

## Supersolid Phases of Bosonic Particles in a Bubble Trap

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Confinement can have a considerable effect on the behavior of particle systems and is therefore an effective way to discover new phenomena. A notable example is a system of identical bosons at low temperature under an external field mimicking an isotropic bubble trap, which constrains the particles to a portion of space close to a spherical surface. Using path integral Monte Carlo simulations, we examine the spatial structure and superfluid fraction in two emblematic cases. First, we look at soft-core bosons, finding the existence of supersolid cluster arrangements with polyhedral symmetry; we show how different numbers of clusters are stabilized depending on the trap radius and the particle mass, and we characterize the temperature behavior of the cluster phases. A detailed comparison with the behavior of classical soft-core particles is provided too. Then, we examine the case, of more immediate experimental interest, of a dipolar condensate on the sphere, demonstrating how a quasi-one-dimensional supersolid of clusters is formed on a great circle for realistic values of density and interaction parameters. Crucially, this supersolid phase is only slightly disturbed by gravity. We argue that the predicted phases can be revealed in magnetic traps with spherical-shell geometry, possibly even in a lab on Earth. Our results pave the way for future simulation studies of correlated quantum systems in curved geometries.

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Computing the equilibrium properties of quantum many-body systems remains a main objective of contemporary physics. In this respect, ultracold atoms provide a framework where geometry and interactions can be tuned almost at will, allowing one to test fundamental many-body theories [1,2]. A system of atoms (loosely) confined to an ellipsoidal surface [3,4] represents a typology that only recently has started to be explored. To achieve this goal, a quantum gas is loaded into a shell trap, where atoms are subject to a quadrupolar field “dressed” by a radio frequency (rf) field [5–9]. In the limit of slow atomic motion in a strong magnetic field, the effective potentials of the internal states are the position-dependent eigenvalues of the Hamiltonian consisting of the bare potentials and the coupling term [10]. For atoms in the upper dressed state, resonance is reached at the surface of an ellipsoid. However, to let atoms explore the full surface, experiments must be performed in outer space [11–13] or adopt some gravity compensation mechanism [14,15]. Coherent and isotropic shells of atoms slowly expanding in microgravity can be generated too [16].

The realization of shell-shaped condensates has fueled a renewed interest in the problem of quantum particles in curved geometries [17], and particularly in the quantum phases that free and interacting bosons can exhibit. Several recent works have investigated Bose-Einstein condensation as well as the superfluid-to-normal fluid transition on a sphere [18–22], while others have studied the dynamics and thermodynamics of the condensate itself [23–26]. Most of these works have considered weakly interacting particles, although some have looked into the condensed phases arising from dipolar interactions [26–29].

While experiments have so far been performed in the same dilute limit, bubble traps open the exciting prospect of investigating the physics of strongly correlated quantum particles in curved spaces, with all the advantages brought by ultracold-atom setups. For example, ultracold atoms may serve as quantum simulators to test fundamental physics; therefore, studying the effects of curvature in a controlled environment can be of interest to other fields ranging from cosmology to biology [30,31]. Similarly, we

can envisage the possibility to control the self-assembly of a many-body system through a convenient choice of the confining surface, i.e., of the shape of the ensuing geometric potential [17]. Equally important is to figure out experiments on curved many-body systems that could be accomplished on Earth, i.e., without the necessity of compensating gravity [32]. In this respect, the most awaited developments concern dipolar atoms [33], which have already been examined in flat space and harmonic traps [34,35] and are known to give rise to supersolid phases [36–39]. At the same time, it will be interesting to see whether the confined geometry of bubble traps can stabilize the supersolid phase in small systems of Rydberg-dressed atoms [40–44].

