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## Time-space variations in the East African Rift magmatism: the role of different mantle domains

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### Abstract

The East African Rift System (EARS) is the classic example of an active continental rift where extensional tectonics and lithospheric thinning have been closely associated to the generation of large volumes of magmas and represents the environment with the largest range of erupted magma types all over the world. The geochemical signature of erupted magmas testifies the involvement of different mantle domains and depths (i.e., subcontinental lithosphere, asthenosphere and deeper mantle sources). Our aim is to investigate the variable contribution of different mantle domains in the genesis of the EARS magmas through space and time, considering not only the geochemical signature of erupted magmas but also the geochemical message of mantle xenoliths. The main goal is to provide a large-scale view of the common process driving the origin of magmas in the EARS beyond the local peculiarities linked to specific settings. To this aim, we screened an exhaustive geochemical database of basalts and mantle xenoliths from the EARS, and we report original trace element and Sr-Nd isotope data of new samples collected from the Main Ethiopian Rift and Turkana depression. The data were subdivided according to spatial and temporal criteria. From a spatial point of view, the samples were ascribed to five groups, namely: Afar, Ethiopia, Turkana, Eastern Branch, and Western Branch; from a temporal point of view, the magmatic activity of the EARS was subdivided into three main temporal intervals: 45-25 Ma, 25-10 Ma and 10-0 Ma. The geochemical and radiogenic isotope (Sr, Nd, Pb) signature of the selected basalts denotes the variable contributions of a mantle plume, a more depleted asthenospheric mantle (DMM), and different SubContinental Lithospheric Mantle (SCLM) domains, depending on their temporal and spatial distribution. The geochemistry of the selected basalts shows a marked correspondence with the compositional heterogeneity of mantle xenoliths, whose isotopic systematics (Sm-Nd, Re-Os) indicates the formation of the local SCLM in the Archean and during the Pan-African orogeny. Both SCLM domains contributed significantly to

1  
2  
3 36 magma genesis in the Western Branch (whose signature points towards a contribution of the Pan-  
4 African lithosphere) and Eastern Branch (which is also affected by Archean SCLM domains)  
5 37  
6 38 magmas. The contribution of the SCLM generally increases with time, possibly related to an increase  
7  
8 39 of the geothermal gradient in response to the arrival and flattening of the plume head at the base of  
9  
10 40 the lithosphere and later extension, thinning and shallower melting. Our interpretation supports a  
11  
12 41 pivotal role of the different SCLM domains in magma genesis that is able to fully explain the large  
13  
14 42 compositional heterogeneity of the EARS basalts and represents a reasonable alternative to the  
15  
16 43 putative presence of multiple mantle plumes or a heterogeneous mantle upwelling.  
17  
18  
19  
20

21 46 **Keywords: East African Rift System; rift-related volcanism; upper mantle domains; radiogenic**  
22 47 **isotopes**  
23  
24 48

## 25 49 1. Introduction

27 50 Deformation and thinning of the lithosphere and magma generation with related volcanic activity are  
28  
29 51 typical expressions of extensional forces on continental plates (e.g., [Ebinger, 2020](#)). Extension-related  
30  
31 52 deformation results in the formation of systems of normal faults and subsiding basins; magma  
32  
33 53 generation, migration and uprising may lead to the development of volcanic systems and/or magmatic  
34  
35 54 underplating. Both these major processes may significantly vary in different rift systems or even  
36  
37 55 within a single rift, and during progressive rift evolution.

38 56 The East African Rift System (EARS; Fig. 1) is an ideal place to investigate extensional deformation,  
39  
40 57 thinning of the continental lithosphere, and magma generation, along with their mutual interactions  
41  
42 58 and variations in space and time (e.g., [Furman, 2007](#); [Rooney, 2017, 2020a,b,c,d](#); [Biggs et al., 2021](#);  
43  
44 59 [Morley and Chantraprasert, 2022](#)). The rift is indeed characterised by significant time and spatial  
45  
46 60 variations in deformation style, lithosphere extension and magma volumes and composition (e.g.,  
47  
48 61 [Ebinger, 2020](#)).

48 62 Many previous studies have highlighted the occurrence of different sources in the magmatism of the  
49  
50 63 EARS, including deep mantle plume material, depleted asthenospheric mantle (DMM-like) and the  
51  
52 64 SubContinental Lithospheric Mantle (SCLM). The exact nature of these sources and their variable  
53  
54 65 contribution in the magma genesis is debated and hard to define, and it is beyond the aim of this paper  
55  
56 66 to provide a comprehensive review of the complex magmatic processes that accompanied the  
57  
58 67 development of the EARS, for which the reader is addressed to the recent, comprehensive works by  
59  
60 68 Rooney ([2017, 2020a,b,c,d](#)). Instead, the goal of this work is to analyse the available trace element  
69  
60 69 and isotopic data on volcanic rocks from the EARS to derive a scenario of the time-space variations

1

2

3 70 of major mantle processes associated with rifting such as lithospheric modification and melting.  
4  
5 71 Although rift lavas may display a significant variability at a local scale (see for instance the diversity  
6  
7 72 of volcanic products of the Western Branch in terms of major element composition and isotopic  
8  
9 73 signature; e.g., [Furman et al., 2006](#)) we grouped the available data in different large-scale sectors of  
10  
11 74 the EARS, and long-time intervals, to isolate the common mantle processes controlling rift-related  
12  
13 75 volcanism in relation to the different tectonic settings and stages of rift development. In addition, our  
14  
15 76 approach was integrated with an analysis of available mantle xenoliths from the whole African  
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17 77 domains; the occurrence of mantle xenoliths in many volcanoes of the rift system provides indeed the  
18  
19 78 opportunity to study the geochemical and isotopic characteristics of the different SCLM domains and  
20  
21 79 their potential contribution to the EARS magmatism.

22 80 Based on this large-scale approach, we discuss a geodynamic scenario which accounts for the spatio-  
23  
24 81 temporal variations of the role of some important processes (e.g., plume and plate dynamics) on the  
25  
26 82 generation of magmas in the EARS.

27 83

## 27 84 **2. Geodynamic setting**

29 85 Extension in East Africa and surrounding regions is controlled by interaction among three major  
30  
31 86 plates: Africa, Arabia and Somalia (e.g., [Chorowitz, 2005](#)). The long-lasting motion of the Africa and  
32  
33 87 Somalia, with intervening minor plates south of the Turkana depression, controls extensional  
34  
35 88 deformation in the EARS, whereas extension in the Gulf of Aden, Red Sea and the Afar depression  
36  
37 89 is controlled by the motion of Arabia with respect to the Africa-Somalia system.

38 90 At its northern termination, the EARS is connected to the oceanic domains of the Gulf of Aden and  
39  
40 91 Red Sea by the Afar depression, which hosts the triple junction between the African, Arabian and  
41  
42 92 Somalian plates. In Afar, the crust has been significantly thinned (up to possibly less than ~15 km),  
43  
44 93 the lithosphere is hot and thin and a focused tectono-magmatic activity within axial magmatic  
45  
46 94 segments has been suggested to document incipient oceanic spreading (e.g., [Bastow and Keir, 2011](#)).  
47  
48 95 South of Afar, the EARS is expressed by the occurrence of the Main Ethiopian Rift (MER), where  
49  
50 96 significant axial tectono-magmatic activity in the northern sector testifies ongoing magma intrusion  
51  
52 97 and significant magmatic modification of the lithosphere (e.g., [Keranen et al., 2004](#)). Important  
53  
54 98 magmatic underplating results in a still rather thick crust in the MER (thickness up to >30km).

55 99 The MER terminates in the south, in the Turkana depression, a lowland where faulting and volcanic  
56  
57 100 activity are widespread over an area up to >300 km wide. The anomalous breadth of the deformed  
58  
59 101 region and the characteristics of the tectonic activity in the area have been generally related to a thin  
60  
61 102 crust (20-25 km) resulting from a Mesozoic-Early Cenozoic extension event which gave rise to NW-  
62  
63 103 SE trending basins (e.g., Anza and South Sudan grabens; [Corti et al., 2022](#) and references therein).

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3 104 South of the Turkana depression, the rift bifurcates because of the thick and strong Tanzanian craton,  
4  
5 105 forming the Eastern and Western Branches. The Eastern Branch is composed of the Kenya Rift and  
6  
7 106 Tanzania Divergence, whereas the Western Branch is made of major rift basins such as Albertine,  
8  
9 107 Tanganyika and Malawi. Both rift arms are localised within old mobile belts (such as the  
10  
11 108 Mozambique belt). Similarly to the MER, the Eastern Branch is characterised by significant volcano-  
12  
13 109 tectonic activity, which is localised within axial segments in the northern part of the Kenya rift (e.g.,  
14  
15 110 [Muirhead et al., 2022](#)). Rifting in this EARS branch has likely propagated southwards, with its  
16  
17 111 youngest expression located on the south-eastern side of the thick Tanzania craton, where the  
18  
19 112 deformation zone widens to form the Tanzania Divergence (e.g., [Ebinger et al., 1997](#)).

20  
21 113 The Western Branch is made of a series of long, deep basins formed in Proterozoic-Paleozoic  
22  
23 114 orogenic belts at the western side of the Tanzanian craton. Geophysical data and models indicate that  
24  
25 115 the lithosphere in this region is still cold and strong and values of bulk extension in these basins are  
26  
27 116 smaller than other sectors of the EARS. However, recent works ([Hopper et al., 2020](#)) imaged a  
28  
29 117 significant lithospheric thinning and modification, in the absence of elevated temperatures and  
30  
31 118 magmatism. Indeed, unlike other sectors of the EARS, most of the basins in the Western rift lack any  
32  
33 119 expression of volcanic activity; the four major volcanic provinces (Kivu, Rungwe, Toro-Ankole,  
34  
35 120 Virunga) seem to be localised within regional transfer zones connecting major basins.

36  
37 121 Overall, geophysical studies reveal significant variations in characteristics of the lithosphere in these  
38  
39 122 different domains of the EARS, with thickness varying from >100 km in Proterozoic and Archean  
40  
41 123 lithospheres (up to ~180-250 km in old cratonic cores), to generally less than 100 km within the rifts  
42  
43 124 (e.g., [Fishwick, 2010](#); [Fishwick and Bastow, 2011](#)). Lithospheric thickness is generally higher  
44  
45 125 beneath the Western Branch than in other areas (e.g., [Fishwick and Bastow, 2011](#); [Nijnju et al., 2019](#)),  
46  
47 126 although thinning to 50-60 km has been imaged in the northern Malawi rift ([Hopper et al., 2020](#)).

48  
49 127 Lateral variations in lithospheric strength (i.e., between the old cratons and the surrounding Pan-  
50  
51 128 African mobile belts) strongly controlled the localization of extensional deformation in large parts of  
52  
53 129 the EARS.

### 54 55 131 **2.1. Deformation and basin evolution in the EARS**

56  
57 132 Cenozoic extensional deformation in East Africa is characterised by a complex spatial-temporal  
58  
59 133 evolution (see for instance [Purcell, 2017](#) for a comprehensive review). Following the Mesozoic-Early  
60  
134 Cenozoic phase of graben development, deformation continued in the Turkana depression during the  
135 Eocene-Oligocene; the other portions of the future EARS lack evidence of significant faulting and  
136 rift-related subsidence at that time. Deformation started during Late Oligocene along the western and  
137 southern Afar margins, as a consequence of a westward rift propagation from the eastern parts of the

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3 138 Gulf of Aden caused by counterclockwise rotation of the Arabian plate (Zwaan et al., 2020).  
4  
5 139 Similarly, normal faulting and subsidence continued in the Turkana depression, but significant  
6  
7 140 deformation still lacked in both Eastern and Western Branches. Tectonic activity increased in the  
8  
9 141 Early Miocene (post-22 Ma): deformation in Ethiopia gave rise the initial development of the  
10 142 southern Main Ethiopian Rift (MER), whereas south of the Turkana depression rifting propagated to  
11  
12 143 form portions of the northern Kenya Rift (Purcell, 2017); some deformation could have affected  
13  
14 144 limited portions of the Western Branch (Simon et al., 2017). A major phase of rifting commenced at  
15 145 around 10-15 Ma, with development of different basins in the Western Branch (e.g., Albertine, Kivu,  
16  
17 146 Tanganyika), subsidence and faulting in the northern MER and southward propagation of the Kenya  
18  
19 147 Rift. Significant deformation continued in the Late Miocene-Pliocene up to recent times. Important  
20  
21 148 subsidence affected the narrow rift valleys of the Western Branch; the Eastern Branch further  
22  
23 149 propagated southwards, with the rift impinging the thick lithosphere of the Tanzanian Craton (Masai  
24  
25 150 block) and creating a region of distributed deformation in the so-called Tanzanian Divergence; rifting  
26  
27 151 accelerated in the wide deformation zone of the Turkana depression where the Kenyan and Ethiopian  
28  
29 152 rifts progressively linked; rifting and tectono-magmatic activity increased markedly in the Kenya  
30  
31 153 Rift, in Afar and in the MER, where large boundary faults accommodated deformation, later replaced  
32  
33 154 by axial systems of faulting and volcanism.  
34  
35 155 Overall, there are significant spatial variations in the timing of extension initiation, with the likely  
36  
37 156 heterogeneous, limited activation of individual basins up to after 10-15 Ma, when more diffuse and  
38  
39 157 continuous extension affected the whole EARS in response to motion between the major Nubia-  
40  
41 158 Somalia plates. Rifting has progressed to focused axial tectono-magmatic activity in parts of the rift  
42  
43 159 (northern Ethiopia and northern Kenya), indicating an advanced rifting stage, and almost break-up  
44  
45 160 and incipient spreading in Afar, where the lithosphere is thin and hot. In other parts of the EARS  
46  
47 161 (e.g., Western Branch) rifting is in its initial stages, and occurs in still thick and cold lithosphere.

## 2.2. *Volcanic activity*

48 164 The initial phases of Cenozoic volcanism in East Africa correspond to the eruption of flood basalts  
49  
50 165 in the Horn of Africa (see review in Rooney, 2017). This large-scale volcanism started in southern  
51  
52 166 Ethiopia and northern Kenya at ~45 Ma until ~34 Ma, followed by a second pulse of flood basalts  
53  
54 167 during the Oligocene, with eruption of thick lava flows characterising the plateaus of northern  
55  
56 168 Ethiopia, Somalia and Yemen (Rooney, 2017 and references therein). This phase of flood basalt  
57  
58 169 activity occurred between ~34 Ma and ~27 Ma, with the majority of volcanic products emplaced in a  
59  
60 170 short time span at ~30 Ma (Hoffman et al, 1997). Another, significantly less important phase of  
61  
62 171 basaltic volcanism occurred between ~27 Ma and ~22 Ma throughout the northern EARS (Early

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3 172 Miocene Resurgence Phase of [Rooney, 2020a](#)), and this activity produced shield-volcanoes until  
4  
5 173 about 10 Ma in the plateaus of Ethiopia ([Kieffer et al., 2004](#)). Limited eruption of carbonatitic  
6  
7 174 magmas in the Western rift occurred in this period (~25-26 Ma) close to Lake Rukwa area ([Roberts  
8  
9 175 et al., 2012](#)). The resurgence phase in the EARS may have continued in an additional phase of  
10 176 widespread basaltic activity which took place between 20 and 16 Ma in the Northern Kenya Rift,  
11  
12 177 southern Ethiopia, and parts of the Turkana depression ([Rooney, 2020a](#)). This phase, dominated by  
13  
14 178 basaltic volcanism was followed by a period of widespread evolved volcanism consisting of  
15 179 phonolites in northern Kenya, and explosive silicic eruptions in the Main Ethiopian Rift. Alkaline  
16  
17 180 volcanism initiated in the Western Branch, specifically in the Kivu-Virunga and Rukwa volcanic  
18  
19 181 provinces.

20 182 At ~12 Ma a renewed phase of widespread basaltic volcanism (Mid-Miocene Resurgence Phase of  
21  
22 183 [Rooney, 2020a](#)) characterised East Africa from Afar to Kenya, with activity in the south slightly  
23  
24 184 predating equivalent basaltic volcanism in the north. This phase of dominantly basaltic activity lasted  
25  
26 185 until ~8-9 Ma and was followed by widespread bimodal volcanism (basalts, trachytes and large-scale  
27 186 silicic volcanism) concomitant with a major rifting episode characterising the subsiding rift valleys  
28  
29 187 of Ethiopia and Kenya, and by focused volcanic activity in the different volcanic provinces of the  
30  
31 188 Western Branch.

32 189 Another major pulse of basaltic activity commenced at about ~4 – 5 Ma in the Afar depression, with  
33  
34 190 the eruption of the so-called Afar Stratoid Series, which affected most of the depression until the  
35  
36 191 Pleistocene. Widespread late Miocene-Pliocene basaltic lavas also interested the Turkana depression  
37  
38 192 ([Franceschini et al., 2020](#); [Rooney, 2020a](#)). Recent volcanic activity is mostly bimodal in the  
39 193 Ethiopian and Kenyan rifts, with axial silicic central volcanoes, in most cases with caldera-forming  
40  
41 194 eruptions, and alignments of basaltic cones and fissures. Basaltic activity is predominant in the  
42  
43 195 northern Afar depression, where incipient continental break-up is observed at axial tectono-magmatic  
44  
45 196 segments. Magma intrusion in the lithosphere and the related thermo-mechanical effects may have  
46 197 favoured magma-assisted rifting in these regions (e.g., [Kendall et al., 2005](#)) and weakened the  
47  
48 198 lithosphere to allow rifting in strong cratonic lithosphere of the southern parts of the EARS (such as  
49  
50 199 the Tanzanian Divergence; e.g., [Ebinger, 2020](#)).

51 200 Overall, the strong difference in the volumes of magmatic products between the highly volcanic  
52  
53 201 Eastern Branch and the almost non-volcanic Western Branch has been interpreted to reflect an  
54  
55 202 asymmetric upraising and emplacement of upwelling mantle material ([Koptev et al., 2015](#)). This  
56  
57 203 process has been influenced by the initial geometry of the lithosphere due to the presence of the thick  
58 204 Tanzanian craton, which diverted the upraising mantle towards the Eastern Branch where warm  
59  
60 205 material accumulated at the base of the thinned lithosphere, increasing decompression melting



1

2

3 206 ([Koptev et al., 2015](#)). A general southward younging of the main volcanic phases (see [Morley and](#)  
4 [Chantraprasert, 2022](#)) is likely related to mantle plume dynamics (see below section 2.3).

