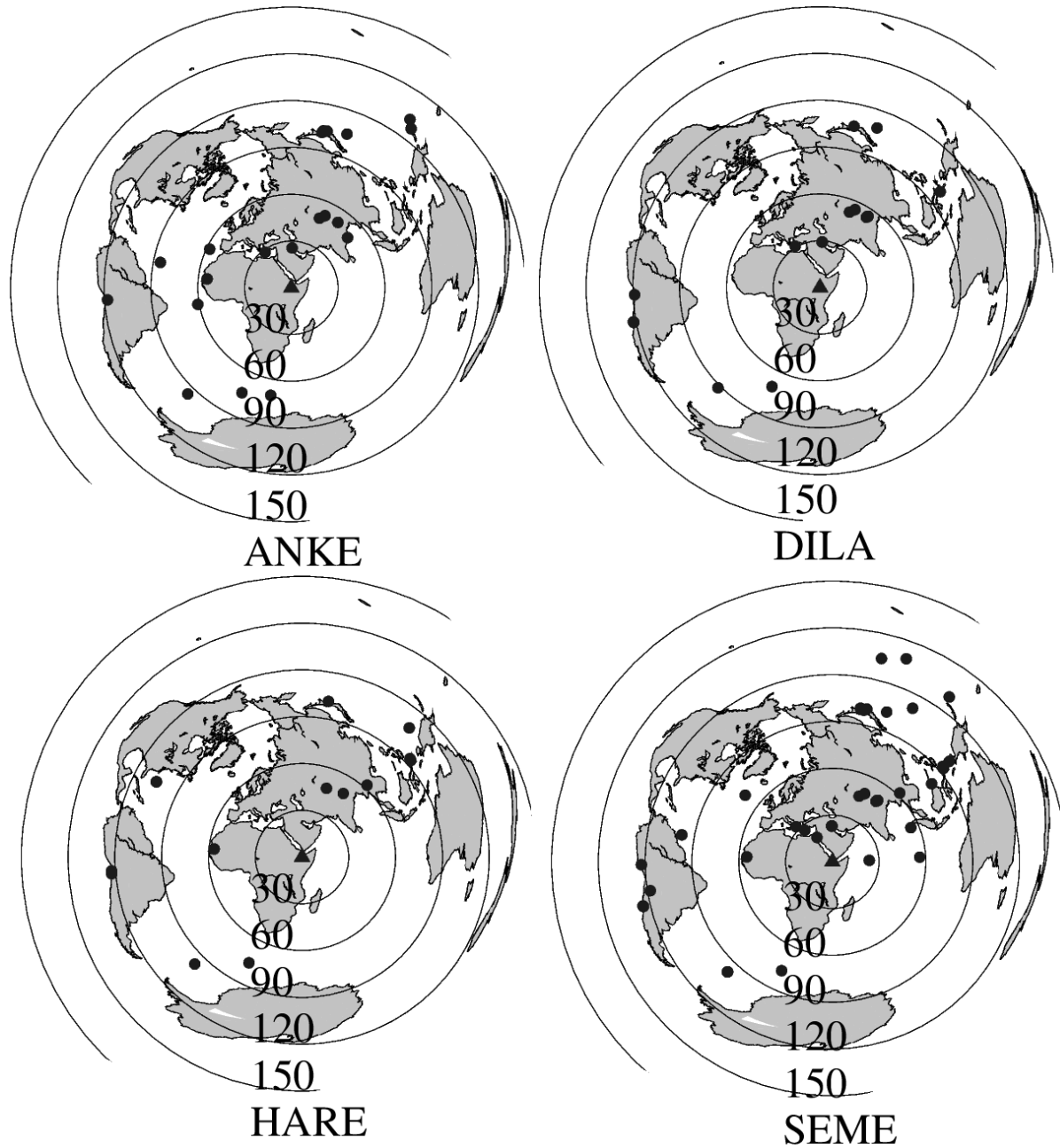
515 **Figures**



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518 Figure 1. Location of the four Ethiopian permanent broadband seismic stations (blue triangles) for which receiver
 519 functions and P wave velocity models were estimated. The black thick lines are border faults that separate the NW
 520 plateau and the SE plateau from the Main Ethiopian Rift and Afar. The red circles represent the location of volcanoes.
 521 Inset: The red rectangle shows the location of the study area on the African continent



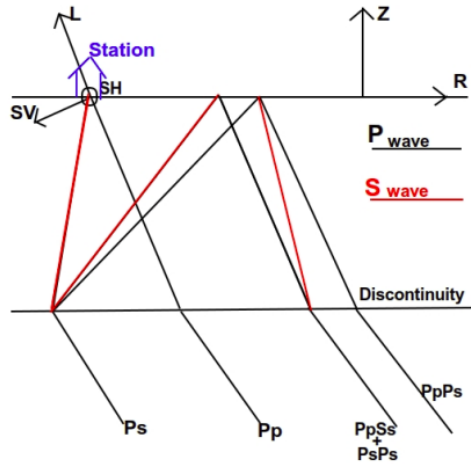
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523 Figure 2. The solid black circular symbols are the locations of the teleseismic earthquakes that occurred between
 524 September 2014 and August 2015 with a magnitude greater than or equal to 5.5mb and within epicentral distances
 525 between 30°and 100°. These earthquake data were recorded by the Ethiopian permanent broad band seismic network
 526 of stations (ANKE, SEME, HARA and DILA) and collected by the team led by the Institute of Geophysics Space
 527 Science and Astronomy (IGSSA), Addis Ababa University. The black triangles in the center of the internal circles are
 528 locations of each seismic station. The teleseismic earthquake's information (latitude, longitude and magnitudes) were
 529 taken from Global Earthquake catalog of the International Earthquake Information Center.