In this Letter, we use the path integral Monte Carlo (PIMC) simulation method to give a first glimpse of the equilibrium phases that can arise in a shell-trapped system of identical bosonic particles. In particular, we will provide the first compelling evidence of supersolid order for two distinct instances of the interaction potential. First, we look at the penetrable-sphere model as an example of soft-core potential and a paradigmatic interaction in condensed matter physics that gives rise to both superfluidity and clusterization [45–48]; for this interaction, we compare our results with various benchmarks. Then, we investigate the effects of a dipole-dipole interaction, more closely related to experiments, and we show, for a realistic choice of parameters, that a supersolid cluster phase indeed occurs in shell geometry. Remarkably, this supersolid is resistant to the gravity of Earth.

To keep contact with the experiments, we simulate spinless bosons in three-dimensional space under the constraint of an external potential analogous to that realized in the lab. For  $N$  particles with mass  $m$ , the Hamiltonian reads

$$H = \sum_{i=1}^N (-\lambda \nabla_{\mathbf{r}_i}^2 + u_{\text{ext}}(|\mathbf{r}_i|)) + \sum_{i<j} v_{\text{int}}(\mathbf{r}_i - \mathbf{r}_j), \quad (1)$$

where  $\mathbf{r}_i$  is the position of the  $i$ th particle,  $\lambda = \hbar^2/2m$ ,  $v_{\text{int}}$  is the (possibly anisotropic) interaction potential, and

$$u_{\text{ext}}(d) = (u_0/\Omega) \sqrt{(d^2 - \Delta)^2/4 + \Omega^2} - u_0 \quad (2)$$

is the external potential appropriate to a spherically symmetric bubble trap [24] centered at the origin. In Eq. (2),  $\Delta$  and  $\Omega$  are square-length parameters related to the detuning and Rabi frequency of the rf field, respectively, while  $u_0$  characterizes the harmonic trap prior to dressing. The potential (2) interpolates between the filled sphere and the spherical shell [see Fig. 1(g)]. In the thin-shell limit,  $u_{\text{ext}}$  becomes harmonic around the minimum at  $\sqrt{\Delta}$ , which thus defines the radius  $R$  of the reference sphere. As  $R$  increases, the potential minimum gets more and more pronounced,

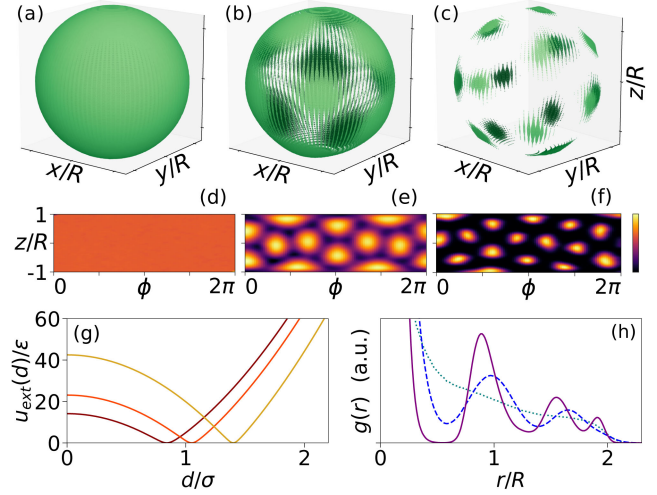


FIG. 1. Soft-core bosons in an isotropic bubble trap. (a)–(c) Particle density for  $N = 120$ ,  $T = 0.125$ ,  $R = 1.4$ ,  $u_0 = 2$ ,  $\Omega = 0.0441$ , and  $\lambda = 0.5$  (a), 0.16 (b), 0.01 (c). The size of the points is proportional to the density averaged along the radial direction, and the shading is a guide for the eye: Brighter points are closer to the observer. (d)–(f) Area-preserving projections of the densities in (a)–(c). Brighter colors indicate a larger density (log scale). (g) External potential (2) for  $u_0 = 2$ ,  $\Omega = 0.0441$ , and  $R = 0.84$  (brown), 1.05 (red), and 1.4 (gold). (h) Pair distribution function for the same parameters in (a)–(c):  $\lambda = 0.5$  (blue dotted line), 0.16 (blue dashed line), 0.01 (purple solid line).

until particles become effectively confined to a spherical surface.