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### 8 209 **2.3. Mantle plume(s) influence on rifting in the EARS**

10 210 Rifting in East Africa has been attributed to the activity of one or more mantle plumes based on  
11 geophysical, geodynamical and geochemical evidences, as outlined hereafter.

12 211  
13 212

#### 14 213 *2.3.1. Geophysical and geodynamical evidences for mantle plume activity*

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16 215

17 215 With the exception of the Turkana depression, the rift valleys of East Africa cut through broad  
18 elevated plateaus; this, together with the initial widespread flood basalt emplacement before  
19 216 significant rifting, has been related to mantle upwelling processes (e.g., [Moucha and Forte, 2011](#);  
20 217 [Hassan et al., 2020](#)). The existence of one or several deep mantle plumes under the EARS is strongly  
21 supported by geophysical investigations, which have highlighted since a long time the occurrence of  
22 218 a large region of low velocity seismic waves rising from the core-mantle boundary, the so-called  
23 African superplume (e.g., [Ritsema et al., 1999](#)). However, the structure of the shallow mantle is not  
24 219 clearly defined by geophysical studies and the number and features of upwelling mantle domains  
25 220 remain debated (e.g., [Boyce et al., 2021](#); [Civiero et al. 2022](#) and references therein).

26 221 Irrespective of the number and structure of mantle plumes, upwelling of hot mantle material is  
27 222 believed to control rifting in the area. Africa and Somalia are indeed surrounded by oceanic ridges,  
28 and the plates are subjected to compression driven by ridge-push force, with no clear regional plate  
29 223 configuration to drive rifting ([Coblentz and Sandiford, 1994](#)). Rifting is instead related to extensional  
30 224 forces imposed by a combination of gravitational potential energy gradients related to mantle-driven  
31 225 uplift and basal drag from horizontal mantle flow at the base of the lithosphere (e.g., [Stamps et al.,](#)  
32 226 [2014, 2015](#)). The age-progressive volcanism in the EARS, characterised by an overall southward  
33 227 younging, has been interpreted as resulting from a southward migrating (Afar) plume relative to the  
34 228 African plate (e.g., [Hassan et al., 2020](#); [Morley and Chantraprasert, 2022](#)).

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36 230

#### 37 231 *2.3.2. Geochemical signature of plume material*

38 232  
39 233

40 234 A robust geochemical evidence for a deep mantle plume in the EARS derives from the occurrence of  
41 235 magmas with He isotopes  $>8 R_A$  (where  $R_A$  represent the He isotope composition of atmosphere, e.g.,  
42 236 [Halldórsson et al 2014](#); [Castillo et al., 2020](#)), which requires the predominant contribution of a deep  
43 237 undegassed mantle (e.g. [Graham et al., 1992](#); [1998](#); [Hanan and Graham, 1996](#)). Additional evidence  
44 238 of anomalously hot rising mantle material is provided by the potential temperature recorded by  
45 239 olivine, which are in excess of  $>150$  °C relative to the ambient upper mantle (e.g., [Wong et al., 2022](#)).

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1  
2  
3 242 Many studies have tried to infer the geochemical and isotopic signature of this plume material from  
4  
5 243 the magmatism of the northern EARS (e.g., [Pik et al., 1999](#); [Rogers et al., 2000](#); [George and Rogers,](#)  
6  
7 244 [2002](#); [Halldórsson et al 2014, 2022](#); [Nelson et al 2012](#); [Rooney et al., 2012](#); [Furman et al., 2016](#);  
8  
9 245 [Castillo et al., 2020](#)). The recent mafic magmas erupted at Erta Ale and Djibouti have been identified  
10 246 as representative of its composition or at least of its tail (e.g., [Furman et al., 2016](#)). This small range  
11  
12 247 of values overlap with the postulated isotopic composition of the plume extrapolated by [Rooney et](#)  
13  
14 248 [al. \(2012\)](#) from MER magmas ( $^{87}\text{Sr}/^{86}\text{Sr}\sim 0.7035$  and  $^{143}\text{Nd}/^{144}\text{Nd}\sim 0.5129$ ), similar to the composition  
15 249 of recent magmas of the Afar region ([Castillo et al., 2020](#)). Early flood basalts (HT2 basalts) of  
16  
17 250 Ethiopia have been also considered a *proxy* for the composition of plume material ([Kieffer et al.,](#)  
18  
19 251 [2004](#); [Muravyeva and Senin 2018](#)) even though a contribution of the SCLM was proposed ([Furman](#)  
20  
21 252 [et al., 2016](#); [Natali et al., 2016](#); [Rooney, 2017](#); [Nelson et al., 2019](#)). In comparison with the previous  
22 253 compositions, the HT2 flood basalts have slightly more radiogenic  $^{87}\text{Sr}/^{87}\text{Sr}$  ( $\sim 0.7040$ ) but similar  
23  
24 254  $^{143}\text{Nd}/^{144}\text{Nd}$ . In terms of Pb isotopes, the isotopic composition of the plume estimated by [Rooney et](#)  
25  
26 255 [al. \(2012\)](#) and [Castillo et al. \(2020\)](#) are similar, with  $^{206}\text{Pb}/^{204}\text{Pb}\sim 19.4\text{-}19.5$  and  $^{208}\text{Pb}/^{204}\text{Pb}\sim 39.2$ ,  
27 256 while the HT2 flood basalts have less radiogenic Pb isotopic composition ([Natali et al., 2016](#), [Nelson](#)  
28  
29 257 [et al., 2019](#)). Finally, magmas from Kenya and Turkana have more radiogenic Pb isotopes and have  
30  
31 258 been originally interpreted to reflect an additional HIMU component within the mantle plume  
32  
33 259 ([Furman et al., 2006](#)) or a second mantle plume ([George et al., 1998](#); [Rogers et al., 2000](#)), even though  
34 260 their Pb isotope signature was later considered to derive from the SCLM via drip melting ([Furman et](#)  
35  
36 261 [al., 2016](#)).

37  
38 262 Regarding the models accounting for the different age and geochemical heterogeneities of the early  
39  
40 263 magmatism of the northern EARS, many interpretations have been proposed: from the existence of  
41 264 two different plumes (Afar and Kenya plumes, e.g., [George et al., 1998](#); [Rogers et al., 2000](#)) to a  
42  
43 265 single, chemically heterogeneous upwelling ([Furman et al., 2006](#)), or a hybrid model involving two  
44  
45 266 branches of the same plume with different compositions ([Nelson et al., 2012](#)). According to the recent  
46 267 review by [Rooney \(2020d\)](#), the prevalent contribution of the deep mantle material beneath the EARS  
47  
48 268 can be considered equivalent to a single plume, i.e. the Afar plume. Hereafter, we refer to the Afar  
49  
50 269 plume to indicate the anomalous hot mantle beneath the EARS.

### 51 270

### 53 271 3. Sample selection and grouping

54  
55 272 The full available database of the EARS magmas was downloaded from GEOROC (<https://georoc.eu/>) on  
56  
57 273 15<sup>th</sup> March 2021, selecting only the less evolved lavas ( $\text{SiO}_2 < 52$  wt%,  $\text{MgO} > 4$  wt%) to minimize the  
58 274 effects of crystal fractionation and crustal assimilation. The published data, consisting of more than  
59  
60 275 1500 samples, were implemented with new unpublished data from Ethiopia (25 samples) and the

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2  
3 276 Turkana depression (2 samples; **Figure 1**; **Supplementary Table1**). These new samples have been  
4  
5 277 analysed for major, trace and Sr-Nd isotopes as detailed in the **Supplementary Material**.

6  
7 278 All samples were then subdivided on the basis of temporal and spatial criteria based on the main  
8  
9 279 tectonic and geodynamic evolution constraints along with the timing of magmatism as described  
10 280 above. Three main temporal periods were defined as follow (**Figure 1**):

- 11  
12 281 i) 45-25 Ma, a time interval that defines the (pre-rift) flood basalt event;
- 13  
14 282 ii) 25-10 Ma, an intermediate period characterised by time-space irregular initiation of  
15 283 deformation and heterogeneous volcanism (from continuation of flood basalt activity to more  
16  
17 284 acid volcanism) among the Arabia, Nubian and Somalian plates;
- 18  
19 285 iii) 10-0 Ma, the youngest interval related to the main rifting phases, with diffuse tectonic and  
20  
21 286 volcanic activity.

22 287 From a spatial point of view, the samples were further subdivided into five different groups following  
23  
24 288 the main tectonic domains of the EARS, roughly from N to S (**Figure 1**):

- 25  
26 289 i) the Afar depression, which is characterised by incipient oceanic spreading and records the  
27 290 interaction;
- 28  
29 291 ii) Ethiopia, which includes the Ethiopian rift valley and surrounding plateaux, characterised by  
30  
31 292 significant tectono-magmatic activity in a lithosphere modified by magmatic processes;
- 32  
33 293 iii) the Turkana depression, characterised by widespread tectonic and volcanic activity in a  
34 294 region of thinned crust following previous (Mesozoic) tectonic events;
- 35  
36 295 iv) the Eastern Branch, where an intense volcanic activity is generated due to warm material  
37  
38 296 accumulated at the base of a thinning lithosphere;
- 39  
40 297 v) the Western Branch, which has a limited extension and forms narrow, deep basins and  
41 298 localised volcanism in a region of cold and strong lithosphere.

42  
43 299  
44  
45 300 In order to correctly understand our final interpretation of the analysed data, it is important to bear in  
46 301 mind that the distribution of samples of the selected database does not actually represent the  
47  
48 302 distribution of the volcanic products in nature, on both volumetric and temporal basis. As an example,  
49  
50 303 the samples from the recent period (10-0 Ma) in the Western Branch are as abundant as the ones from  
51  
52 304 the Eastern Branch, whereas the volumes of erupted magmas from the latter area are much larger than  
53 305 those from the former region. Also, not all different magmas have been retrieved for their isotopic  
54  
55 306 composition, limiting the completeness and the representativeness of our database. For example, no  
56  
57 307 isotopic analysis on samples from the Western Branch older 10 Ma are available, despite limited  
58 308 volcanism has occurred in the area since ~25-26 Ma (**Figure 1**). Nevertheless, our database may be  
59  
60 309 considered well representative for our purpose: from a spatial point of view, all the selected samples

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2  
3 310 are evenly distributed within the main volcanic regions. The temporal distribution of samples is in  
4  
5 311 line with the overall southward younging of volcanism described above: consequently, the oldest  
6  
7 312 interval is sampled in the northern areas only (Afar, Ethiopia, Turkana), whereas progressively  
8  
9 313 younger volcanic rocks are sampled moving southwards (Eastern Branch and Western Branch). In  
10 314 detail, most of the selected samples are from the recent (10-0 Ma) volcanic activity (Afar 19%,  
11  
12 315 Ethiopia 13%, Turkana 2%, Eastern Branch 23%, Western Branch 24%) although magmas from the  
13  
14 316 other two periods of activity in the different areas are represented as well (19% in total) with the  
15 317 exception of the Eastern Branch and Western Branch having no samples with an emplacement age  
16  
17 318 >25 Ma and >10 Ma, respectively.

#### 20 320 4. Geochemical characteristics of mafic magmas

22 321 The chemical composition of mafic magmas varies from basalt, picobasalt, trachybasalt, basanite  
23  
24 322 and foidite with the Eastern and Western Branch samples having the most undersaturated  
25  
26 323 compositions. All the selected samples, but the magmas from the Eastern and Western Branches, have  
27  
28 324 a relatively constant  $\text{CaO}/\text{Al}_2\text{O}_3$  with decreasing MgO (Figure 2) in agreement with the compositional  
29 325 variation of mantle-derived magmas from oceanic areas (MORB and OIB; GEOROC database). This  
30  
31 326 signature provides evidence for the reliability of the screened samples in terms of minimising  
32  
33 327 geochemical variations due to low-pressure fractionation processes. Figure 2 shows that olivine is the  
34 328 main mineral responsible for the observed decrease in MgO. Admittedly, at  $\text{MgO} < 7\%$  a subset of  
35  
36 329 samples have a combined decrease in MgO and  $\text{CaO}/\text{Al}_2\text{O}_3$ , suggesting an additional minor role of  
37  
38 330 clinopyroxene fractionation. The effect of clinopyroxene fractionation could be removed by limiting  
39  
40 331 the database to samples with  $\text{MgO} > 7\%$  but, since we will mostly discuss variations in incompatible  
41 332 trace elements that are negligibly affected by clinopyroxene and olivine, we decided not to restrict  
42  
43 333 the database. The magmas from the Eastern and Western Branches have higher  $\text{CaO}/\text{Al}_2\text{O}_3$  than other  
44  
45 334 EARS magmas (Figure 2), demanding for an origin from different mantle source domains, which will  
46 335 be considered in the following sections.

48 336 Another useful comparison between our EARS samples and MORBs and OIBs is reported in the  
49  
50 337 Th/Yb vs. Ta/Yb diagram (Figure 3). Most the mafic magmas of the older two periods of activity are  
51  
52 338 superimposed to OIB compositions and denote a provenance from fertile mantle sources. Half of the  
53 339 samples from Ethiopia have, however, lower Ta/Yb and Th/Yb indicating either an origin from more  
54  
55 340 depleted mantle sources (Pik et al., 1999) or higher mantle melting degrees, as suggested by their  
56  
57 341 high MgO contents ( $\text{MgO} > 10$  wt.%, Figure 2a). The latter hypothesis comes from the observation  
58 342 that the degree of incompatibility during mantle melting increases from Th to Ta and Yb (e.g., Sun  
59  
60 343 and McDonough, 1989; McKenzie and O'Nions, 1991; Kelemen et al., 2003), meaning that, starting

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2  
3 344 from the same mantle source, the higher the melting degree, the lower Th/Yb and Ta/Yb in produced  
4  
5 345 melts. Most mafic magmas of the youngest period of activity (Figure 3b) cluster at similar Th/Yb and  
6  
7 346 Ta/Yb as the older mafic magmas, except samples from the Eastern and Western Branches, which  
8  
9 347 show higher Ta/Yb and Th/Yb. This characteristic requires the contribution of other mantle source  
10 348 domains as also attested by their high CaO/Al<sub>2</sub>O<sub>3</sub> (Figure 2b). The Afar samples with low Th/Yb and  
11  
12 349 Ta/Yb (Figure 3a) recall a depleted mantle sources as suggested for the Ethiopia mafic magmas of  
13  
14 350 the oldest period of activity (Barrat et al., 2003; Daoud et al., 2010).

15 351  
16

#### 17 352 **4.1. Partial melting of the mantle source**

18  
19 353 Constraints on the depth and degrees of mantle melting originating the EARS mafic magmas can be  
20  
21 354 assessed by Rare Earth Element (REE) fractionation. One of the most useful diagrams is Yb vs.  
22 355 La/Yb, which is strongly dependent upon mantle melting in the garnet and spinel stability fields. We  
23  
24 356 have modelled the REE signature of the EARS mafic magmas (Figure 4) applying a non-modal batch  
25  
26 357 melting to a nominal Primitive Mantle (McDonough and Sun, 1995) and SCLM (McDonough, 1990)  
27 358 using partition coefficients from the compilation of McKenzie and O’Nions, (1991) and Kelemen et  
28  
29 359 al. (2003). Both sources are not meant to be the actual mantle domains of the overall EARS mafic  
30  
31 360 magmas; rather, they are used as *proxies* to have a qualitative information on melting degrees and  
32  
33 361 depths (e.g. Ayalew et al., 2018; Feyssa et al., 2019; Tortelli et al., 2022).

34 362 The REE fractionation of mafic magmas belonging to the oldest two periods of activity (Figure 4a)  
35  
36 363 are consistent with variable melting degrees of a primitive mantle source both in the garnet and spinel  
37  
38 364 stability fields, with a few samples from Ethiopia, Turkana, and Eastern branch indicating lower  
39 365 degrees of mantle melting (<1%) predominantly in the garnet stability field. In contrast, the EARS  
40  
41 366 mafic magmas of the youngest period of activity reveal a significant dichotomy. The Afar, Turkana,  
42  
43 367 and Ethiopia samples suggest an origin from variable melting degrees of a primitive mantle source  
44  
45 368 almost entirely in the spinel stability field (Figure 4b), whilst the Western Branch and, to a lesser  
46 369 extent, the Eastern Branch samples have extreme La/Yb values (up to 320), which are not consistent  
47  
48 370 with melting of a primitive mantle-like garnet lherzolite (Figure 4b) even at very low melting degrees.  
49  
50 371 Their REE fractionation fits better with an origin from a SCLM source in the garnet and spinel  
51  
52 372 stability fields at low melting degrees (Figure 4c).

53 373  
54

#### 55 374 **4.2. Mantle source mineralogy**

56  
57 375 Incompatible trace element ratios can provide constraints on minor mineral phases occurring in the  
58 376 mantle source in addition to a *normal* lherzolite (olivine + orthopyroxene + clinopyroxene ± spinel ±  
59  
60 377 garnet). Ba/Rb vs. Rb/Sr yields information on the role of K-bearing phases in the mantle source (e.g.,

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2  
3 378 Furman and Graham, 1999). Deviations from primitive mantle and SCLM estimates indicate  
4  
5 379 metasomatic enrichment events: mantle-derived magmas in equilibrium with phlogopite-bearing  
6  
7 380 lherzolite are expected to have significantly higher Rb/Sr and lower Ba/Rb than those in equilibrium  
8  
9 381 with amphibole-bearing lherzolite (Figure 5). The EARS mafic magmas of the older two periods of  
10 382 activity cluster from primitive mantle composition towards high Ba/Rb in the case of Ethiopia and  
11  
12 383 Turkana mafic magmas (Figure 5a), suggesting an origin from a lherzolite variably metasomatised  
13  
14 384 by amphibole. In contrast, a few samples from Ethiopia, and the Western and Eastern Branches have  
15 385 high Rb/Sr pointing to a phlogopite-bearing mantle source. Mafic magmas of the youngest period of  
16  
17 386 activity show a more pronounced affinity for a lherzolite variably metasomatised by amphibole and/or  
18  
19 387 phlogopite (Figure 5b). Phlogopite-bearing lherzolite melts are prevalent, albeit not exclusively, in  
20  
21 388 the Western and Eastern Branches mafic magmas (Foley et al., 2012, Rosenthal et al., 2009), whilst  
22 389 amphibole-bearing lherzolite melts are more abundant in the Eastern Branch, Turkana, Ethiopia, and  
23  
24 390 Afar mafic magmas. The presence of phlogopite in the mantle beneath the EARS is supported by the  
25  
26 391 occurrence of phlogopite-bearing mantle xenoliths, which are common in the northern sector of the  
27 392 Western Branch (Foley et al., 2012, Lloyd et al., 2002). In the Western Branch, the presence of  
28  
29 393 phlogopite and mantle melting at depth > 100 km is also attested by the occurrence of kamafugites  
30  
31 394 (and possibly kimberlites) (Foley et al., 2012, Tappe et al., 2020).