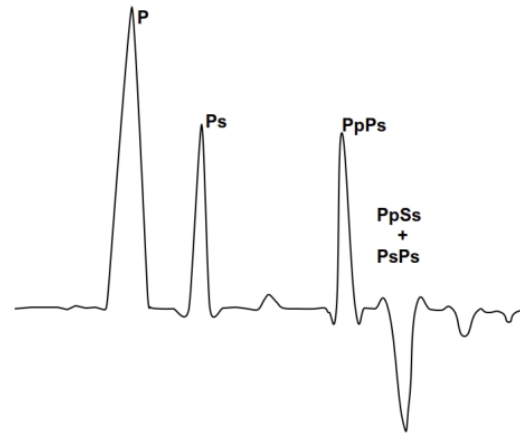
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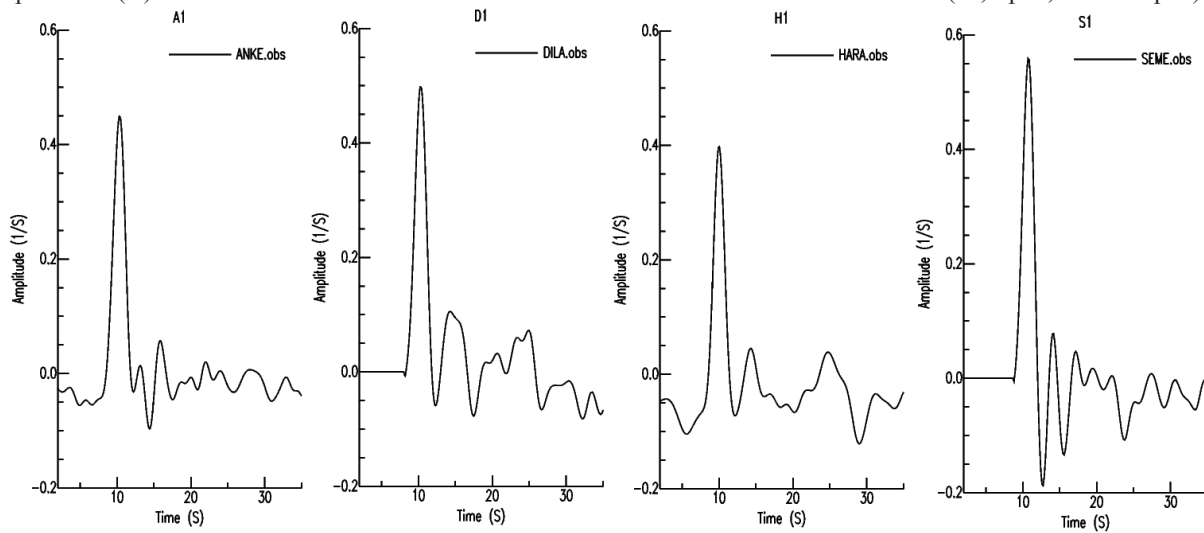
(A)



(B)

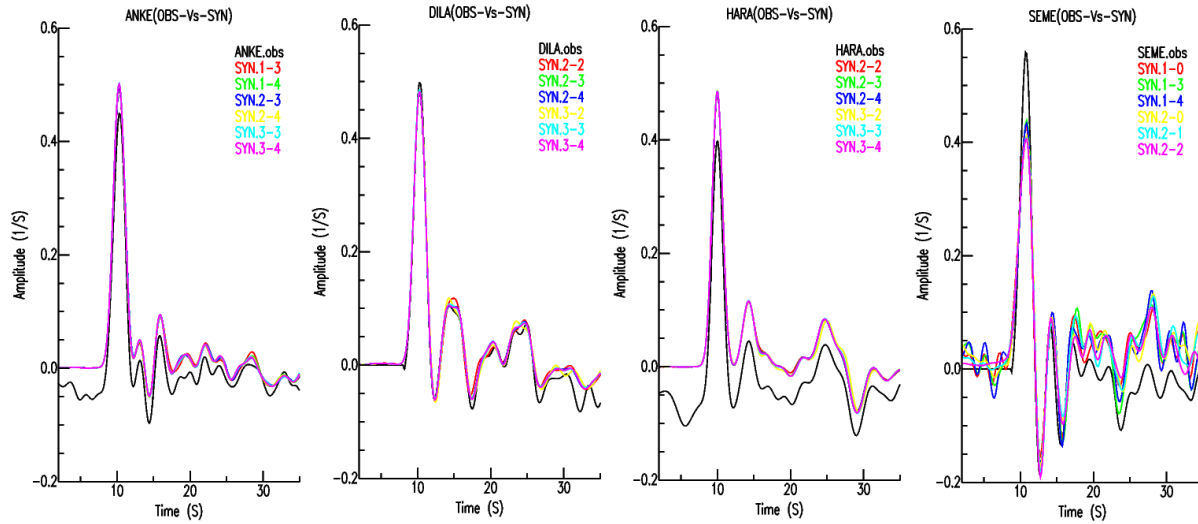
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541 Figure 3. (A) is a graphical representation ray diagram of receiver functions identifying the major P- to S converted
 542 phases that illustrates the direct P, Ps, PpPs and PpSs phases that comprises the receiver function for a single half-
 543 space and (B) is a receiver function that shows the direct P-wave and the reverberations (Ps, PpPs, PsPs + PpSs).