In our PIMC simulations [49], we employ the worm algorithm [50] to sample the equilibrium statistics of bosons at finite temperature and estimate the superfluid fraction,  $f_s$  [51]. To deal with the strong spatial constraint due to (2), we introduce a biased version of the PIMC method [51]. We also perform classical Monte Carlo (MC) and molecular dynamics (MD) simulations to probe the classical limit in the soft-core case.

We begin by considering bosons interacting through the soft-core potential:

$$v_{\text{int}}(\mathbf{r}) = \epsilon \theta(\sigma - r), \quad (3)$$

where  $\epsilon > 0$ , and  $\sigma$  is the core diameter. In the discussion of the soft-core model,  $\epsilon$  and  $\sigma$  are taken as units of energy and length, respectively; temperatures are expressed in units of  $\epsilon/k_B$ . Because of the peculiar nature of the repulsion, at high density particles are gathered in droplets or clusters [77]. Employing mean-field theory at  $T = 0$ , the authors of Ref. [78] explored various possible arrangements of clusters on the sphere, finding the configuration of lowest enthalpy as a function of the radius. The evidence of supersolid phases is, however, not conclusive: Condensation of clusters is assumed, not derived, while there is no guarantee that the true ground state has been identified. Using, for the first time, *ab initio* PIMC simulations, we provide for the

same interaction (3) conclusive evidence of supersolid order. Moreover, we show how the supersolid withstands finite temperature in a realistic bubble trap, i.e., beyond a purely two-dimensional setup.

First, we investigate the behavior of the system at a temperature  $T \ll 1$ , choosing  $\Omega = 0.0441$  (an arbitrary value much smaller than  $\Delta$ ; see below). In Figs. 1(a)–1(f), we give a graphical account of the structures found for decreasing  $\lambda$  at  $T = 0.125$ , which are indicative of a superfluid-supersolid-insulator “transition,” similar to what occurs on a plane [79]. Specifically, we perform a size scaling analysis of  $f_s$  at fixed  $N/(4\pi R^2) = 4.87$  for two strengths of the trap potential to see how the system approaches the planar limit. On a plane, the phases were characterized in terms of the dimensionless interaction strength  $\alpha = m\rho\sigma^4\epsilon/\hbar^2 = \rho\sigma^4\epsilon/(2\lambda)$  (for a number density  $\rho = 4.4$ ), finding that the superfluid-supersolid transition occurs at  $\alpha \approx 13$ , whereas the system becomes insulating at  $\alpha \approx 22$  [79]. In both two and three dimensions, the former transition is marked by a jump in  $f_s$  [45]. Our results are reported in Fig. 2. In the strongly confined case ( $u_0 = 50$ , top panel), the planar behavior is already recovered for  $N = 120$  particles, while for smaller sizes we find a smooth crossover. For a weaker external potential ( $u_0 = 2$ , bottom panel), the transition to supersolid is milder and shifted toward higher  $\alpha$  values; moreover, convergence to the planar limit is much slower. In both cases, the system is a density wave at large  $\alpha$ ; as  $\alpha$  is reduced, quantum fluctuations increase, eventually leading to a disruption of polyhedral order which is faster when more freedom is given to particles in the radial direction, i.e., for  $u_0 = 2$ . The same sequence of transitions is reflected in the shape of the pair distribution function  $g(r)$  [see Fig. 1(h) and [51]]. As