32  
33 395 In addition to phlogopite and amphibole, the trace element signature of EARS mafic magmas reveals  
34 396 another minor component likely occurring in their mantle source. This is displayed by Zr/Nb vs.  
35  
36 397 Nb/Ta (Figure 6). Zr/Nb is fractionated during mantle melting and represents a proxy of source  
37  
38 398 fertility, with MORBs having higher Zr/Nb than OIBs because the former originate from more  
39 399 depleted mantle domains. In contrast, Nb and Ta have similar partition coefficients during mantle  
40  
41 400 melting, as attested by the rather constant Nb/Ta in MORBs and OIBs (Figure 6). The EARS mafic  
42  
43 401 magmas of the older two periods of activity suggest an origin from a variable depleted/enriched  
44  
45 402 mantle source ( $2 < \text{Zr/Nb} < 40$ ), similar to MORBs and OIBs, and without any significant Nb/Ta  
46 403 fractionation (Figure 6a). Admittedly, most samples are from the Ethiopia albeit a few samples are  
47  
48 404 also from Turkana and Eastern Branch. As with other trace element ratios, the EARS mafic magmas  
49  
50 405 of the youngest period of activity display differences with respect to the oldest mafic magmas (Figure  
51 406 6b). While most of Afar, Ethiopia, and Turkana mafic magmas plots within the range defined by the  
52  
53 407 oldest lavas and oceanic area basalts, some samples from the Eastern and Western Branches are  
54  
55 408 displaced toward higher Nb/Ta (up to ca. 50), with the latter shifting also towards low Zr/Nb values  
56  
57 409 (down to <1). This peculiar signature suggests carbonatitic metasomatism of the mantle source  
58  
59 410 (Green, 1995; Pfänder et al., 2012; Bragagni et al., 2022) and can also explain their high  $\text{CaO/Al}_2\text{O}_3$   
60 411 (Figure 2 and Foley et al., 2012). Carbonatites are indeed widespread in these two regions (e.g., van

1  
2  
3 412 Straaten and Bell, 1989; Bell and Blenkinsop, 1987; Foley et al., 2012; Guzmics et al., 2019) and  
4  
5 413 possess the unique Nb/Ta and Zr/Nb signature (Hoernle et al., 2002; Bizimis et al., 2003;  
6  
7 414 Chakhmouradian, 2006) required to explain the offset of the Eastern and Western Branch magmas  
8  
9 415 from typical values of oceanic (MORBs, OIBs) and continental (e.g., other EARS) mafic magmas.  
10 416 Overall, the trace element characteristics of the EARS mafic magmas are roughly consistent with a  
11  
12 417 progressive increase of the contribution of metasomatic mantle sources, especially in the southern  
13  
14 418 part of the rift and in the youngest volcanic activity.

### 15 419 16 17 420 **4.3. Radiogenic isotopes**

18  
19 421 Radiogenic isotope compositions of EARS mafic magmas are reported in Figure 7 along with fields  
20  
21 422 of MORBs and OIBs. The Afar and Ethiopia mafic magmas of the older two periods of volcanic  
22  
23 423 activity (Figure 7a) cluster around the composition that is considered to represent the Afar mantle  
24  
25 424 plume (Rogers et al., 2000; George and Rogers, 2002; Rooney et al., 2012; Castillo et al., 2020) with  
26  
27 425 trends towards the MORB, EM I, and EM II mantle components of the mantle zoo (e.g., Hofmann,  
28  
29 426 1997; Stracke and Hofmann, 2005). The Turkana and a subset of the Ethiopia samples are displaced  
30  
31 427 towards a HIMU component suggesting, following some interpretations, the occurrence of a different  
32  
33 428 plume (i.e., the so-called Kenya plume, Rogers et al., 2000). The Eastern Branch samples are  
34  
35 429 displaced from the Afar and the putative Kenya plume to less radiogenic Nd isotope compositions  
36  
37 430 and also display a trend towards radiogenic Sr isotope compositions ( $^{87}\text{Sr}/^{86}\text{Sr}$  up to 0.706), which is  
38  
39 431 offset from the EM II mantle component (Figure 7a).

40  
41 432 In contrast, the EARS mafic magmas of the youngest period of activity have a widespread range of  
42  
43 433 isotopic compositions starting from the Afar mantle plume to the MORB-like mantle and to extremely  
44  
45 434 radiogenic Sr isotope and unradiogenic Nd isotope compositions, far exceeding the EM II and EM I  
46  
47 435 mantle components (Hofmann, 1997), respectively (Figure 7b). The Afar mantle plume component  
48  
49 436 is present, with no exception, in all of the mafic magmas of this period, whilst the radiogenic Sr  
50  
51 437 isotope mantle component is predominant in the Western Branch mafic magmas, with a few samples  
52  
53 438 from the Eastern Branch, in agreement with the contribution of a phlogopite-bearing lherzolite  
54  
55 439 (Figure 5). The unradiogenic Nd isotope mantle component is restricted to the Eastern Branch mafic  
56  
57 440 magmas and is even more extreme than that of the previous periods of activity (Figure 7), requiring  
58  
59 441 a low time-integrated Sm/Nd.

60 442 Lead isotope compositions of EARS mafic magmas have less, albeit significant, differences between  
61  
62 443 the oldest and youngest periods of volcanic activity (Figure 8). As with Sr and Nd isotopes (Figure  
63  
64 444 7), the samples, starting from the common Afar plume component, are offset towards MORB, EM I,  
65  
66 445 EM II, and HIMU mantle components. Most of mafic magmas are aligned along the NHRL defined

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2  
3 446 by MORBs and OIBs, with Afar, Ethiopia, and Turkana mafic magmas having generally less  
4  
5 447 radiogenic Pb isotope compositions than those of the Western and Eastern Branches (Figure 8). In  
6  
7 448 particular, the recent Western Branch mafic magmas deviates from the other EARS magmas, pointing  
8  
9 449 towards more radiogenic  $^{207}\text{Pb}/^{204}\text{Pb}$  at the same  $^{206}\text{Pb}/^{204}\text{Pb}$  (Figure 8b). Such a radiogenic  
10 450  $^{207}\text{Pb}/^{204}\text{Pb}$  signature exceeds the range observed in magmas with EM II affinity, requiring the  
11  
12 451 contribution of an old crustal component with high time-integrated U/Pb. Instead, the Eastern Branch  
13  
14 452 mafic magmas of the three periods of activity are offset towards HIMU-like compositions (Figure 8).  
15 453 Such a shift to a HIMU mantle component is odd with their extreme unradiogenic Nd isotope  
16  
17 454 composition (Figure 7b). Another peculiar characteristic is that the Turkana and the subset of the  
18  
19 455 Ethiopia samples, although suggesting the occurrence of a HIMU mantle component in the Sr-Nd  
20  
21 456 isotope diagram (Figure 7a), do not have Pb isotope compositions recalling the HIMU mantle  
22 457 component (Figure 8).

23  
24 458 As with trace element ratios (Figures 3-6), Sr-Nd-Pb isotope compositions of the EARS mafic  
25  
26 459 magmas reveal significant spatial and temporal variations and allow to place important constraints on  
27 460 the contribution of different mantle domains in magma genesis during the evolution of the rift. In  
28  
29 461 particular, since there is geochemical evidence for the influence of the SCLM in the magmatism of  
30  
31 462 the EARS (e.g. Furman et al., 2016), in addition to the Afar mantle plume, we will focus on the role  
32  
33 463 of the SCLM in affecting the isotopic signature of magmas in space and time. The geochemical and  
34 464 isotopic compositions of mantle xenoliths from different lithospheric African domains provides the  
35  
36 465 opportunity to assess their role in the EARS magmatism.

## 37 38 466

### 39 467 **5. The message from the radiogenic isotope signature of mantle xenoliths**

40  
41 468 In order to investigate the contribution of the SCLM on magma genesis, we collected available  
42  
43 469 radiogenic isotope data (Sr-Nd-Pb-Os) on ultramafic mantle xenoliths (GEOROC database) that are  
44  
45 470 representative of SCLM fragments sampled by ascending magmas of the EARS and the surrounding  
46 471 regions (Arabia peninsula and South Africa). These data are compared with the observed trends in  
47  
48 472 the EARS mafic magmas to constrain their genesis in space and time at the scale of the whole rift  
49  
50 473 system.

#### 51 474

#### 52 475 **5.1. *Sm-Nd isotope composition***

53 476 The Sm-Nd isotope systematics can be successfully used to retrieve age information on ultramafic  
54  
55 477 garnet-bearing rocks. In the Sm-Nd isochron diagram (Figure 9), the ultramafic xenoliths are roughly  
56  
57 478 distributed along two different trends, revealing two distinct formation ages of the African SCLM.  
58 479 The reference isochrons reported in Figure 9 were calculated starting from the Nd isotope



1  
2  
3 480 composition of either the primitive (dashed lines) or the depleted mantle (solid lines). The most recent  
4  
5 481 reference isochrons comprise most of the mantle xenoliths and are consistent with a relatively young  
6  
7 482 event at ~600 Ma. This age, recorded by fragments of the SCLM, includes samples from the African  
8  
9 483 continent (South Africa, Tanzania, Ethiopia) and Arabia (Figure 9). The other reference isochrons  
10 484 encompass samples from the Tanzanian craton and record an older event at ~2.6 Ga. These two pairs  
11  
12 485 of reference isochrons are consistent with the two main formation events of the SCLM during the  
13  
14 486 Archean and the Pan-African (Chesley et al., 1999, Burton et al., 2000).

15 487 Finally, carbonatites occurring in the southern portion of the EARS (both Western and Eastern  
16  
17 488 Branches) are incidentally broadly aligned along both the Pan-African and Archean reference  
18  
19 489 isochrons (Figure 9) suggesting an origin from a metasomatised SCLM (Foley et al., 2012).

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21 490

## 22 491 5.2. Os isotope composition

23  
24 492 The Re-Os isotope systematic provide further age constrains on events recorded by mantle xenoliths.  
25  
26 493 Since Re is incompatible and Os is compatible during mantle melting, the Os isotope signature of the  
27  
28 494 mantle is “frozen” after melt extraction. As such, the so-called Re-depletion model ages ( $T_{RD}$ )  
29 495 represent a tool to date the formation of the lithospheric mantle (Walker et al. 1989). Importantly,  
30  
31 496 metasomatic events in the SCLM can overprint the original Re-Os isotopic signature, making  $T_{RD}$   
32  
33 497 ages only reliable “minimum ages”. If metasomatism acts in an unsystematic way, information can  
34  
35 498 be still obtained when considering the distribution of  $T_{RD}$  ages. A survey of  $T_{RD}$  Os ages from mantle  
36 499 xenoliths of the EARS shows three main peaks at ca. 2.8, 1.5, and 0.5 Ga (Figure 10). The  $T_{RD}$  ages  
37  
38 500 are all calculated using the primitive mantle estimates of Meisel et al. (2001) and might slightly differ  
39  
40 501 from those reported in the original publications, where different reference reservoirs had been used.  
41 502 Despite the limited number of data, the three  $T_{RD}$  age peaks are also observed in mantle xenoliths  
42  
43 503 from individual volcanic centres.

44  
45 504 In north Ethiopia, near the Afar depression, most  $T_{RD}$  ages are clustered around 0.6 Ga. Even if these  
46 505 model ages were originally interpreted as “disturbed” with no age meaning (Alemayehu et al., 2019),  
47  
48 506 it is striking that they define a relative narrow range, matching the Pan-African event. Although  
49  
50 507 limited to only two samples, older  $T_{RD}$  ages of 1.6 and 1.2 Ga, points to portions of the lithosphere  
51  
52 508 that was Archean in age and later (partially) overprinted (Alemayehu et al., 2019). Mantle xenoliths  
53 509 from the northern Turkana depression show a rough bimodal  $T_{RD}$  age distribution with a first peak  
54  
55 510 around 500 Ma and an older asymmetric peak at 1.5 Ga, which includes  $T_{RD}$  ages up to 2.4 Ga. The  
56  
57 511 oldest  $T_{RD}$  ages of the EARS (up to 3.4 Ga) are recorded in mantle xenoliths from the Eastern Branch  
58 512 (Burton et al., 2000; Chesley et al., 1999; Meisel et al., 2001) with a clear Archean signature, resulting  
59  
60 513 in a main peak at around 2.6 Ga. This ancient signature is consistent with the formation of the

1  
2  
3 514 Tanzanian craton at 2.8-3.4 Ga (Chesley et al., 1999; Burton et al., 2000). Unfortunately, no  $T_{RD}$  ages  
4  
5 515 are available for mantle xenoliths from the Western Branch of the EARS.

6 516 The three  $T_{RD}$  age peaks observed in the complete dataset of the EARS, reflect also variable  
7  
8 517 contribution of different lithospheric domains. In particular, Archean  $T_{RD}$  ages are recorded only  
9  
10 518 around the Tanzania Craton, while Pan-African ages are mostly recorded in the north, near the Afar  
11  
12 519 region. These ages are consistent with what observed with Sm-Nd isotopes (Figure 9), stressing the  
13  
14 520 importance of two lithospheric mantle domains, affected by (at least) two distinct events. The  
15 521 additional peak at 1.5 Ga, recorded by the samples from northern Turkana depression in the  $T_{RD}$   
16  
17 522 density probability plot (Figure 10) could reflect a rejuvenation of the Archean lithosphere (i.e., no  
18  
19 523 age meaning) or express an additional event in the lithospheric mantle. Following the second  
20  
21 524 possibility, a 1.5 Ga geological event was proposed by Alemayehu et al. (2019) on the basis of a  
22 525 linear array of xenoliths in a Re-Os isochron diagram. The regression through these data yielded a  
23  
24 526 1.5 Ga age, synchronous with the break-up of the Nuna supercontinent (Alemayehu et al., 2019).  
25  
26 527 Although the error on the inferred initial Os isotopes is large ( $^{187}\text{Os}/^{188}\text{Os} = 0.117 \pm 0.003$ , Alemayehu  
27 528 et al., 2019), it is consistent with the composition of the primitive mantle at that time ( $^{187}\text{Os}/^{188}\text{Os} =$   
28  
29 529 0.118), possibly reflecting the formation of specific domains of the lithospheric mantle at that time  
30  
31 530 as also suggested by the peak in the  $T_{RD}$  ages.

32 531 Despite the strong evidence of metasomatism in many mantle xenoliths of the EARS, which is  
33 532 expected to affect the Os isotope signature (Chesley et al., 1999; Burton et al., 2000; Reisberg et al.,  
34 533 2004; Alemayehu et al., 2019), it is surprisingly that whole rock  $T_{RD}$  ages are able to preserve age  
35  
36 534 information. For instance, a positive correlation between  $^{187}\text{Os}/^{188}\text{Os}$  and  $^{187}\text{Re}/^{188}\text{Os}$  (Chesley et al.,  
37  
38 535 1999, Burton et al., 2000, Reisberg et al., 2004) is inconsistent with a “frozen” Os isotopic signature  
39  
40 536 after melt extraction. Similarly, the inverse correlation between  $^{187}\text{Os}/^{188}\text{Os}$  and indices of melt  
41  
42 537 depletion (e.g.,  $\text{Al}_2\text{O}_3$ , Lu, CaO) can be explained by incomplete extraction of Re after partial melting,  
43  
44 538 which is however at odd with what expected after large degrees of partial melting required for the  
45  
46 539 formation of highly refractory cratonic mantle roots. This points towards metasomatic event that  
47  
48 540 introduced radiogenic Os along with  $\text{Al}_2\text{O}_3$ , Lu, CaO and Re (e.g., Chesley et al., 1999; Reisberg et  
49  
50 541 al., 2004). The fact that these correlations are not ubiquitously observed could reflect a limited  
51  
52 542 metasomatic influence on  $T_{RD}$  ages, thus affecting only few samples. Moreover, Re-enrichment,  
53 543 which is characteristic of melt/fluid infiltrations shortly before or during the eruption of the magma  
54  
55 544 hosting the xenolith (e.g., Chesley et al., 1999), does not have a large effect on the Os isotope  
56  
57 545 composition when the eruption is relatively recent. More speculative scenarios to explain the  
58 546 occurrence of primary Os isotope signatures in mantle domains affected by metasomatism include Os  
59  
60 547 mobility within the mantle in the form of sulfide melts (Reisberg et al., 2004). Thus, through

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3 548 migration of sulfide melts, the Os isotopic composition could be transferred from its original mantle  
4 domain to other ones without being significantly affected and, therefore, reflecting the formation of  
5 549 residual mantle domains (Bragagni et al., 2017).

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## 6. Mantle domains involved in the genesis of rift-related volcanism

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The geochemical features of EARS mafic magmas strongly depend on the interaction and mixing between a common deep plume-derived material and other mantle domains (e.g., Rooney, 2020d and references therein). A depleted asthenospheric component (DMM), which is clearly observed in the magmatism of the Gulf of Aden and Red Sea (e.g. Altherr et al., 1990), is likely involved in some of the magmatism of the EARS. Indeed, the contribution from DMM can explain the low Th/Yb e Ta/Yb (Figure 3) and the radiogenic Nd isotope composition (Figure 7) of the first products of Ethiopia magmas (Pik et al., 1999; Kieffer et al., 2004) and some of the younger Afar magmas (Barrat et al., 2003, Daoud et al., 2010). Similarly, a DMM involvement is also invoked in the Turkana area (Furman et al., 2006, Figure 7). As such, a variable contribution of a depleted mantle to the inferred Afar plume composition can explain the isotope variability towards high  $^{143}\text{Nd}/^{144}\text{Nd}$  and low  $^{87}\text{Sr}/^{86}\text{Sr}$ . However, the largest variations in Sr-Nd-Pb isotopes in the EARS magmas, roughly defining two trends, one with radiogenic  $^{87}\text{Sr}/^{86}\text{Sr}$  (exceeding the oceanic EMII end-member) and moderately unradiogenic  $^{143}\text{Nd}/^{144}\text{Nd}$ , and the other with extremely unradiogenic  $^{143}\text{Nd}/^{144}\text{Nd}$  and moderately low  $^{87}\text{Sr}/^{86}\text{Sr}$  (Figure 7), point towards the involvement of others contributions from different mantle end-members such as the SCLM.