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545 (A)

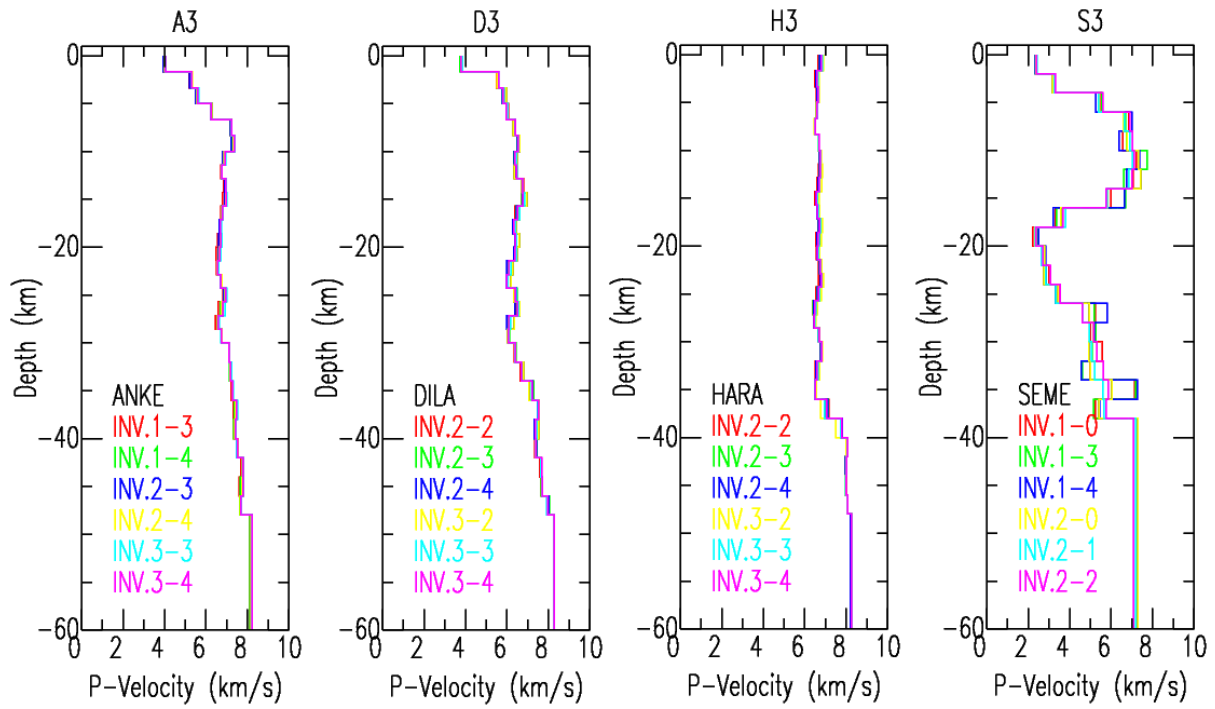


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547 (B)

548 Figure 4. (A)The black wave form represented by A1, D1, H1 and S1 are the observed waveform receiver functions
 549 obtained from the deconvolution of the radial component from the horizontal component waveform of the teleseismic
 550 earthquakes recorded at ANKE, DILA, HARA and SEME permanent broad band seismic stations. (B) Observed
 551 receiver functions (Figure A) fit with the synthetic coloured receiver functions obtained from iterative frequency
 552 domain water level deconvolution technique. In all plots the delay time given for the P wave to show contrast is 10
 553 seconds and Ps marked in each plot is the conversion from direct P wave to S wave. The coloured waveform models
 554 are the synthetic receiver functions obtained by varying the smoothing weight parameters using the program
 555 "smthin".

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Figure 5. P wave velocity models obtained by inverting receiver functions for the seismic stations A3 (ANKE), D3 (DILA), H3 (HARA) and S3 (SEME). The inversion results in these plots were obtained using programs by Ammon[1991]. The Moho is interpreted as the depth at which there is a sharp increase in seismic velocity from ~6.5 km/s in the lower crust to 7—7.5 km/s in the upper mantle.