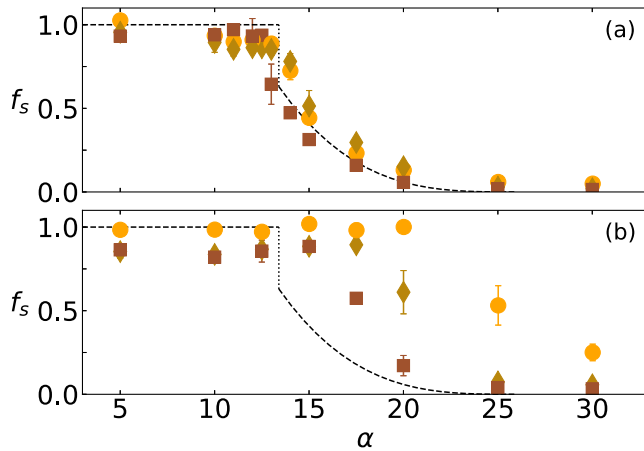


FIG. 2. Superfluid fraction across the superfluid-supersolid-insulator transition plotted as a function of  $\alpha$  for  $T = 0.125$  [the  $\rho$  entering the expression of  $\alpha$  is  $N/(4\pi R^2)$ ]. (a)  $u_0 = 50$ . (b)  $u_0 = 2$ . Different symbols correspond to different system sizes:  $N = 40$  and  $R = 0.808$  (circles),  $N = 80$  and  $R = 1.143$  (diamonds),  $N = 120$  and  $R = 1.4$  (squares). The dashed lines refer to the planar limit [80].

long as  $f_s > 0$ ,  $g(r)$  is nonzero at any distance, which is consistent with a condensate wave function being nonzero everywhere on the sphere. Instead, in a normal solid near zero temperature,  $g(r)$  would ideally be zero in the interstitial region between two successive shells of neighbors. The deviations from this behavior at moderate to large distances are a temperature effect, ultimately due to the finite extension of the clusters.

Next, we fix the number  $N$  of particles (120) and the temperature  $T$  (0.5), and set  $u_0 = 2$  and  $\Omega = 0.0441$  [we require  $\Omega \ll 1$ , so as to have  $u_{\text{ext}}(0) \approx \Delta/\Omega \gg 1$  for  $R \approx 1$ ]. We collect in a diagram the equilibrium arrangements for various  $\lambda$  and  $R$  [Fig. 3(a)]. On account of the  $f_s$  value and the evidence or not of polyhedral order, we can reasonably distinguish between superfluid, supersolid, and normal-solid states. Throughout the “solid” regions, the number  $N_c$  of clusters may vary, but clusters are invariably found at the vertices of a regular or semiregular polyhedron: for example, an octahedron for  $N_c = 6$ , a square antiprism for  $N_c = 8$ , and an icosahedron for  $N_c = 12$  [51]. For some  $(\lambda, R)$  pairs, the cluster structure agrees with those predicted in [78] (for example, the icosahedral structure at  $\lambda = 0.16$  and  $R = 1.4$ ). However, for several other pairs, the equilibrium configuration is a novel cluster phase not seen before. Moreover, contrary to the mean-field prediction, we see that decreasing  $\lambda$  at fixed  $R$  leads to