The contribution of the SCLM to magma genesis in the EARS might be extremely complex and variable (e.g., Furman and Graham 1999, Rosenthal et al., 2009; Foley et al 2012; Rooney, 2020d), although its role seems widespread and ubiquitous, as already observed in Ethiopia (e.g. Natali et al 2016; Feyissa et al 2017, 2019; Beccaluva et al 2009; Rooney, 2017), Western Branch (e.g., Spath 2001, Rosenthal et al 2009; Furman and Graham 1999), Eastern Branch (e.g. Rogers, 2006), and Turkana (e.g. Meshesha et al 2011; Furman et al 2016). The Sr-Nd isotope signature of mantle xenoliths (Figure 11) can help to understand the link between SCLM and the EARS mafic magmas. In the same diagram, the isotope compositions of carbonatites from this area are also reported, as these magmas are thought to derive from strongly metasomatised SCLM sources (e.g. Foley et al., 2012; Rooney, 2017; Rooney 2020d). It is striking to observe that the isotopic signature of the mafic magmas nicely overlap with that of mantle xenoliths and carbonatites. Notably, mantle xenoliths points to even more extreme compositions (Figure 11). Since the isotopic composition of the mantle xenoliths are consistent with two events (Archean and Pan-African, section 5.1), the two trends in the Sr-Nd space observed in EARS magmas can be ascribed to the interaction of the Afar mantle plume

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3 583 with these two different SCLM domains. In this scenario, magmas trending towards extremely  
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5 584 radiogenic  $^{87}\text{Sr}/^{86}\text{Sr}$  and moderately unradiogenic  $^{143}\text{Nd}/^{144}\text{Nd}$  are affected by the Pan-African SCLM  
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7 585 (Mobile Belt of [Rooney, 2020b](#)), while magmas pointing towards very low  $^{143}\text{Nd}/^{144}\text{Nd}$  show an  
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9 586 affinity with the Archean lithosphere (Craton of [Rooney, 2020c](#)). Although there are several lines of  
10 587 evidence that the SCLM experienced multiple events and types of metasomatism (e.g., [Rosenthal et](#)  
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12 588 [al., 2009; Rooney 2020d](#)), its Sr-Nd isotope signature is roughly consistent with these two  
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14 589 components, at least at the very large scale of [Figure 7](#).

15 590 The extreme radiogenic isotope compositions recorded in the Western and Eastern Branches can be  
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17 591 therefore obtained by admixing the Afar mantle plume with the two different domains of the SCLM  
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19 592 (Pan-African and Archean) as recorded by ultramafic mantle xenoliths ([Figures 7, 8, 11](#)). In this  
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21 593 scenario, the trend observed in the older magmas of the Eastern Branch can be explained by a  
22 594 contribution of both SCLM domains: the interaction with the Archean lithosphere can result in the  
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24 595 general lower  $^{143}\text{Nd}/^{144}\text{Nd}$  of these magmas than those of the Afar mantle plume, while the Pan-  
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26 596 African lithosphere can be responsible for the trend towards high  $^{87}\text{Sr}/^{86}\text{Sr}$  with only a moderate  
27 597 decrease in  $^{143}\text{Nd}/^{144}\text{Nd}$  ([Figures 7a, 11a](#)). Most of the younger samples, with extreme unradiogenic  
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29 598 Nd isotope compositions, point towards a stronger contribution from the Archean SCLM, whereas  
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31 599 the remaining samples, with radiogenic Sr isotope compositions, indicate the contribution of the Pan-  
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33 600 African SCLM ([Figures 7b, 11a](#)). Notably, the samples of the Eastern Branch with the strongest  
34 601 Archean signature are from volcanic centres located within the Tanzania craton, while the Pan-  
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36 602 African signature is recorded in samples from the Mozambique mobile belt along the craton  
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38 603 boundary. Overall, there is a trend of increasing contributions with time of SCLM domains to magma  
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40 604 genesis (see below).

41 605 The Western Branch, where only samples from the youngest period of volcanic activity are available,  
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43 606 shows a more uniform trend towards the Pan-African lithosphere, reaching the most extreme  
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45 607 radiogenic Sr isotope compositions ([Figures 7b and 11a](#)). Many studies highlighted the complexity  
46 608 of the different metasomatic processes, involving a significant role of silicate and carbonatitic melts  
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48 609 in the SCLM of the Western Branch ([Furman and Graham, 1999; Foley et., 2012; Rosenthal et al.,](#)  
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50 610 [2009](#)). However, from the large scale of [Figure 7](#), the isotopic composition of the Western Branch  
51 611 magmas spreads along a rather distinct trend, dominated by Pan-African SCLM, suggesting that the  
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53 612 overall metasomatism within the Western Branch mimics this general trend. [Furman and Graham](#)  
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55 613 [\(1999\)](#) identified a common lithospheric mantle (CLM) with a composition of  $^{87}\text{Sr}/^{86}\text{Sr} = 0.7050$  and  
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57 614  $^{143}\text{Nd}/^{144}\text{Nd} = 0.51264$ , which falls within our Pan-African trend, although not reaching the extreme  
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59 615 radiogenic Sr isotopes observed in the Western Branch magmas ([Figure 7](#)).

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3 616 The observed temporal and spatial variations based on Sr and Nd isotopes are also consistent with  
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5 617 what observed on Pb isotopes (Figures 8, 11b). The spread towards radiogenic Pb isotopes observed  
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7 618 especially in the Eastern Branch samples are consistent with a SCLM contribution that was enriched  
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9 619 in carbonatite-like domains characterised by high time-integrated U/Pb (Figure 11b). Such a  
10 620 component can be observed both in ultramafic mantle xenoliths and in carbonatite magmas (Figure  
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12 621 11b). This extreme radiogenic Pb isotope signature points towards a HIMU-like mantle component.  
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14 622 Such a component was previously interpreted to potentially reflect an additional mantle plume or a  
15 623 distinct portion of the African Superplume (e.g., Rooney 2020d for a review on the topic). In  
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17 624 particular, the HIMU plume component was proposed for the Turkana and South Ethiopia areas (the  
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19 625 Kenya plume, Rogers et al., 2000), possibly reflecting the deep recycling in the mantle of about 30%  
20 626 ancient (1.7–2 Ga) hydrothermally altered subducted oceanic crust (Furman et al., 2006, Nelson et  
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22 627 al., 2012). The Turkana and South Ethiopia magmas, however, do not have the Pb isotope signature  
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24 628 of HIMU basalts as displayed by the Eastern Branch magmas (Figures 8). This suggests that the  
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26 629 HIMU signature, coupled with extreme Sr-Nd isotope compositions (Figures 11a), can be obtained  
27 630 with carbonatite metasomatism (Figures 11a), which is well known to affect portions of the local  
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29 631 SCLM (e.g., Rooney et al., 2014; Muirhead et al., 2020). The  $^{187}\text{Os}/^{188}\text{Os}$  of Turkana magmas are  
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31 632 slightly more radiogenic than the inferred composition of the Afar mantle plume, which could reflect  
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33 633 deep recycling of oceanic crust (Nelson et al., 2012) but also a similar component in the local SCLM.  
34 634 Indeed, metasomatism can enrich the SCLM in Re as attested, for example, by the radiogenic  
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36 635  $^{187}\text{Os}/^{188}\text{Os}$  of eclogite and pyroxenites from cratonic settings (e.g., Aulbach et al., 2009). As such  
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38 636 our interpretation may also support an origin of the HIMU signature in the SCLM by metasomatism  
39 637 or dripping (Rooney et al., 2014; Furman et al., 2016). Whatever the name of this signature, the  
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41 638 important aspect is that chemical and isotopic variations of the EARS magmas require that the SCLM  
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43 639 was affected by mantle metasomatism (e.g., Rooney, 2020d; Furman 2007), which is best observed  
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45 640 in Pb isotopes (likely in the form of carbonatite metasomatism), but that does not seem to have fully  
46 641 overprinted the Sr-Nd-Os isotope signature inherited from the Archean and Pan-African events.  
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48 642 Interestingly, based on Pb isotopes, a metasomatic age of ca. 500 Ma (i.e., Pan-African) was proposed  
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50 643 for the source of the Western Branch (Vollmer and Norry 1983) along with an older event of ca. 1  
51 644 Ga (Rogers et al., 1998).  
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53 645 A slightly different mechanism is required to explain the Pb isotope composition of the Western  
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55 646 Branch magmas, pointing to an EM II-like mantle component (Figure 8). This signature, coupled with  
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57 647 their Sr-Nd isotope composition (Figure 7), requires the contribution of an old crustal component in  
58 648 the SCLM (e.g., Furman et al., 2007; Castillo et al., 2014), likely referred to the Pan-African event  
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60 649 (Figures 9, 11).

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3 650 In summary, the radiogenic isotope signature of the SCLM, as recorded by mantle xenoliths and  
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5 651 carbonatites, covers the whole compositional range observed in EARS mafic magmas. Although the  
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7 652 picture might be more complex when considering the variable chemical composition of metasomatic  
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9 653 melts/fluids and their impact on the small-scale chemical variations of the EARS magmas (e.g.,  
10 654 [Furman et al., 2016](#); [Rooney 2020d](#)), our findings highlight that the SCLM contains the geochemical  
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12 655 signature required to explain the large variations observed in the EARS magmatism throughout the  
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14 656 interaction with the Afar mantle plume and to a lesser extent the DMM asthenospheric mantle.  
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### 16 17 658 **6.1. Mechanism for the contribution of the SCLM in the EARS magmas**

18 659  
19 660 The SCLM, which comprises an old Archean domain and a more recent Pan-African component,  
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21 661 plays a crucial role in the genesis of the EARS magmas and their geochemical fingerprints. Another  
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23 662 important outcome of our analysis is that, at large-scale, the SCLM contribution seems to increase  
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25 663 with time, as magmas of the third younger period of activity (<10 Ma) show the more extreme isotopic  
26 664 variation.

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28 665 Melt production from the SCLM is, however, paradoxical because it is much colder than the  
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30 666 convective asthenosphere, and it should be highly depleted in those major elements necessary for  
31 667 basalt generation (e.g., [Arndt and Christensen, 1992](#)). Enrichment processes involving percolation of  
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33 668 silicate melts and volatile-rich fluids have been, however, proposed to explain re-fertilisation and  
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35 669 melting of the SCLM at temperatures below the solidus of dry peridotite (e.g., [Hawkesworth et al.,](#)  
36 670 [1984](#); [Menzies and Hawkesworth, 1987](#); [Stolz and Davies, 1988](#); [Ionov et al., 2002](#)). Other factors  
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38 671 including extension rate, lateral temperature gradients, inputs of external heat by mantle plumes  
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40 672 impinging and flattening at the base of the lithosphere may also contribute to SCLM melting (e.g.,  
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42 673 [McKenzie and Bickle, 1988](#)).

43 674 In the case of the EARS, re-fertilisation of the SCLM likely occurred during the Archean and Pan-  
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45 675 African events ([Figure 9](#)) and different mechanisms for explaining SCLM melting have been  
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47 676 proposed. The arrival of the hot plume material at the base of the lithosphere could provide the heat  
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49 677 necessary to trigger melting at the base of the lithosphere (e.g., [Rogers et al., 1998](#); [Beccaluva et al.,](#)  
50 678 [2009](#)), especially in easily fusible metasomatic portions ([Steiner et al., 2022](#)). Alternatively, portions  
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52 679 of the bottom of the lithosphere could be physically transported through later advection into a region  
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54 680 of melting near the boundary of thick lithospheric domains such as those expected at the margin of  
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56 681 cratons ([Muirhead et al., 2020](#)). Another possibility is represented by drip melting, which requires  
57 682 metasomatised lithospheric portions that first sink into the asthenosphere due to their higher density  
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59 683 and then melt ([Furman et al., 2016](#)). Geochemically it is difficult to discriminate between these  
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3 684 different scenarios, but it is possible to do some considerations according to the time variations in the  
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5 685 magmatism as discussed in previous sections.  
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7 686 The scenario we propose is represented by initial plume arrival with uplift and volcanism in the  
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9 687 absence of significant extension (e.g., [Corti, 2009](#)). In these conditions, the interaction between the  
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11 688 uprising plume and the lithosphere is limited, with the lithosphere being too cold to melt. Instead, in  
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13 689 the rigid lithosphere, the development of throughgoing fractures and faults during plume-related uplift  
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15 690 facilitates direct transport of mantle melts *en route* to the surface ([Figures 7, 8](#); e.g., [Beccaluva et al.,](#)  
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17 691 [2009](#)). This well explains the predominance of the plume component in the initial melts (40-25 Ma  
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19 692 and 25-10 Ma), a scenario which is appropriate for Afar, Ethiopia and Turkana ([Figures 7, 8](#)). The  
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21 693 minor SCLM signature in these magmas may be related to local melting of limited portions of the  
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23 694 lithosphere in response to an increase of the geothermal gradient related to the arrival of the plume  
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25 695 (e.g., [Beccaluva et al., 2009](#)). Successively, due to the continued spreading of the plume head beneath  
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27 696 the plate and the continued heating of the lithosphere, the SCLM overpassed the solidus and started  
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29 697 to contribute significantly to melt production. This process was further enhanced in the youngest  
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31 698 period of volcanic activity (10-0 Ma) by the onset of the main rifting phases, with progressive thinning  
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33 699 of the lithosphere and progressively shallower melting. In this phase, the increase in the SCLM  
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35 700 component may be also enhanced by an overall progressive weakening of the uprising plume and a  
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37 701 relative motion of the plume head away from Ethiopia, Afar and the Turkana depression ([Hassan et](#)  
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39 702 [al., 2020](#)). This process therefore may explain the overall large-scale increase of the SCLM-like  
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41 703 signature (Archean and Pan-African) in the recent EARS magmas (10-0 Ma, [Figures 7, 8](#)).  
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43 704 It is worth noting, however, that the picture may be more complex and apparently contradictory at a  
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45 705 smaller temporal and spatial scale. For instance, the isotopic signature in the most recent magmatism  
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47 706 of the Turkana and Afar regions (< 2 Ma) points to that of the Afar plume ([Furman et al., 2006](#)), with  
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49 707 an absence of SCLM. In the case of the Afar, the lack of SCLM signature in the recent magmas is  
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51 708 readily explained by the extremely thin or absent lithospheric mantle of the area (e.g., [Rychert et al.,](#)  
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53 709 [2012](#)), which is close to a phase of oceanisation (e.g., [Bastow and Keir, 2011](#)). For the Turkana  
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55 710 depression, the waning of the SCLM component might be related to vanishing of an episode of drip  
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57 711 melting ([Furman et al., 2016](#)). Similarly, also the locally thin lithosphere in portions of the Western  
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59 712 Branch was interpreted as due to drip melting and the variation in Sr-Nd isotope composition as a  
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61 713 consequence of removal of the lower portions of a layered lithosphere ([Lawrence et al., 2022](#)).  
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## 7. Summary: spatial and temporal variations in magma production through the EARS

58 715 The scenario of magma production along the EARS obtained from our critical analysis of the  
59 716 available geochemical data of mafic magmas and mantle xenoliths from different sectors of the rift,  
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3 718 integrated with new trace element and Sr-Nd isotope data from the Main Ethiopian Rift and Turkana  
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5 719 depression, is summarised and illustrated in [Figure 12](#). It is important to remark here that the  
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7 720 distribution of samples in our database does not represent, from both volumetric and temporal points  
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9 721 of view, the actual distribution of the volcanic products (see paragraph 3). Despite these limitations,  
10 722 our database is fully representative and our analysis offers a geochemical picture which is consistent  
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12 723 with previous investigations of magma production in the EARS (e.g., [Rooney, 2020d](#) and reference  
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14 724 therein) and allows additional constraints on these processes.

15 725 In the time interval between 45 and 25 Ma ([Figure 12a](#)) volcanic activity was limited to Ethiopia,  
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17 726 Afar and the Turkana tectonic domains, affected by impingement of the upraising mantle plume  
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19 727 which caused volcanism and plateau uplift well before the main rifting episodes. Consequently,  
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21 728 magmas from this older phase mainly involved a plume component with minor contributions from  
22 729 the SCLM (Archean domain in the Turkana depression and Pan-African domain in Afar and Ethiopia)  
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24 730 and the DMM.

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26 731 In the second time interval (25-10 Ma), the DMM component disappeared in both the Afar and  
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28 732 Ethiopia, whereas it was still involved in the magma genesis of the Turkana depression which did not  
29 733 significantly change its signature with respect to the first period of volcanism ([Figure 12b](#)). The  
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31 734 available data from the Eastern Branch indicate the presence of a mantle plume component associated  
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33 735 with variable contributions from the SCLM in both its Archean and Pan-African domains ([Figure](#)  
34 736 [12b](#)).

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36 737 The youngest period of volcanic activity (10-0 Ma) corresponds to the main rifting phases all along  
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38 738 the EARS ([Figure 12c](#)): widespread volcanism still records a significant contribution from the  
39 739 upraising plume material and an increasing signature of the Archean and Pan-African SCLM  
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41 740 domains. Specifically, the Archean SCLM contributed to the magmatism of the Eastern Branch,  
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43 741 where rifting propagated within the Tanzanian craton, while the Pan-African SCLM influenced the  
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45 742 magmatism outside the craton along the Pan-African mobile belts in both the Eastern and Western  
46 743 Branches. Carbonatite magmas as well, which might have been generated directly within the carbon-  
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48 744 rich Tanzanian lithospheric mantle ([Eggler and Bell, 1989](#)) or through liquid immiscibility from  
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50 745 silicate melts ([Brooker and Kjarsgaard, 2011](#)), contributed to the geochemical signature of the Eastern  
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52 746 and Western mafic magmas of the EARS ([Figures 6, 11b](#)).

53 747 Overall, our critical analysis of the screened database indicates that the trace element and radiogenic  
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55 748 isotope signatures of the EARS mafic magmas are fully consistent with a relatively homogenous,  
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57 749 single mantle plume ([Rogers et al., 2000](#); [George and Rogers, 2002](#); [Rooney et al., 2012](#); [Castillo et](#)  
58 750 [al., 2020](#)) which is variably contaminated, depending on temporal and spatial distribution ([Figure 12](#)),  
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60 751 by different Archean and Pan-African SCLM domains (phlogopite- and amphibole-bearing



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3 752 metasomatised peridotites) plus a depleted asthenospheric component (DMM). The plume  
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5 753 contribution in the volcanism of the Eastern Branch starting at 25 Ma, and of the Western Branch  
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7 754 starting at 10 Ma is consistent with the model invoking a single Afar plume migrating southwards  
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9 755 relative to the African plate (Hassan et al., 2020). The increasing contribution of the different SCLM  
10 756 domains is related to the main rifting phases especially in Ethiopia and Eastern and Western Branches,  
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12 757 in which extensional processes have progressively allowed melting of the SCLM. In this view and at  
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14 758 our scale, a simple increasing of lithosphere melting, provides the simplest mechanism to explain the  
15 759 overall variability with time without the need to evoke other complex models involving the  
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17 760 occurrence of multiple mantle plumes.

### 22 763 **Acknowledgments**

23  
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28  
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## Figure captions

**Figure 1.** Cenozoic magmatic and tectonic activity in East Africa and parts of the Arabian Peninsula (modified from [Chorowicz, 2005](#); [Rooney, 2017](#)), superimposed on a NASA-SRTM digital elevation model. The coloured circles indicate the location of samples used in this work; the different colours indicate the three main temporal periods in which the samples are grouped. Enclosed by black dashed lines are the five main tectonic domains in which the samples have been spatially subdivided. Inset shows the location of the East African Rift System.

**Figure 2.** Correlation plot of  $\text{CaO}/\text{Al}_2\text{O}_3$  vs MgO wt% for the EARS mafic (>4 wt% MgO) magmas during the two oldest periods (a) compared to the youngest period (b) of activity. Symbols and colours identify the different tectonic domains as described in the legend and detailed in the text. The compositional field of MORB and OIB lavas are also shown for comparison.

**Figure 3.** Distribution of the key trace elements ratio Th/Yb vs Ta/Yb during the oldest two periods of activity (a) and during the youngest period (b) of activity. Symbols and colours identify the different tectonic domains as described in legend and detailed in the text. The compositional field of MORB and OIB lavas are also shown for comparison. Logarithmic scale is used to better highlight sample variation.

**Figure 4.** Correlation plot of Yb (ppm) vs La/Yb for the products erupted during the two oldest periods (a) and during the youngest period (b-c) of activity showing the effect of mantle melting in the garnet and spinel stability fields, using a non-modal batch melting model. Mantle source compositions are a nominal Primitive Mantle ([McDonough and Sun, 1995](#)) (a-b), and a nominal SubContinental Lithospheric Mantle (SCLM, [McDonough and Sun 1995](#)) (c). Partition coefficients are from the compilation of [McKenzie and O'Nions \(1991\)](#) and [Kelemen et al., 2003](#). Symbols and colours are reported in the legend. The solid black and red lines represent melts derived from different melting degrees (from 0.1% to 20%) of a garnet and spinel lherzolite, respectively, whereas dashed grey lines represent mixing between garnet- and spinel-derived melts.

**Figure 5.** Distribution of Rb/Sr vs Ba/Rb trace element ratios showing the role of amphibole-bearing vs phlogopite bearing metasomatized mantle in the genesis of EARS mafic magmas during the oldest two periods (a) and the youngest period of activity (b). The composition of PM and SCLM is also shown ([McDonough and Sun, 1995](#)).

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5 809 **Figure 6.** Zr/Nb vs Nb/Ta during the two oldest periods (a) and during the youngest period (b) of  
6 810 activity. The compositional field of MORB and OIB lavas are also shown for comparison.  
8 811 Logarithmic scale is used to better highlight sample variation.

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12 813 **Figure 7.** Sr-Nd isotope variation through the EARS for the products of the two oldest periods (a) of  
13 814 activity compared to the youngest period of activity (b). The compositional field of MORB and OIB  
15 815 lavas are also shown for comparison. The star represents the composition of the Afar mantle plume  
16 proposed by [Rooney \(2020d\)](#) consistent with the values of Afar basalts with high  $^3\text{He}/^4\text{He}$  of [Castillo  
17 816 et al. \(2020\)](#).

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22 819 **Figure 8.**  $^{207}\text{Pb}/^{204}\text{Pb}$  vs  $^{206}\text{Pb}/^{204}\text{Pb}$  through the EARS for the products of the two oldest periods (a)  
23 of activity compared to the youngest period of activity (b). The compositional field of MORB and  
24 820 OIB lavas are also shown for comparison. Star as in [Figure 7](#).

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29 823 **Figure 9.** Sm-Nd isotope composition of the ultramafic xenoliths. The reference isochrons intercept  
30 the y-axis at Nd isotope composition consistent with the evolution of either the primitive (dashed  
31 824 lines) or the depleted mantle (solid lines) at 600 Ma (violet lines) and 2.6 Ga (green lines).

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36 827 **Figure 10.** Rhenium-depletion model ages ( $T_{\text{RD}}$ ) from mantle xenoliths of the EARS. The density  
37 probability plot is made assuming an uncertainty ( $1\sigma$ ) of 200 Ma on  $T_{\text{RD}}$  model ages (cf. [Pearson et  
38 828 al., 2007](#)) and using the PM composition of  $^{187}\text{Os}/^{188}\text{Os}=0.1296$  and  $^{187}\text{Re}/^{188}\text{Os}=0.4353$  ([Meisel et  
39 829 al., 2001](#)). All available literature whole rock data are plotted ([Burton et., al 2000](#), [Chesley et al.,  
40 830 1999](#), [Alemayehu et al., 2019](#), [Reisberg et al., 2004](#), [Meisel et al., 2001](#), [Becker et al., 2006](#)). A  
41 831 frequency histogram of  $T_{\text{RD}}$  ages from different regions is reported in the background. The frequency  
42 of each bin is shown on the right axis.

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50 835 **Figure 11.** Sr-Nd (a) and Pb (b) isotope signature of ultramafic xenoliths. The isotope compositional  
51 836 field of the mafic magmas characterising the five EARS tectonic districts have been reported for  
52 comparison together with the indicative composition of the DMM and Afar Plume ([Castillo et al.,  
53 837 2020](#)) mantle components.

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58 840 **Figure 12.** Schematic sketch-map cartoon summarising the scenario of magma production along the  
59 EARS inferred from the geochemical and isotopic characteristics of mafic magmas correlated with  
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842 carbonatites and ultramafic xenoliths and subdivided according to the main tectonic domains (Afar,  
843 Ethiopia, Turkana, Eastern Branch, and Western Branch), and to the three main temporal intervals  
844 (45-25 Ma, 25-10 Ma and 10-0 Ma), see text for detail. DMM: depleted asthenospheric mantle;  
845 SCLM: SubContinental Lithospheric Mantle.

For Review Only

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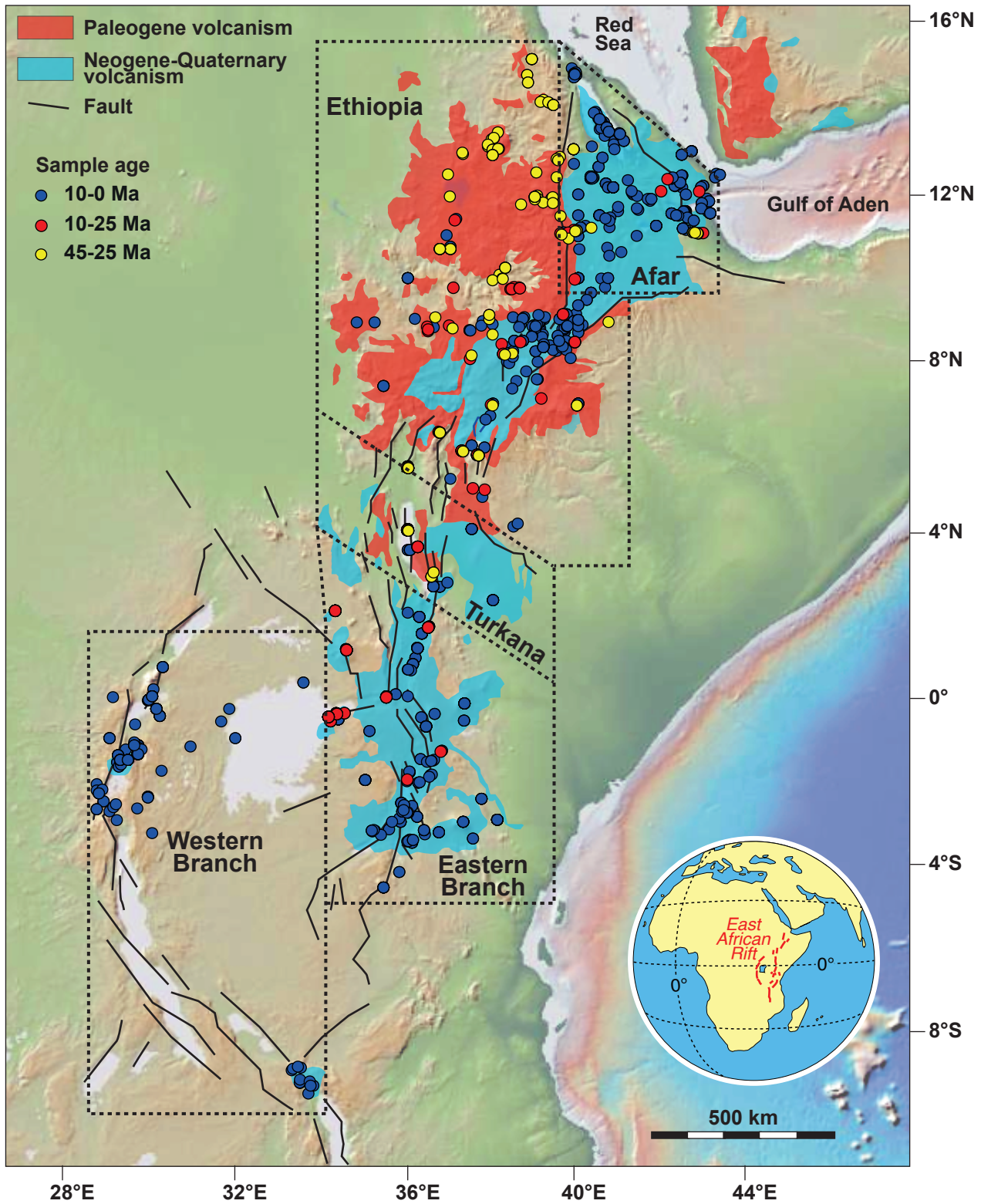


Figure 1

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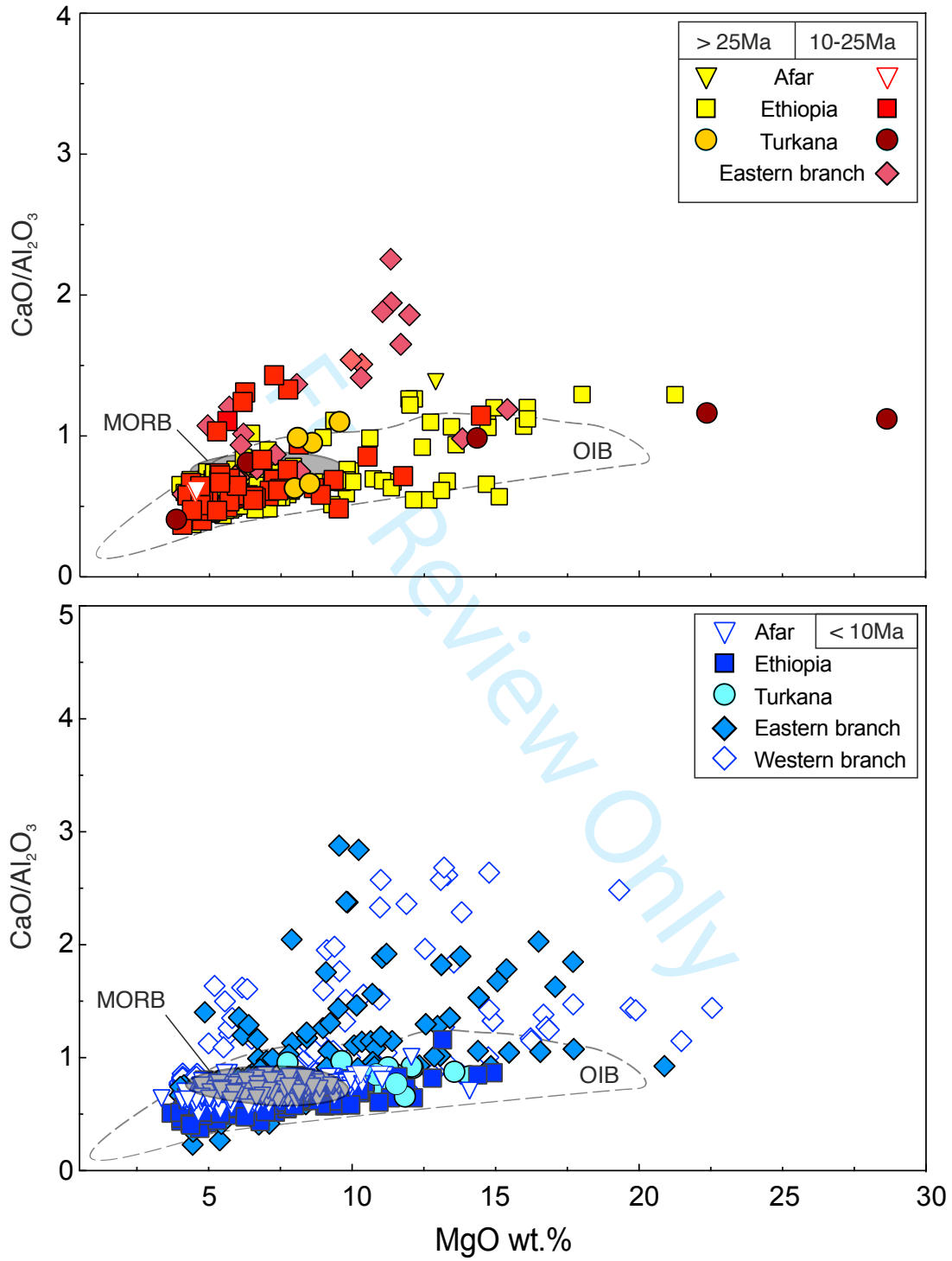