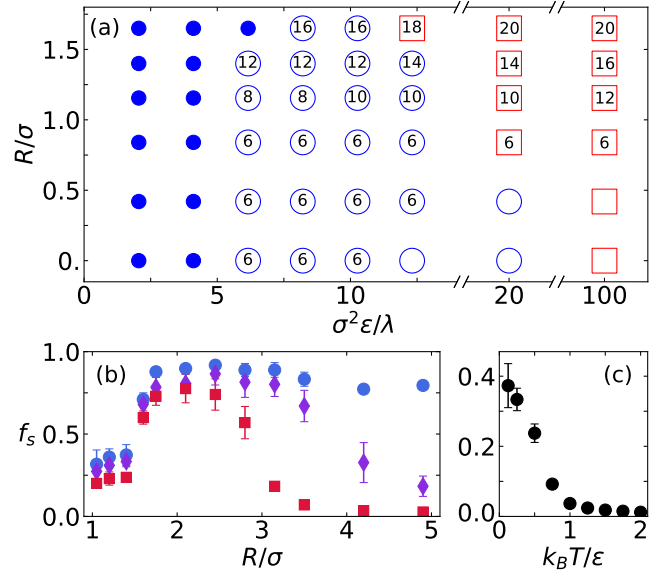


FIG. 3. (a) A diagram showing for  $N = 120$  and  $T = 0.5$  the number  $N_c$  of clusters as a function of  $\lambda^{-1}$  and  $R$ . The blue dots are state points where the system is superfluid. Open circles and squares mark supersolid and normal-solid states, respectively. When  $\lambda$  and  $R$  are both small,  $N_c$  depends on the initial conditions (see more in [51]); therefore, no number is printed in the symbol. (b) Superfluid fraction  $f_s$  plotted as a function of  $R$  at fixed  $\lambda = 0.16$  and  $u_0 = 50$  for  $T = 0.5$  (red squares), 0.25 (purple diamonds), and 0.125 (blue circles). (c)  $T$  dependence of  $f_s$  for  $R = 1.4$  and  $\lambda = 0.16$ .

transitions between equilibrium configurations with different numbers of clusters. As we increase the radius, the supersolid region shrinks, while the number of clusters grows. At large radii, the particles assemble in smaller clusters, making it more difficult for the system to sustain superfluidity; eventually, above a certain  $R$  the system is so dilute that clusters are washed away for every  $\lambda$ .

In the  $T \rightarrow 0$  limit, the superfluid phase is stable at all large radii. Upon heating, similar to flat space, a superfluid-to-normal-fluid transition eventually occurs. This is illustrated for  $\lambda = 0.16$  in Fig. 3(b), where we plot the superfluid fraction as a function of  $R$  at different temperatures. As  $T$  goes up,  $f_s$  is gradually reduced until it vanishes, starting from larger  $R$  values. At high enough temperatures,  $f_s$  is also depleted for the supersolid, as we show in Fig. 3(c) for  $R = 1.4$  and  $\lambda = 0.16$ : Interestingly, however, polyhedral order is preserved throughout the range of temperatures.

When  $\lambda \rightarrow 0$ , the number of clusters approaches a value dependent on  $R$ ; this regime corresponds to the classical limit, regardless of  $R$  and  $T > 0$ . To make sure that this is indeed the case, we have considered the classical counterpart of the quantum problem. The first observation of cluster phases in classical particles on a sphere goes back to Ref. [81], where density-functional-theory calculations are presented. Here, to keep the same level of accuracy of the quantum treatment, we carry out extensive MC and MD simulations of classical soft-core particles, using the same parameters of the quantum simulations. The results expressed in terms of the number and final arrangement of clusters are reported in [51] for various  $R$  values. Except for small radii, where the uncertainty in  $N_c$  is large, the aggregates formed by quantum particles have a comparatively smaller number of clusters. This is an effect of quantum delocalization, which causes the effective diameter of bosons to be larger than  $\sigma$ .

We now move to examine the behavior of dipolar bosons in the thin-shell limit. As is common in experiments, we assume atoms to be polarized along  $\hat{z}$ . We work with a number density on the sphere of about  $4 \mu\text{m}^{-2}$ , of the same order as that used in experimental realizations [11]. The (anisotropic) interaction potential reads

$$v_{\text{int}}(\mathbf{r}) = v_{\text{HS}}(r) + \frac{\mu_0 d_m^2}{4\pi} \frac{1 - 3 \cos^2 \theta}{r^3}, \quad (4)$$

where  $\mu_0$  is the vacuum permeability,  $d_m = 9.93\mu_B$  is the magnitude of the dipole moment of a  $^{164}\text{Dy}$  atom,  $\mu_B$  is the Bohr magneton, and  $\cos \theta = \hat{\mathbf{r}} \cdot \hat{\mathbf{z}}$  [82]. Finally,  $v_{\text{HS}}(r)$  is the hard-sphere potential with core diameter equal to the scattering length  $a$  [51].