Figure 2

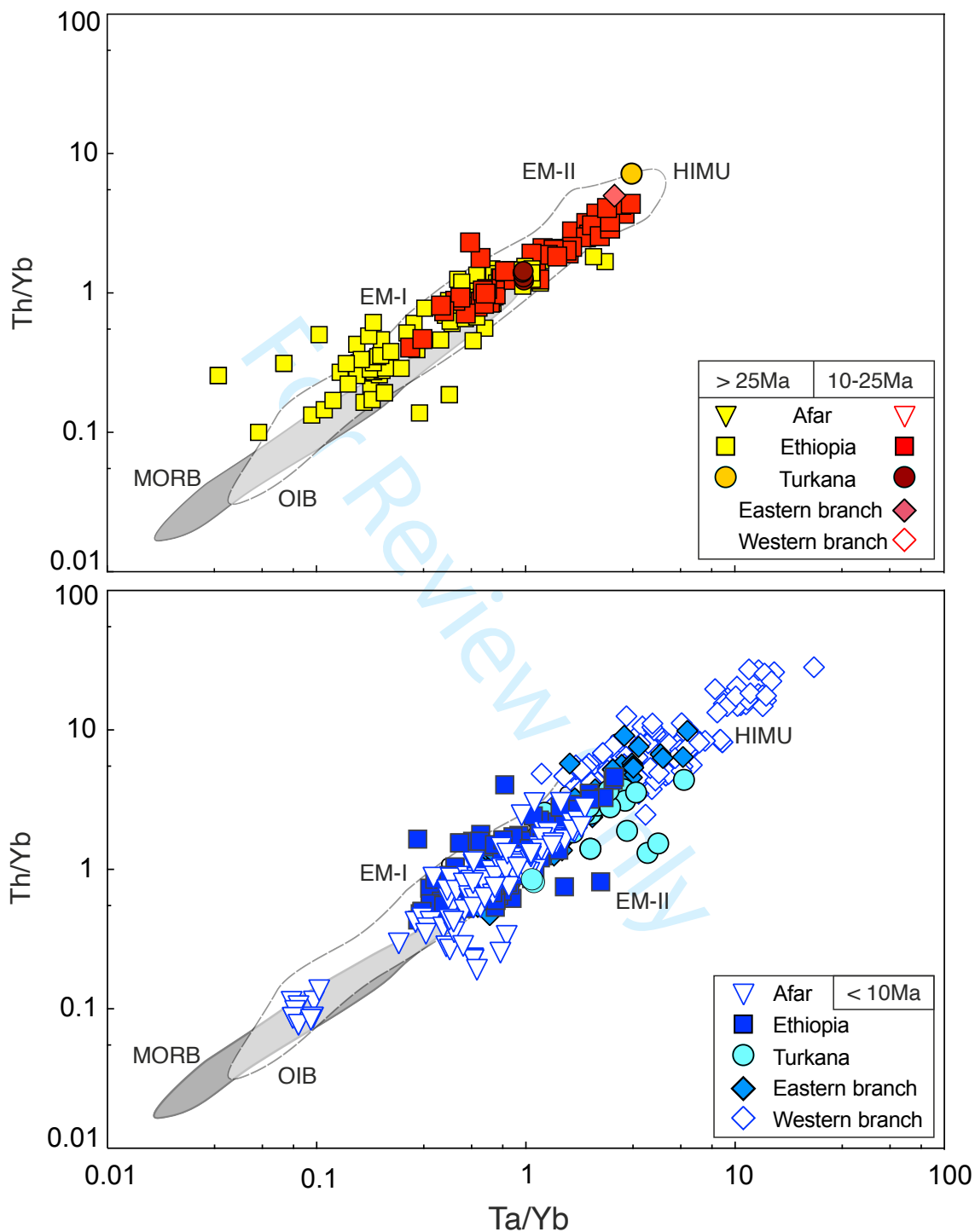


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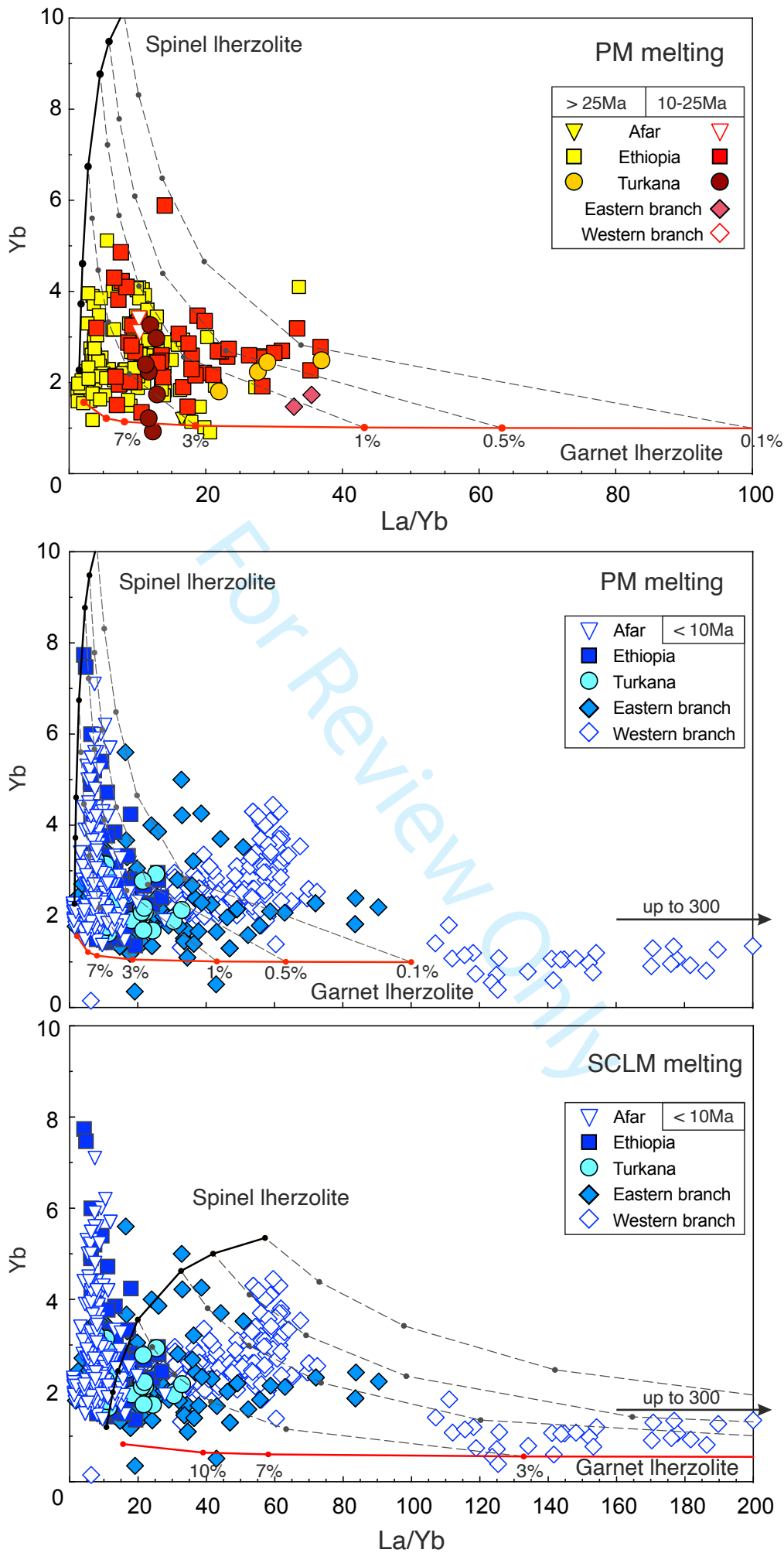


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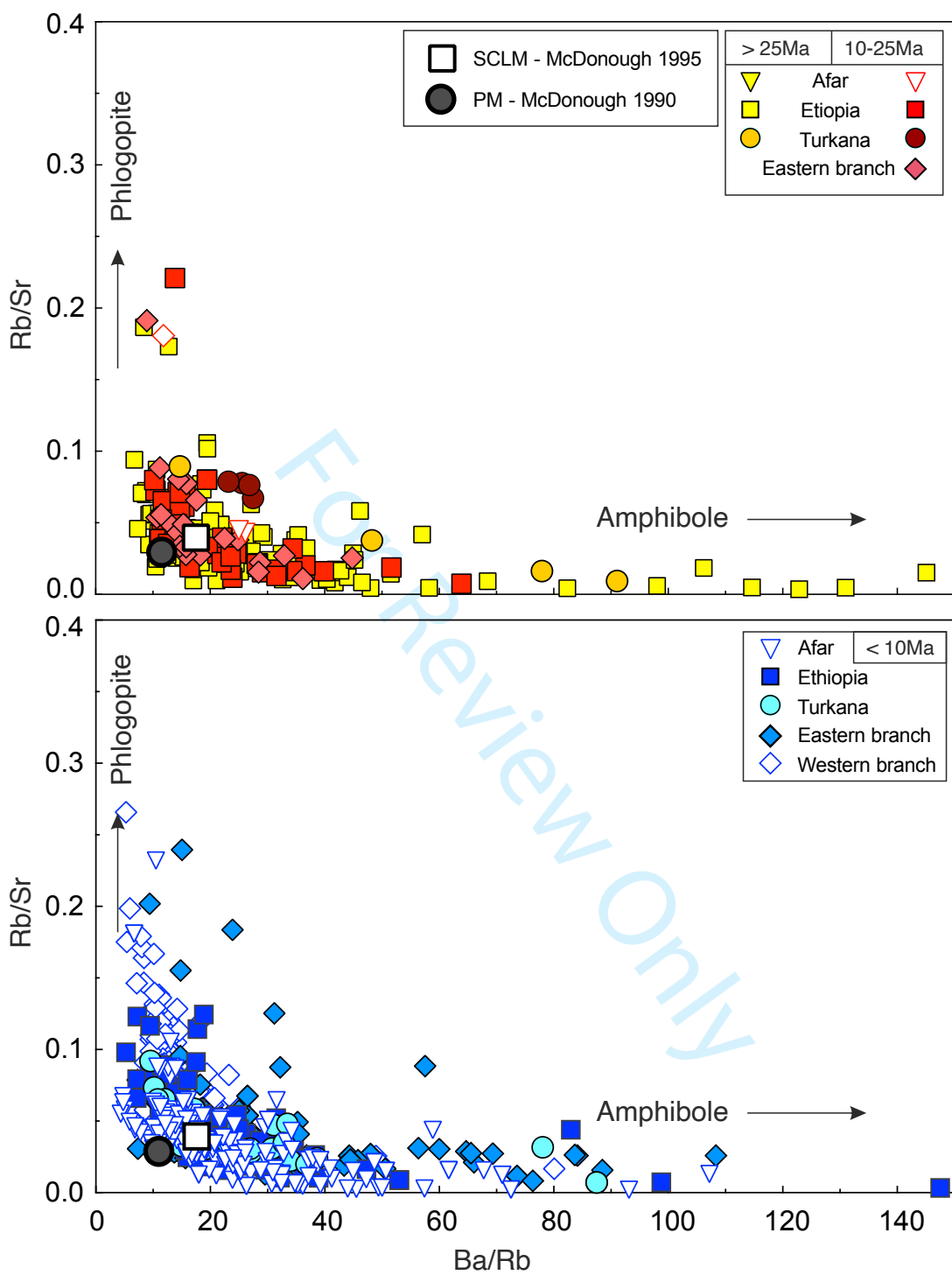


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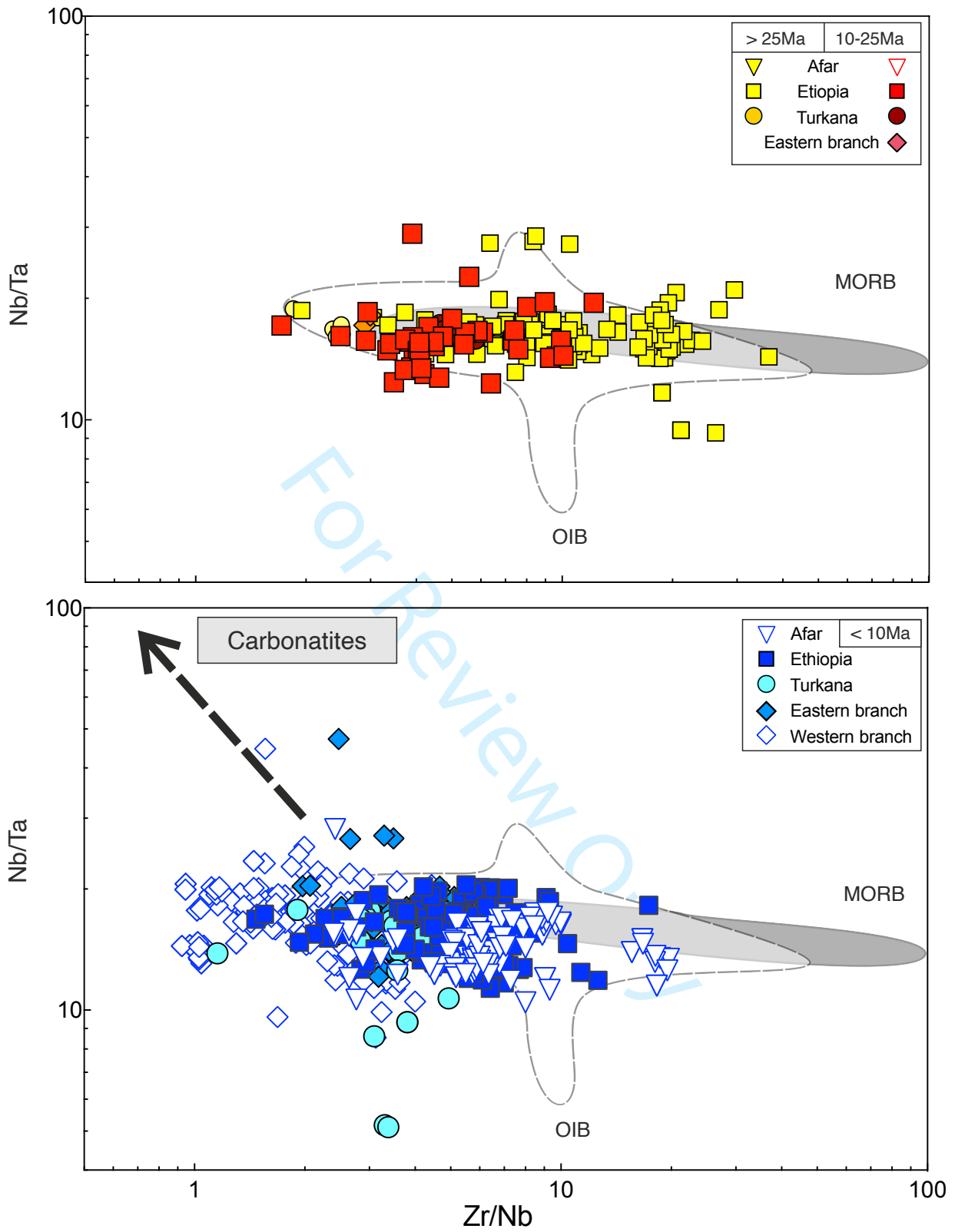


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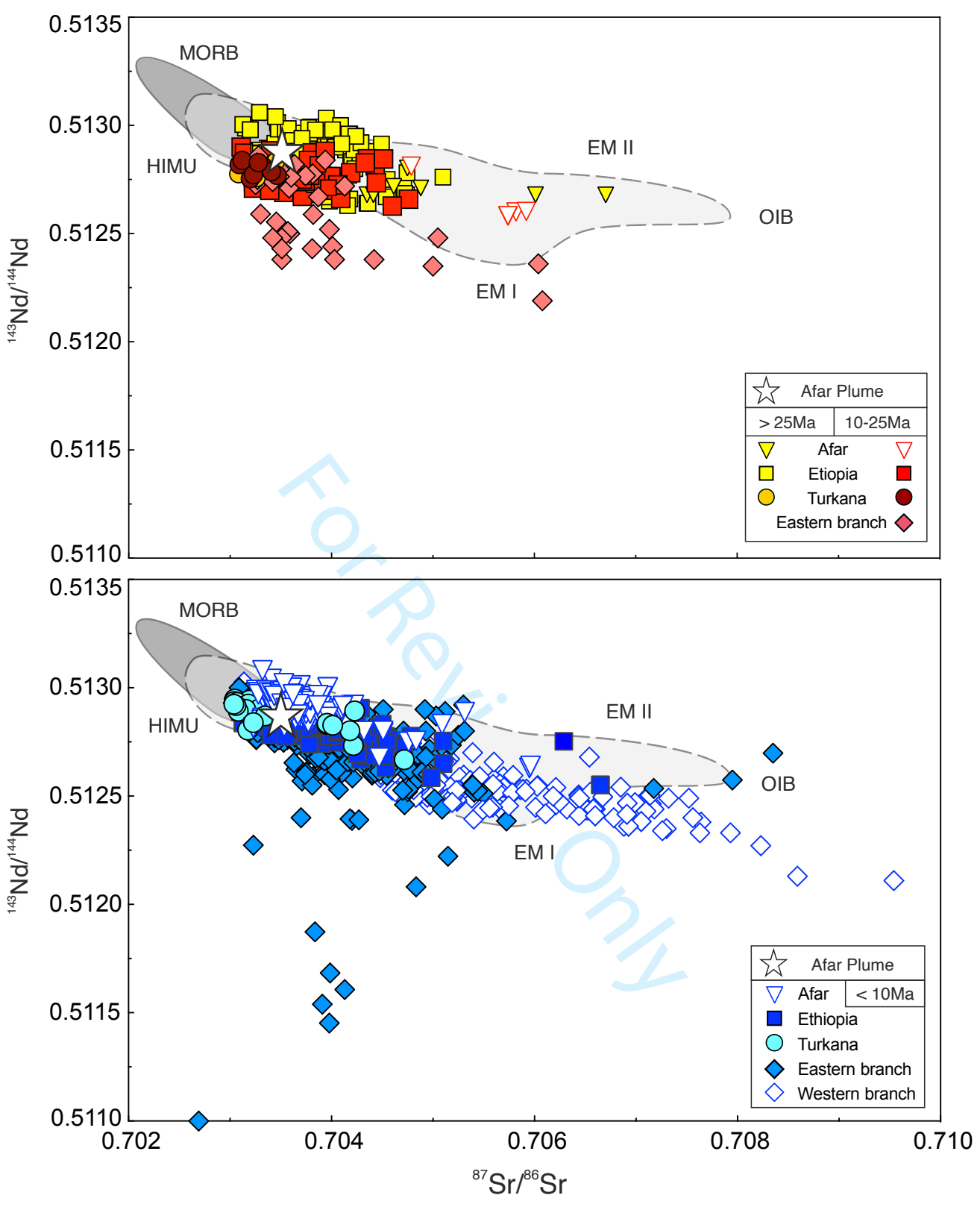


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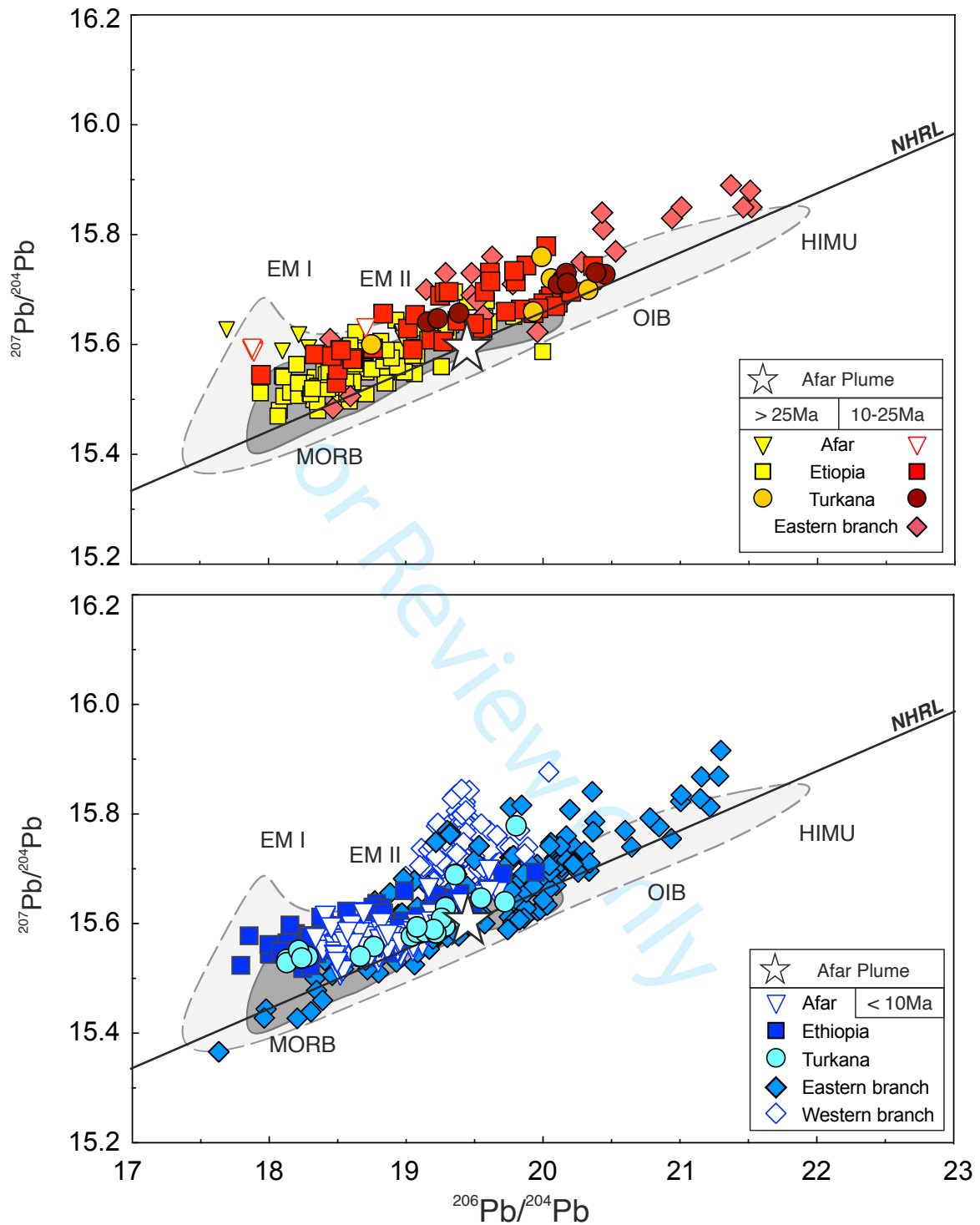


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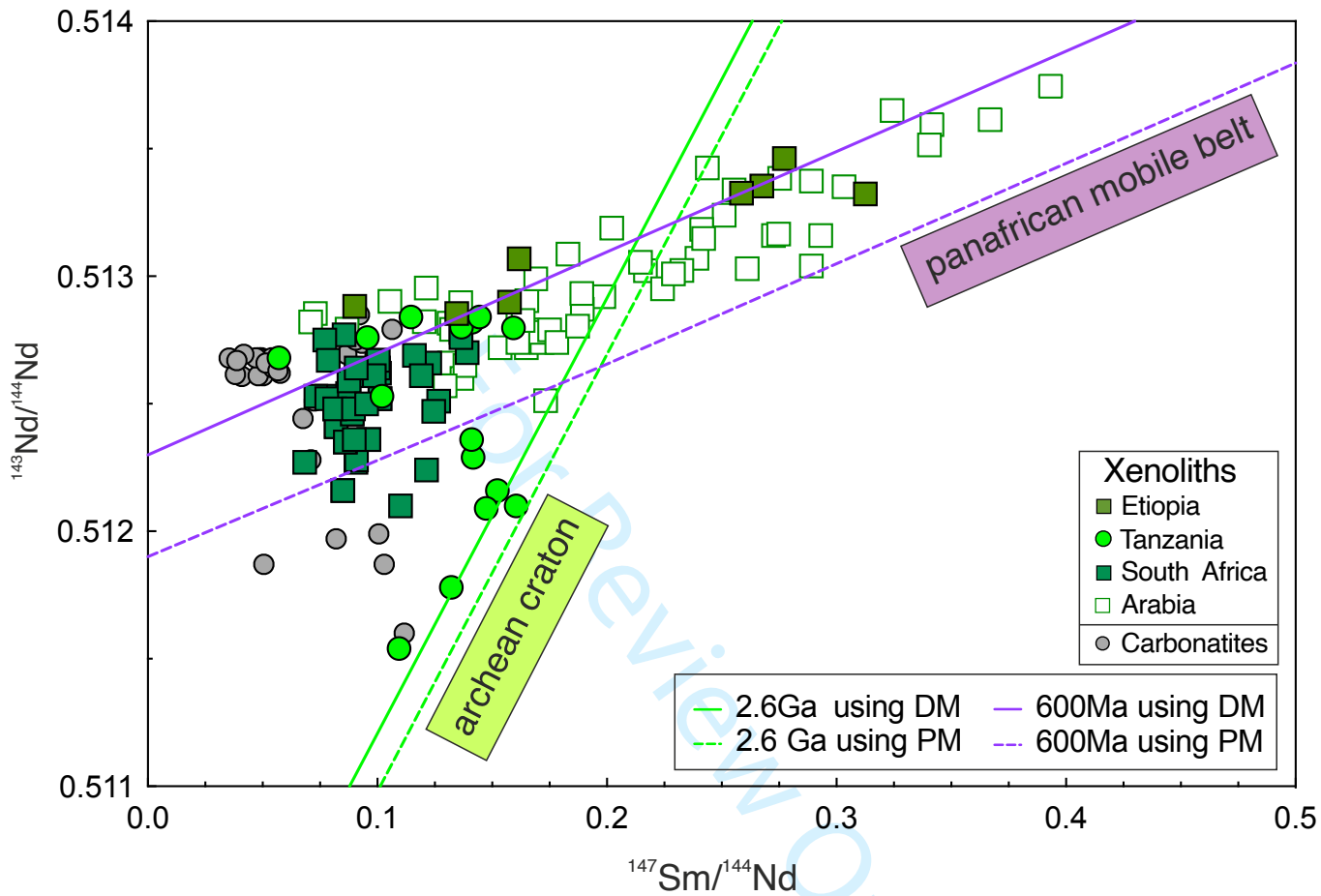


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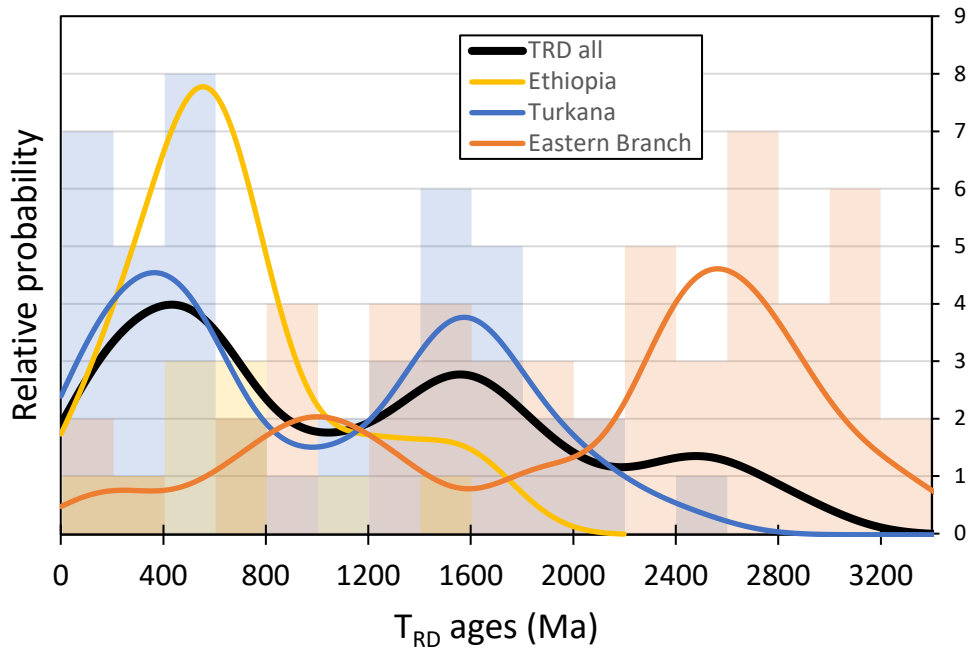
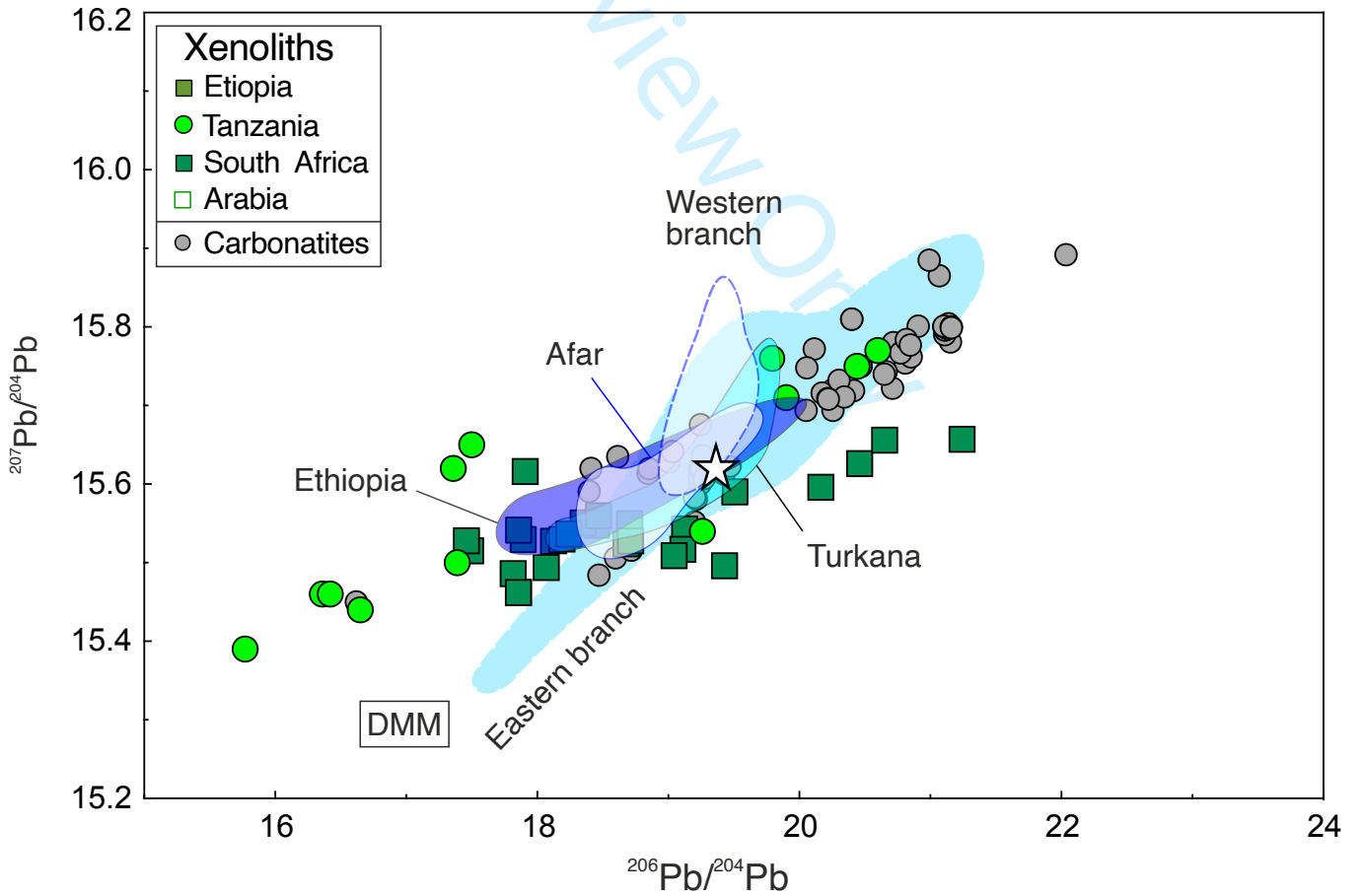
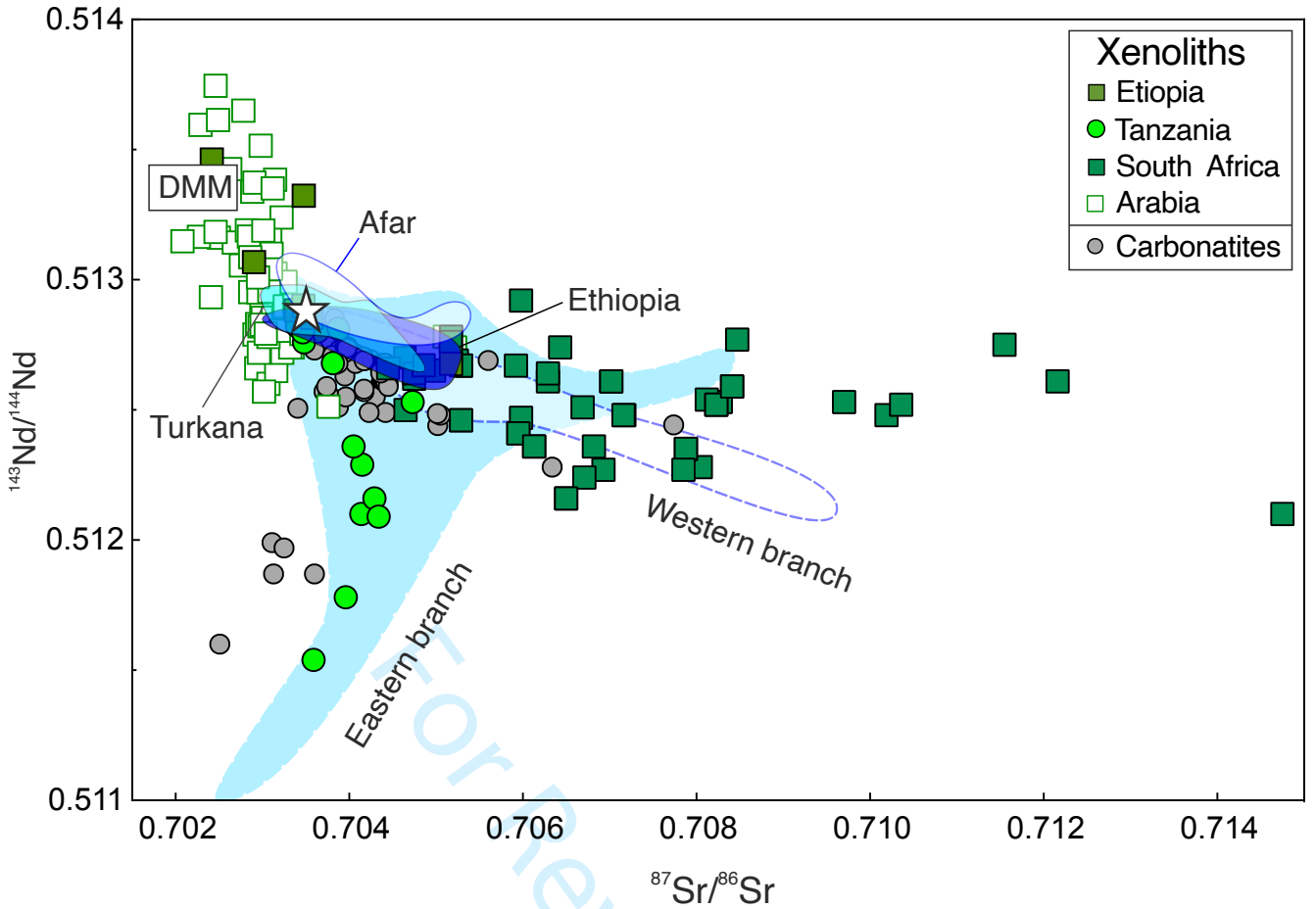


Figure 10



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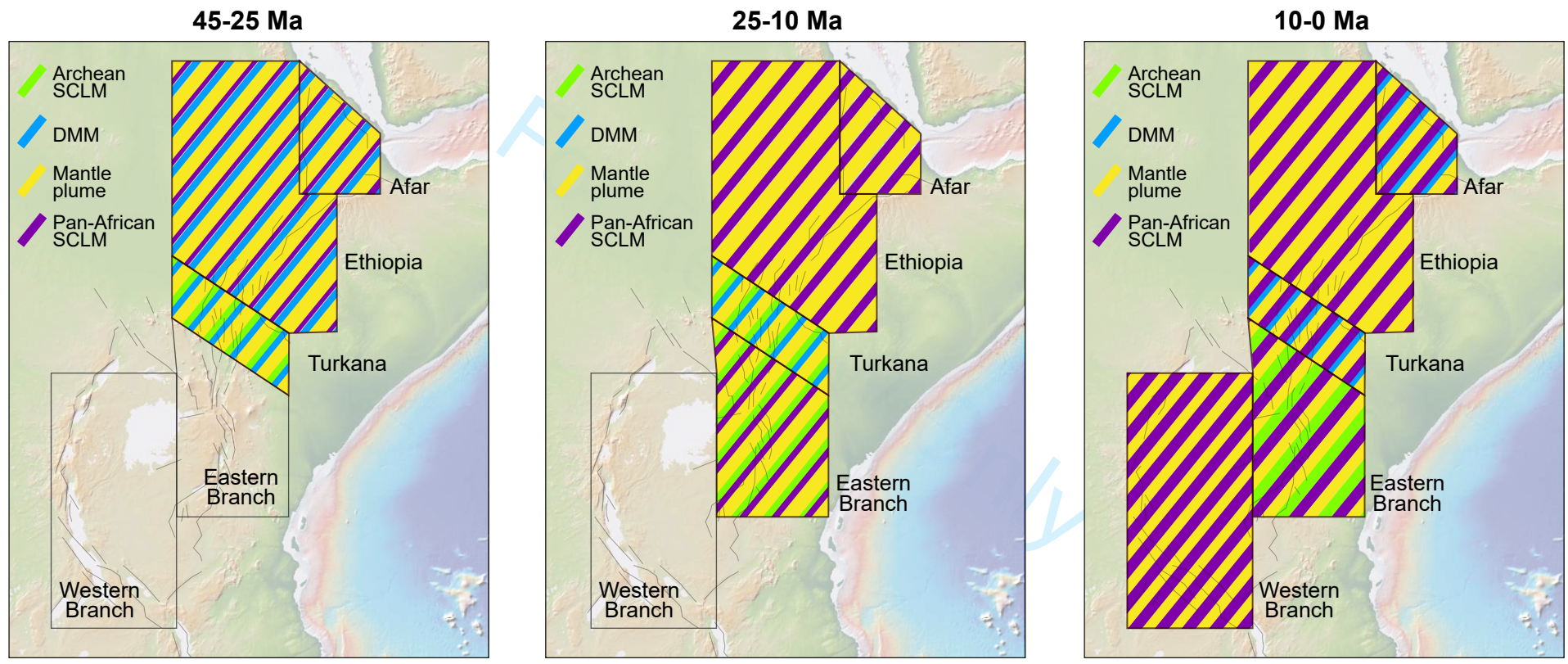


Figure 12

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## Supplementary Material

### Analytical methods

A new set of 27 samples were collected from a relatively wide area of the EARS from the Ethiopia (25 samples) and Turkana (2 samples) tectonic domains ([Supplementary Table1](#)).

The samples were prepared and powdered for whole-rock characterization of major, trace elements and radiogenic isotopes of Sr and Nd. The reported data are unpublished and presented here for the first time.

Major elements were determined at the Department of Earth Sciences of the University of Florence by X-Ray Fluorescence (XRF), according to the procedure of [Franzini et al. \(1972\)](#). MgO and Na<sub>2</sub>O were analysed through atomic absorption spectroscopy (AAS) and FeO by trititation. Loss On Ignition (LOI) was determined through gravimetry after heating the sample powders at 950° C.

Trace elements were determined by Inductively Coupled Plasma Mass Spectrometry (ICP-MS) at the Geowissenschaftliches Zentrum der Universität Gottingen (GZG) on a VG PQ2 system. Analytical uncertainties are within the significant digits reported in [Supplementary Table 1](#).