We simulate the system for various  $a$ , considering values up to  $10^3 a_0$ , with  $a_0$  the Bohr radius, thus much smaller than the sphere radius taken to be  $R = 2.6 \mu\text{m}$  (or  $R \approx 50 \times 10^3 a_0$ ). Our results for  $N = 360$  and  $T = 1$  nK are illustrated in Fig. 4. As is clear from Eq. (4), particles

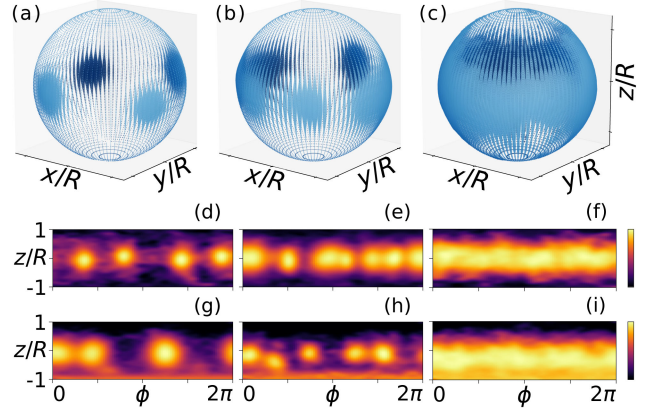


FIG. 4. Dipolar atoms on a sphere. (a)–(c) Particle density for  $N = 360$ ,  $T = 1$  nK, and  $R = 2.6 \mu\text{m}$ , at  $a/a_0 = 5$  (a),  $a/a_0 = 50$  (b), and  $a/a_0 = 350$  (c). The density increases with the size of the points, while colors are a guide for the eye. (d)–(f) Area-preserving projections of the densities in (a)–(c). (g)–(i) Effect of gravity. Brighter colors indicate larger values of the density (log scale).

attract each other along the  $z$  direction and repel each other along  $x$  and  $y$ . As long as  $a$  is much smaller than  $R$ , particles move away from the poles of the sphere and bunch together around the equator. Indeed, up to  $a \approx 150a_0$ , particles form clusters lined up along the equator, as seen in Figs. 4(a) and 4(b). At variance with what is observed for dipolar clusters in trapped Euclidean geometries [36], the cluster phase on the sphere excludes zigzag configurations: In Figs. 4(a) and 4(b), clusters remain on the equator due to the lack of a significant repulsion along  $z$  that could counteract the mutual attraction. Unless  $a$  is very small, the superfluid fraction for rotations around  $z$  is finite and significantly larger than  $f_s$  for rotations around  $x$  and  $y$ , thus qualifying this state as supersolid [Fig. 5(a)] [83]. For

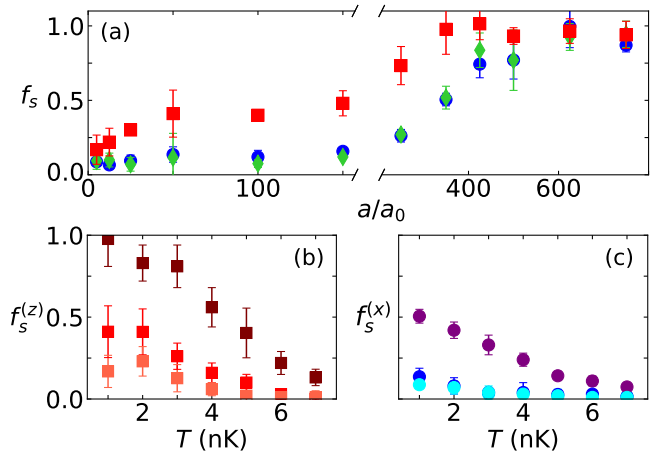


FIG. 5. (a) Superfluid fractions  $f_s^{(x)}$  (blue circles),  $f_s^{(y)}$  (green diamonds), and  $f_s^{(z)}$  (red squares) as a function of  $a$ . (b), (c) Thermal behavior of  $f_s^{(z)}$  and  $f_s^{(x)}$ , for  $a/a_0 = 5$  (pink and light blue), 50 (red and blue), and 350 (brown and purple).