For isotope analyses, powder digestion and Sr-Nd purification were carried out in the clean laboratory ("Class 1000") of the Department of Earth Sciences of the University of Florence. Sample digestion procedure was performed by sequential HF-HNO<sub>3</sub>-HCl and elemental separation through specific chromatographic columns, as described in [Avanzinelli et al. \(2005\)](#). All measurements were performed at the Department of Earth Sciences of the University of Florence using a Thermo-Finnigan Triton-Ti<sup>®</sup> Thermal Ionisation Mass Spectrometer (TIMS), equipped with 9 movable collectors. <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd were measured dynamically and corrected using an exponential mass fractionation law to <sup>86</sup>Sr/<sup>88</sup>Sr = 0.1194 and <sup>146</sup>Nd/<sup>144</sup>Nd = 0.7219, respectively, as described by [Avanzinelli et al. \(2005\)](#). Replicate measurements of NBS 987 and La Jolla reference materials (0.710249 and 0.511856, respectively, [Thirlwall 1991](#)) gave mean values of <sup>87</sup>Sr/<sup>86</sup>Sr = 0.710248 ± 0.000013 (2s, n = 86) and <sup>143</sup>Nd/<sup>144</sup>Nd = 0.511846 ± 0.000007 (2s; n = 67). The Sr procedural blank was 270 pg, which is safely within the blank range of our lab for whole rocks procedures ([Avanzinelli et al. 2005](#)).

All data were age corrected to the initial isotope value and errors were fully propagated.

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**Table S1:** Major elements (wt%), trace elements (ppm) and Sr-Nd isotope composition of selected samples.

Sample	Note	Tectonic domain	Locality	Latitude	Longitude	Temporal Period	SiO <sub>2</sub>	TiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	P <sub>2</sub> O <sub>5</sub>	LOI
Ets 209	1	Ethiopia	Gonder	11°24'58"	37°08'07"	3	44.51	1.78	15.61	4.18	6.40	0.20	11.34	10.08	3.33	1.29	0.30	0.97
Ets 213	1	Ethiopia	Gonder	11°01'30"	36°54'38"	3	46.93	1.32	16.88	2.95	7.20	0.22	7.73	9.38	4.65	1.70	0.35	0.67
Ets 124	3	Ethiopia	Gibat	8°45'00"	37°26'55"	3	42.21	2.72	14.67	2.16	9.10	0.18	11.87	10.72	3.08	0.97	0.83	1.49
Ets 126	3	Ethiopia	Wenchi	8°51'45"	37°43'30"	3	47.48	1.69	16.75	6.49	3.32	0.16	9.12	9.80	2.32	0.97	0.48	1.41
Ets 145	3	Ethiopia	Wechacha	9°04'00"	38°44'00"	3	46.27	1.73	15.60	2.25	7.90	0.16	10.94	9.45	2.88	0.74	0.45	1.62
Ets 161	3	Ethiopia	Wenchi	8°58'25"	37°45'40"	3	46.76	2.02	17.09	2.63	6.52	0.16	8.19	10.84	3.08	1.25	0.59	0.87
Ets 162	3	Ethiopia	Gibat	8°45'25"	37°27'00"	3	45.63	2.19	17.34	1.78	8.20	0.17	9.55	10.06	2.70	0.87	0.46	1.06
Ets 183	3	Ethiopia	Wechacha	9°03'50"	38°28'15"	3	48.59	1.77	17.90	1.66	7.34	0.16	6.80	7.78	4.41	1.56	0.92	1.11
Ets 21	3	Ethiopia	Debre Zeyt	8°50'45"	38°55'45"	3	49.53	2.34	17.73	2.54	8.10	0.17	4.81	8.64	3.35	1.06	0.74	1.00
Ets 243L	1	Ethiopia	Wollega	9°01'31"	36°09'59"	3	43.79	2.96	14.87	3.74	7.20	0.19	10.25	10.54	3.90	1.60	0.62	0.35
Ets 25	3	Ethiopia	Wechacha	8°53'25"	38°39'05"	3	50.00	2.60	18.41	2.57	7.36	0.16	4.36	7.65	3.95	1.27	0.68	1.00
Ets 39	3	Ethiopia	Wenchi	8°53'50"	37°59'20"	3	49.01	2.20	17.52	7.43	3.36	0.19	4.78	9.08	3.74	1.31	0.75	0.63
Ets 43	3	Ethiopia	Gibat	8°44'45"	37°28'05"	3	46.32	2.86	17.10	1.93	8.12	0.18	7.33	9.99	3.23	1.23	0.65	1.06
Ets 48	3	Ethiopia	Nekemt	8°49'25"	36°36'30"	3	44.93	3.08	15.49	2.30	8.96	0.17	7.69	10.61	2.28	1.38	0.63	2.47
Ets 62	3	Ethiopia	Debre Zeyt	8°39'00"	39°10'05"	3	48.62	1.75	16.81	3.84	6.46	0.19	7.13	9.74	2.89	0.84	0.60	1.14
Ets 81	3	Ethiopia	TulluWellel	8°56'30"	35°13'20"	3	49.09	2.61	16.10	3.38	9.20	0.20	5.41	8.18	3.24	0.73	0.47	1.38
Ets 84	3	Ethiopia	TulluWellel	8°56'30"	34°48'10"	3	51.58	2.73	16.20	4.93	5.00	0.22	4.38	6.55	4.49	1.84	1.48	0.60
Ets 241	3	Ethiopia	Arjo-Nekemt	08°51'53"	36°28'39"	2	50.82	2.56	16.31	1.32	10.24	0.19	4.71	7.66	3.57	1.11	0.51	1.01
Ets 281	3	Ethiopia	Guraghe	08°27'05"	38°15'19"	2	47.68	2.94	15.42	5.85	8.64	0.21	4.69	7.86	3.43	0.79	0.98	1.49
Ets 267	3	Ethiopia	Injbara (Tana)	11°24'58"	37°08'07"	2	48.76	1.33	17.95	0.09	8.80	0.15	9.52	8.73	3.10	0.86	0.24	0.48
AA3	2	Ethiopia	Debre Zeyt	8°43'25"	38°59'0"	2	47.03	1.90	17.03	3.86	6.20	0.17	8.89	9.93	2.92	1.18	0.51	0.38
Ets 110	3	Ethiopia	Konchi	8°53'35"	37°00'05"	2	47.13	3.03	16.74	3.97	7.78	0.19	5.27	7.86	4.25	1.63	0.66	1.50
Ets 245	3	Ethiopia	Arjo-Nekemt	08°45'52"	36°29'44"	2	47.02	2.99	17.93	2.18	8.08	0.21	4.39	8.49	4.85	1.79	0.89	1.17
Ets 93	3	Ethiopia	Nekemt	9°06'45"	36°38'30"	1	50.46	3.02	15.02	3.06	9.70	0.21	4.44	7.65	3.20	1.49	0.44	1.32
Ets 108	3	Ethiopia	Konchi	8°50'50"	37°03'05"	1	43.17	2.29	14.60	2.83	7.46	0.16	15.12	8.32	2.89	1.00	0.35	1.81
Etn 9	3	Turkana	Sidamo - Megga	4°03'32"	38°28'48"	3	43.41	2.49	14.13	3.33	8.72	0.19	11.84	9.29	3.87	1.58	0.52	0.62
Etn 13	3	Turkana	Sidamo - Megga	4°07'51"	38°34'48"	3	44.58	2.25	14.14	3.19	9.36	0.23	9.35	8.83	4.58	2.01	0.60	0.88

Footnotes : All isotopic data are unpublished. Notes, 1: major and trace elements from Conticelli et al. (1999), 2: major and trace elements from Gasparon et al. (1993), 3: all data unpublished. Temporal period, 1: 45-

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	Sc	V	Cr	Co	Ni	Zn	Rb	Sr	Y	Zr	Nb	Cs	Ba	Hf	Ta	Pb	Th	U	La	Ce	Nd	Sm	Eu	Gd	Tb	Dy	Ho	Er
6	n.a.	225	518	53	270	77	95	768	28	167	54	0.90	684	3.94	3.25	1.17	6.33	1.47	43	82	35	6.7	2.2	7.04	0.995	5.65	1.13	3.21
7	n.a.	164	259	19	130	75	105	901	29	211	136	1.17	994	4.58	7.86	4.68	13.6	3.04	77	131	48	7.8	2.5	8.13	1.02	5.52	1.11	3.19
8	n.a.	297	524	57	278	n.a.	27	858	29	141	23	0.32	487	2.67	1.6	0.541	4.22	1.1	45	97	48	9.4	3.0	9.29	1.25	6.71	1.25	3.36
9	n.a.	275	481	50	196	n.a.	22	815	25	156	34	0.098	824	3.78	1.71	2.1	2.81	0.677	26	54	27	5.6	2.0	5.86	0.849	4.87	0.961	2.64
10	n.a.	248	600	53	291	n.a.	21	626	25	184	42	0.27	335	4.14	2.17	3.81	2.58	0.682	26	55	28	5.8	1.9	6.04	0.868	4.82	0.922	2.46
11	n.a.	275	330	43	111	n.a.	34	740	30	174	41	0.32	568	4.12	2.04	0.882	4.03	0.934	39	69	34	6.8	2.3	7.18	1	5.53	1.09	2.98
12	n.a.	345	395	51	147	n.a.	23	735	26	127	44	0.25	369	3.19	2.35	0.696	3.19	0.825	32	64	30	6.0	2.0	6.3	0.885	5.04	0.995	2.73
13	n.a.	166	143	37	83	n.a.	50	1240	31	325	86	0.61	871	6.49	4.9	5.18	7.73	2.04	65	127	54	9.7	3.1	9.59	1.22	6.23	1.14	3.03
14	n.a.	336	72	41	25	n.a.	20	746	41	201	39	0.176	502	4.57	2.27	3.62	2.45	0.563	37	63	39	8.1	2.8	8.85	1.28	7.13	1.4	3.85
15	n.a.	254	281	49	167	83	44	1050	30	274	47	0.48	638	6.07	2.9	1.88	6.47	1.78	59	116	51	9.7	3.1	9.7	1.28	6.76	1.26	3.39
16	n.a.	270	20	39	17	n.a.	23	1170	45	194	51	0.25	698	4.68	3.28	1.36	3.22	0.863	52	80	54	10.5	3.3	11.2	1.59	9.01	1.82	5.14
17	n.a.	221	126	40	56	n.a.	25	684	32	249	47	0.26	716	6.01	2.95	1.46	3.19	0.813	42	89	47	9.4	3.3	9.46	1.33	7.31	1.41	3.81
18	28	314	248	44	82	n.a.	30	976	30	193	61	0.28	494	4.5	3.13	1.44	4.27	1.17	45	92	43	8.4	2.7	8.49	1.14	6.11	1.17	3.13
19	26	287	436	50	146	n.a.	49	747	30	216	29	0.36	354	4.95	1.79	0.547	3.81	1.07	37	84	45	9.7	3.1	9.51	1.33	7.11	1.29	3.37
20	n.a.	260	428	48	128	n.a.	14.7	565	67	155	24	0.071	501	3.92	1.35	0.469	1.84	0.429	57	51	48	9.3	2.9	11.4	1.57	8.82	1.9	5.06
21	n.a.	323	57	48	27	n.a.	14.3	603	28	161	22	0.53	511	4.37	1.23	1.41	2.11	0.52	25	57	32	7.2	2.5	7.35	1.08	6.13	1.2	3.24
22	n.a.	146	8	22	3	n.a.	33	1300	53	408	24	0.26	716	8.82	1.29	2.68	6.99	2.17	76	179	95	19.0	5.7	18.2	2.36	12	2.18	5.69
23	n.a.	300	52	39	19	n.a.	20	625	32	174	22	0.155	692	4.44	1.15	2.73	2.08	0.502	26	59	33	7.5	2.7	7.75	1.13	6.32	1.23	3.33
24	n.a.	335	23	49	43	n.a.	15	546	48	324	27	0.206	342	7.59	1.38	1.93	2.01	0.573	28	68	42	10.1	3.2	10.6	1.61	9.33	1.85	5.06
25	n.a.	204	296	45	163	n.a.	21	449	22	105	19	0.19	284	2.8	0.84	n.a.	1.73	0.454	15	31	17	3.9	1.41	4.45	0.71	4.31	0.887	2.43
26	n.a.	249	398	n.a.	119	n.a.	31	638	24	172	42	0.37	444	4.47	2.84	0.343	3.05	0.814	32	67	32	6.6	2.2	6.76	0.961	5.34	1.05	2.87
27	n.a.	259	17	39	12	n.a.	39	1030	31	255	63	0.43	658	6.71	4.05	3.5	5.22	1.27	50	104	50	9.9	3.1	9.87	1.34	7.06	1.33	3.59
28	n.a.	188	22	31	20	n.a.	52	1470	37	316	36	0.53	657	6.89	1.82	2.45	7.72	2.31	67	139	64	12.2	3.9	12	1.58	8.24	1.53	4.14
29	n.a.	350	22	45	13	n.a.	38	619	40	324	51	0.46	567	7.91	2.96	5.83	5.04	1.28	44	97	48	10.0	3.0	10	1.44	7.89	1.54	4.18
30	n.a.	337	110	48	39	n.a.	32	801	27	169	51	0.293	547	4.46	2.96	1.52	4.28	0.987	38	78	37	7.5	2.4	7.65	1.07	5.93	1.14	3.08
31	n.a.	232	476	57	349	92	49	875	26	170	41	0.52	621	3.71	2.69	1.17	5.59	1.33	48	96	44	8.5	2.8	8.58	1.15	6	1.1	2.86
32	n.a.	178	473	45	245	103	70	1080	34	317	90	0.79	761	6.84	5.58	2.77	7.83	1.97	60	117	51	9.7	3.2	9.85	1.31	6.93	1.29	3.43

25 Ma, 2: 25-10 Ma, 3: <10Ma. n.a. = not analyzed, m=measured, i=initial. 2 s.e. = 2\* standard error of the mean

Acta Vulcanologica, 11, 143-160.

of African Earth Sciences (and the Middle East), 17(2), 145-165.

Tm	Yb	Lu	$(^{87}\text{Sr}/^{86}\text{Sr})_m$	2 s.e.	$(^{87}\text{Sr}/^{86}\text{Sr})_i$	2 s.e.	$(^{143}\text{Nd}/^{144}\text{Nd})_m$	2 s.e.	$(^{143}\text{Nd}/^{144}\text{Nd})_i$	2 s.e.
0.444	2.8	0.431	0.703165	± 0.000007	0.703159	± 0.000007	0.512923	± 0.000005	0.512922	± 0.000005
0.448	2.96	0.468	0.703094	± 0.000006	0.703061	± 0.000006	0.512886	± 0.000006	0.512881	± 0.000006
0.426	2.65	0.396	0.703300	± 0.000006	0.703298	± 0.000006	0.512852	± 0.000005	0.512851	± 0.000005
0.359	2.23	0.34	0.704506	± 0.000006	0.704505	± 0.000006	0.512835	± 0.000005	0.512834	± 0.000005
0.328	2.02	0.307	0.703480	± 0.000008	0.703479	± 0.000008	0.512863	± 0.000005	0.512862	± 0.000005
0.39	2.35	0.366	0.703320	± 0.000007	0.703314	± 0.000007	0.512898	± 0.000005	0.512895	± 0.000005
0.364	2.26	0.345	0.703415	± 0.000008	0.703404	± 0.000008	0.512801	± 0.000004	0.512794	± 0.000004
0.389	2.42	0.364	0.703556	± 0.000007	0.703542	± 0.000007	0.512828	± 0.000006	0.512822	± 0.000006
0.506	3.19	0.491	0.704071	± 0.000007	0.704070	± 0.000007	0.512803	± 0.000005	0.512802	± 0.000005
0.44	2.67	0.405	0.703336	± 0.000006	0.703335	± 0.000006	0.512856	± 0.000006	0.512855	± 0.000006
0.727	4.72	0.758	0.703780	± 0.000008	0.703779	± 0.000008	0.512776	± 0.000006	0.512775	± 0.000006
0.502	3.09	0.472	0.703572	± 0.000007	0.703571	± 0.000007	0.512817	± 0.000004	0.512816	± 0.000004
0.404	2.48	0.371	0.703396	± 0.000006	0.703394	± 0.000006	0.512780	± 0.000006	0.512779	± 0.000006
0.429	2.53	0.367	0.703239	± 0.000007	0.703185	± 0.000008	0.512839	± 0.000006	0.512822	± 0.000006
0.61	3.33	0.525	0.704717	± 0.000008	0.704695	± 0.000008	0.512730	± 0.000005	0.512714	± 0.000005
0.439	2.71	0.416	0.703782	± 0.000006	0.703773	± 0.000006	0.512737	± 0.000004	0.512728	± 0.000004
0.704	4.24	0.628	0.703548	± 0.000008	0.703535	± 0.000008	0.512773	± 0.000005	0.512762	± 0.000005
0.441	2.81	0.427	0.703970	± 0.000007	0.703961	± 0.000007	0.512712	± 0.000005	0.512706	± 0.000005
0.689	4.3	0.661	0.703798	± 0.000008	0.703793	± 0.000008	0.512764	± 0.000007	0.512759	± 0.000007
0.342	2.13	0.327	0.703388	± 0.000007	0.703373	± 0.000007	0.512889	± 0.000005	0.512881	± 0.000005
0.389	2.44	0.369	0.703867	± 0.000006	0.703865	± 0.000006	0.512820	± 0.000004	0.512819	± 0.000004
0.465	2.86	0.429	0.703507	± 0.000007	0.703495	± 0.000007	0.512779	± 0.000004	0.512773	± 0.000004
0.531	3.35	0.5	0.703129	± 0.000007	0.703128	± 0.000007	0.512877	± 0.000005	0.512877	± 0.000005
0.554	3.45	0.521	0.703949	± 0.000006	0.703929	± 0.000006	0.512722	± 0.000004	0.512716	± 0.000004
0.402	2.5	0.373	0.703522	± 0.000006	0.703514	± 0.000006	0.512764	± 0.000006	0.512760	± 0.000006
0.361	2.19	0.322	0.703060	± 0.000006	0.703057	± 0.000006	0.512918	± 0.000004	0.512917	± 0.000004
0.447	2.78	0.418	0.703039	± 0.000006	0.703036	± 0.000006	0.512924	± 0.000005	0.512923	± 0.000005