larger  $a$ , due to the increased particle exclusion, clusters grow in size and merge together forming a ribbon wrapped around the sphere; see Figs. 4(c) and 4(f). This specific configuration is consistent with the density profile seen in [26,27]. Beyond  $a \approx 200a_0$ , the ribbon is homogeneous; however, between  $a = 150a_0$  and  $a = 200a_0$ , there is a wide crossover region where the clusters are still present as azimuthal density modulations. The modulated-to-unmodulated transition is shifted to slightly larger  $a$  compared to other trapped geometries (see, e.g., [85] and references cited therein) because of the enhanced stability of spherically confined clusters. The superfluid fraction along  $z$  is still significantly larger than in the other directions. Finally, for very large values of  $a > 500a_0$ , a homogeneous superfluid state emerges, similar to the one shown in Figs. 1(a) and 1(d). Simulations at different values of  $N$  but fixed density show no significant scaling, a strong indication that the same supersolid phase persists at larger numbers of particles, up to experimental values. The temperature analysis indicates that the supersolid behaves differently from the superfluid ribbon [see Figs. 5(b) and 5(c)]. In the former case,  $f_s^{(z)}$  remains nonzero up to  $\approx 5$  nK due to a nonzero superfluid signal of the system around  $z$ . As for  $f_s^{(x)}$ , the signal is smaller and substantially constant as a result of the thickness of the supersolid in the latitudinal direction. Instead, the ribbon remains superfluid up to much higher temperatures, while keeping its anisotropic character for all  $T$ .

Figures 4(g)–4(i) depict simulations carried out by including the effect of gravity on  $^{164}\text{Dy}$  atoms. Interestingly, for small enough  $a$  the atoms still form a ring of clusters perpendicular to  $z$ , although their center of mass is no longer at  $z = 0$ , but slightly shifted downward. For larger values of  $a$ , even though part of the atoms are found near the south pole, a shifted superfluid ribbon is still evident. This is to be contrasted with what happens when we remove the dipolar component of the interaction, but still keep the hard-core repulsion: In this case, all particles are pushed by gravity to the bottom of the trap [51].

To conclude, we have investigated the equilibrium phases of identical bosons in a shell trap. As the bubble expands, the system switches from three to two dimensional in a continuous fashion. In this peculiar setting, the PIMC algorithm needs an *ad hoc* modification that we have discussed at length in [51]. With this tool at hand, we have considered two species of bosons, i.e., soft-core atoms and dipolar atoms, providing evidence at small nonzero temperature of two unconventional supersolid behaviors. We have argued that the realization of these phases is within reach of present-day technology. Very significantly, we have shown that the dipolar supersolid on the sphere is robust to the effects of gravity. This surprising result indicates that it could be probed experimentally even without sending the apparatus to space. More generally, the ideas underlying our original implementation of the PIMC algorithm on the

sphere can profitably be replicated for other curved surfaces, allowing us to gain insight into the peculiarities of bosons in spaces where the geometric potential is nonzero (see more in the last section of [51]).

In this Letter, we have focused on demonstrating the existence of a supersolid phase on the sphere. Using the same numerical techniques, it would be also possible to investigate the nature of the superfluid phase on the sphere, e.g., Berezinskii-Kosterlitz-Thouless behavior or reentrant effects [85,86]. Our work might also be relevant to supersolid phases in the crust of neutron stars [87]. The dynamical properties of shell-trapped bosons are another open problem and an active area of research [88]. For example, it would be interesting to see whether the superfluid ribbon in Figs. 4(c) and 4(f) can support persistent currents in the same way as does a superfluid ring [89–91].

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